

# Analysis of Selected Models for Inelastic Electron-Scattering in the KATRIN Gaseous Tritium Source

Master's thesis of

**Fabian Leven**

at the Institute of Experimental Particle Physics (ETP)  
of the Department of Physics  
at the Karlsruhe Institute of Technology

Reviewer: Prof. Dr. G. Drexlin  
Institute of Experimental Particle Physics, KIT

Second Reviewer: Dr. K. Valerius  
Institute for Nuclear Physics, KIT

Advisor: M. Sc. M. Machatschek  
Institute of Experimental Particle Physics, KIT

July 15th, 2018 – July 14th, 2019



# Erklärung der Selbstständigkeit

Ich versichere, dass ich diese Arbeit selbstständig verfasst habe und keine anderen als die angegebenen Quellen und Hilfsmittel benutzt habe, die wörtlich oder inhaltlich übernommenen Stellen als solche kenntlich gemacht und die Satzung des KIT zur Sicherung guter wissenschaftlicher Praxis in der gültigen Fassung vom 24.05.2018 beachtet habe.

Karlsruhe, den \_\_\_\_\_ , \_\_\_\_\_  
Fabian Leven

Als Prüfungsexemplar genehmigt von

Karlsruhe, den \_\_\_\_\_ , \_\_\_\_\_  
Prof. Dr. G. Drexlin



# Introduction

To our current knowledge, neutrinos are at the same time the most elusive and most abundant massive particles in the Universe. Their detection is challenging and pushes experiments to the edge of technical frontiers. Yet, their understanding might shed light on long-standing open questions of modern physics: What is dark matter? Do Majorana particles exist? Why does matter predominate over antimatter? What happened in the earliest stages of our Universe just after the Big Bang? And in all these regards, the yet unknown mass of neutrinos is a key physics parameter.

The Karlsruhe TRItium Neutrino (KATRIN) experiment aims to measure the effective mass of the electron antineutrino with an unprecedented sensitivity of 200 meV (90 % C.L.) by precision spectroscopy of the endpoint region of the tritium- $\beta$ -decay spectrum. This sensitivity is one order of magnitude more constraining than the one of the predecessor experiments.

In order to provide this sensitivity, KATRIN features i. a. a gaseous tritium source. Its special characteristics must be well controlled and understood. This thesis focuses on two selected models concerning electron-scattering in said gaseous source:

- ① The dependence of the cross section for inelastic electron-scattering off hydrogen isotopologues on the energy of the incident electrons should be investigated. Former works primarily used a constant cross section when simulating KATRIN neutrino mass measurements. An energy-dependent scattering model should be established and applied within simulated neutrino mass inference.
- ② In a commissioning measurement that was conducted in fall 2018, the KATRIN collaboration has established a preliminary model of the probability density for the energy an electron loses when scattering in the KATRIN gaseous tritium source. The aim of this thesis was the implementation of statistical methods that enable the general treatment of model uncertainties. The methods were then tested in the context of the preliminary energy loss model in order to assess systematic effects in neutrino mass inference.

The results are listed in the summary in chapter 7.

## Outline

This thesis is structured as follows:

Chapter 1 is a brief introduction to neutrino physics with special emphasis on the measurement of the neutrino mass.

Chapter 2 focuses on the experimental setup of the KATRIN experiment.

Chapter 3 introduces a mathematical model of a KATRIN neutrino mass measurement that can be used in parameter inference.

Chapter 4 integrates the mathematical model into a statistical framework for neutrino mass inference.

Chapter 5 investigates a refinement of the mathematical model by incorporating the energy-dependence of the cross section for inelastic electron scattering in the KATRIN gaseous tritium source as per item ①.

Chapter 6 focuses on the treatment of uncertainties from a preliminary model for the energy loss of electrons scattering inelastically off deuterium molecules as per item ②.

And chapter 7 summarizes the results.

# Contents

|  |             |
|--|-------------|
| <b>Introduction</b>  | <b>v</b>    |
| <b>List of Figures</b>   | <b>xi</b>   |
| <b>List of Tables</b>  | <b>xiii</b> |
| <b>1. Neutrino Physics</b>   | <b>1</b>    |
| 1.1. Neutrinos until the 1960s . . . . .                             | 1           |
| 1.2. Neutrinos in the Standard Model of Particle Physics . . . . .   | 2           |
| 1.2.1. General Particle Properties . . . . .                         | 3           |
| 1.2.2. Neutrino Properties . . . . .                                 | 3           |
| 1.3. Mechanisms that Generate Neutrino Masses . . . . .              | 3           |
| 1.3.1. Neutrino Mass Terms . . . . .                                 | 4           |
| 1.3.2. Neutrino Mixing . . . . .                                     | 5           |
| 1.4. Neutrino Oscillations . . . . .                                 | 5           |
| 1.4.1. Relation to Neutrino Masses . . . . .                         | 6           |
| 1.4.2. Experimental Considerations . . . . .                         | 7           |
| 1.4.3. Experimental Case Study: The Solar Neutrino Problem . . . . . | 7           |
| 1.4.4. Summary of Experimental Results . . . . .                     | 8           |
| 1.5. Absolute Neutrino Mass Measurements . . . . .                   | 8           |
| 1.5.1. Observational Cosmology . . . . .                             | 8           |
| 1.5.2. Search for Neutrinoless Double- $\beta$ Decay . . . . .       | 10          |
| 1.5.3. Kinematic Measurements of Weak Decays . . . . .               | 10          |
| <b>2. The KATRIN Experiment</b>                                      | <b>13</b>   |
| 2.1. Overview of the KATRIN Experimental Setup . . . . .             | 13          |
| 2.2. Windowless Gaseous Tritium Source . . . . .                     | 14          |
| 2.3. Rear Section . . . . .  | 15          |
| 2.4. Differential Pumping Section . . . . .                          | 16          |
| 2.5. Cryogenic Pumping Section . . . . .                             | 16          |
| 2.6. Pre and Main Spectrometer . . . . .                             | 18          |
| 2.6.1. MAC-E-Filter Principle . . . . .                              | 18          |
| 2.6.2. Background Mitigation Strategies . . . . .                    | 21          |
| 2.7. Detector Section . . . . .                                      | 21          |
| <b>3. Mathematical Formalism of a KATRIN Measurement</b>             | <b>23</b>   |
| 3.1. Differential Tritium- $\beta$ -Decay Spectrum . . . . .         | 23          |
| 3.2. Response Function . . . . .                                     | 25          |
| 3.2.1. Concepts and Nomenclature . . . . .                           | 25          |
| 3.2.2. Gas Dynamics . . . . .  | 26          |
| 3.2.3. Transmission Function . . . . .                               | 27          |
| 3.2.4. Probability for Electron-Scattering within the WGTS . . . . . | 28          |
| 3.2.5. Energy Loss of Electrons due to Scattering . . . . .          | 29          |

|           |  |           |
|-----------|--|-----------|
| 3.2.6.    | Assembly of the Response Function . . . . .  | 30        |
| 3.2.7.    | Discussion of the Response Function . . . . .  | 31        |
| 3.3.      | Integral Rate . . . . .  | 32        |
| 3.4.      | Detector Counts . . . . .  | 33        |
| 3.5.      | Model Amendments . . . . .   | 33        |
| 3.6.      | Measurement Time Distribution . . . . .  | 33        |
| 3.7.      | A Simulated KATRIN Neutrino Mass Measurement . . . . .   | 34        |
| <b>4.</b> | <b>Statistical Methods and Neutrino Mass Inference at KATRIN</b>   | <b>37</b> |
| 4.1.      | Maximum Likelihood Estimation . . . . .  | 37        |
| 4.2.      | The Likelihood of a KATRIN Measurement . . . . .   | 37        |
| 4.3.      | A Nominal KATRIN Neutrino-Mass Fit . . . . .   | 38        |
| 4.4.      | Uncertainty Intervals . . . . .  | 38        |
| 4.4.1.    | Bayesian Credible Intervals . . . . .  | 39        |
| 4.4.2.    | Frequentist Confidence Intervals . . . . .   | 39        |
| 4.5.      | The KaFit and SSC Software Frameworks . . . . .  | 41        |
| 4.6.      | KATRIN's Sensitivity to the Neutrino Mass . . . . .  | 41        |
| 4.6.1.    | Definition and Construction of KATRIN's Sensitivity . . . . .  | 42        |
| 4.6.2.    | Sensitivity from Ensemble Tests . . . . .  | 42        |
| 4.6.3.    | Sensitivity using the Profile-Likelihood Method . . . . .  | 43        |
| <b>5.</b> | <b>Energy-Dependence of the Cross Section for Inelastic Electron-Scattering within the WGTS</b>  | <b>45</b> |
| 5.1.      | Cross Section Values for Inelastic Electron-Scattering off Molecular Hydrogen Isotopologues . . . . .  | 45        |
| 5.1.1.    | Theoretical Cross Section Formulae . . . . .   | 46        |
| 5.1.2.    | Relation to Cross Section Values Used in Former Works . . . . .  | 47        |
| 5.2.      | An Energy-Dependent Scattering Model using the Poisson Distribution . . . . .  | 48        |
| 5.2.1.    | Formalism of the Poisson Model for Electron-Scattering . . . . .   | 49        |
| 5.2.2.    | Properties of the Poisson Model for Electron-Scattering . . . . .  | 50        |
| 5.3.      | Effects of Systematic Cross-Section Offsets in Neutrino Mass Inference . . . . .   | 53        |
| 5.3.1.    | Neutrino Mass Shifts for an Energy-Independent Cross-Section Offset from Former Works . . . . .  | 53        |
| 5.3.2.    | Neutrino Mass Shifts for an Energy-Dependent Cross-Section Offset . . . . .  | 54        |
| 5.3.3.    | Relation of the Shifts for an Energy-Dependent and -Independent Cross-Section Offset . . . . .   | 55        |
| 5.4.      | An Energy-Dependent Scattering Model beyond Using the Poisson Distribution . . . . .   | 57        |
| 5.4.1.    | Formalism of the Extended Model for Electron-Scattering . . . . .  | 58        |
| 5.4.2.    | Numerical Accuracy and Cross-Check of the Extended Model for Electron-Scattering . . . . .   | 59        |
| 5.4.3.    | Discussion of the Extended Model for Electron-Scattering . . . . .   | 62        |
| 5.5.      | Conclusion and Outlook . . . . .   | 63        |
| <b>6.</b> | <b>Statistical Methods in the Context of the Empirical Energy Loss Model Derived from KATRIN Data for Electrons Scattering Inelastically off Deuterium</b> | <b>65</b> |
| 6.1.      | The Empirical KATRIN Energy Loss Model . . . . .   | 66        |
| 6.1.1.    | Description and Parametrization of the KATRIN Energy Loss Model . . . . .  | 66        |
| 6.1.2.    | Nuisance Parameters that Further Influence the KATRIN Energy Loss Model . . . . .  | 67        |
| 6.2.      | Scope of the Conducted Study . . . . .   | 68        |
| 6.3.      | Statistical Prerequisites . . . . .  | 70        |
| 6.3.1.    | Combination of a Calibration and a Neutrino Mass Measurement . . . . .   | 71        |



|                     |  |           |
|---------------------|--|-----------|
| 6.3.2.              | Nuisance Parameters and the Profile-Likelihood Method . . . . .                                  | 72        |
| 6.3.3.              | An Asimov Data Set in Relation to Ensemble Tests in Sensitivity<br>Studies . . . . .             | 73        |
| 6.4.                | Free Fit Parameters of the Energy Loss Model in Neutrino Mass Inference .                        | 74        |
| 6.4.1.              | Sensitivity from an Asimov Data Set . . . . .  | 74        |
| 6.4.2.              | Cross-Check and Extension of the Asimov Data Set via Ensemble<br>Testing . . . . .               | 76        |
| 6.4.3.              | Reduction of Systematic Effects by Additional Fit Parameters . . . .                             | 78        |
| 6.5.                | Conclusion . . . . .   | 81        |
| <b>7.</b>           | <b>Summary</b>   | <b>83</b> |
| <b>Appendix</b>     |  | <b>85</b> |
| A.                  | Proof that the Poisson Model for Electron-Scattering is a Probability Density                    | 85        |
| B.                  | Measurement Time Distributions . . . . .   | 86        |
| C.                  | Documentation of KaFit-Likelihood Extensions . . . . .   | 87        |
| D.                  | Preliminary Parameter Values of the KATRIN Energy Loss Model . . . . .                           | 88        |
| E.                  | Configuration of the Sensitivity Study using the Empirical KATRIN Energy<br>Loss Model . . . . . | 89        |
| <b>Acronyms</b>     |  | <b>91</b> |
| <b>Bibliography</b> |  | <b>93</b> |
|                     | References . . . . .   | 93        |
|                     | Preliminary . . . . .  | 100       |
|                     | Software . . . . .   | 100       |



# List of Figures

|  |    |
|--|----|
| 1.1. The Standard Model of Particle Physics . . . . .  | 4  |
| 1.2. Neutrino mixing parameters and mass ordering . . . . .  | 9  |
| 1.3. Feynman graph of neutrinoless double-beta decay . . . . .   | 11 |
| 2.1. KATRIN beamline . . . . .   | 14 |
| 2.2. KATRIN windowless gaseous tritium source (WGTS) . . . . .   | 15 |
| 2.3. KATRIN rear section (RS) . . . . .  | 16 |
| 2.4. KATRIN differential pumping section (DPS) . . . . .   | 17 |
| 2.5. KATRIN cryogenic pumping section (CPS) . . . . .  | 18 |
| 2.6. Scheme of the KATRIN main spectrometer and the MAC-E filter principle.  | 19 |
| 2.7. KATRIN detector section . . . . .   | 22 |
| 3.1. Tritium- $\beta$ spectrum for a vanishing and non-vanishing neutrino mass . . . .   | 24 |
| 3.2. The KATRIN transmission function . . . . .  | 27 |
| 3.3. The probability density for the energy loss of electrons due to scattering off<br>tritium molecules . . . . .                               | 29 |
| 3.4. The KATRIN response function . . . . .  | 32 |
| 3.5. Simulated KATRIN neutrino mass measurement for a non-vanishing neutrino<br>mass . . . . .   | 35 |
| 4.1. Illustration of a confidence interval obtained via the likelihood ratio . . . .   | 40 |
| 4.2. Illustration of KATRIN's sensitivity to the neutrino mass . . . . .   | 42 |
| 5.1. Inelastic cross section for electrons scattering off molecular hydrogen iso-<br>topologues . . . . .  | 48 |
| 5.2. Probability for severalfold electron-scattering in the WGTS . . . . .   | 51 |
| 5.3. Probability for one-fold electron-scattering in the WGTS in an extended<br>energy range . . . . .   | 52 |
| 5.4. Shift of an inferred squared neutrino mass induced by a neglected energy-<br>dependence of the inelastic scattering cross section . . . . . | 54 |
| 5.5. Shift of an inferred squared neutrino mass induced by a neglected constant<br>offset of the inelastic scattering cross section . . . . .    | 56 |

|  |    |
|--|----|
| 5.6. Simulated integral $\beta$ -electron rate using different models of the cross section for inelastic electron-scattering . . . . .                               | 57 |
| 5.7. Accuracy of the numerical evaluation of the extended model of the probability for electron-scattering in the WGTS . . . . .                                     | 61 |
| 5.8. Comparison of the Poisson and the extended model for one-fold scattering of electrons in the WGTS . . . . .   | 63 |
| 6.1. The preliminary empirical KATRIN energy loss model for electrons scattering off deuterium molecules . . . . .   | 69 |
| 6.2. Profile-likelihood ratio of a KATRIN measurement from an Asimov data set  | 75 |
| 6.3. Test of coverage for the ensemble of confidence intervals obtained in the sensitivity study using the KARTRIN energy loss model . . . . .                       | 77 |
| 6.4. Chi-square distribution for the simulated ensemble of neutrino mass measurements obtained in the sensitivity study using the KATRIN energy loss model . . . . . | 78 |

# List of Tables

|   |    |
|---|----|
| 1.1. Neutrino source classes for neutrino oscillation experiments . . . . .   | 7  |
| 1.2. Overview of neutrino oscillation parameters . . . . .  | 10 |
| 1.3. Constrains on the neutrino mass by double- $\beta$ decay experiments . . . . .   | 10 |
| 1.4. Constraints on the neutrino mass by kinematic measurements . . . . .   | 11 |
| 3.1. Probability for severalfold electron-scattering in the WGTS . . . . .  | 29 |
| 4.1. KATRIN's sensitivity to the neutrino mass from ensemble tests . . . . .  | 43 |
| 5.1. Probability for severalfold electron-scattering in the WGTS - reviewed . . .   | 49 |
| 6.1. Neutrino mass confidence intervals and sensitivities obtained from an Asimov data set . . . . .                        | 76 |
| 6.2. Systematic shifts of the squared neutrino mass induced by offsets of the parameters of the energy loss model . . . . . | 80 |



# 1. Neutrino Physics

This chapter is an introduction to neutrino physics. The primary aim is to give an experimentally-rooted definition of a neutrino. Therefore, in section 1.1, selected experimental milestones are outlined that led to today's description of a neutrino within the established Standard Model of Particle Physics (SM). In section 1.2, it follows an outline of the SM and how it relates to the neutrino. Special attention is paid to the neutrino mass: First, in section 1.3, an extension of the SM that allows for neutrino masses is summarized. Second, in section 1.4, the phenomenon of neutrino flavor oscillations is introduced. Corresponding experiments proved that neutrinos have mass. Third, in section 1.5, experiments for an absolute neutrino mass measurement are presented because as such they relate particularly to the KATRIN experiment.

## 1.1. Neutrinos until the 1960s

Albeit the neutrino as a hypothetical new particle was not postulated until 1930, its rich scientific history might be seen as already heralded during the preceding 35 years.

In 1895, Becquerel reported results on experiments with phosphorescent substances, especially uranium salts, on photographic plates [Bec95]. These experiments mark the discovery of radioactivity and triggered manifold subsequent investigations.

In 1899, Rutherford published a classification of radioactive decays into  $\alpha$  and  $\beta$  types according to their penetration strength [Rut99].

In 1900, Becquerel determined the mass-charge ratio of  $\beta$ -decay particles [Bec00] and identified them as the electron previously described by Thomson [Tho97].

In 1914, Chadwick measured a continuous electron energy spectrum in the  $\beta$  decay of lead-214 and bismuth-214 [Cha14].

In 1927, Ellis and Wooster conducted a calorimetric measurement of the  $\beta$ -decay energy of radium and demonstrated that the continuity of the  $\beta$  spectrum was intrinsic to the decay as opposed to be caused by secondary effects [EW27].

In 1930, a  $\beta$  decay was thought of as a two-body decay  ${}^z\text{A} \rightarrow {}^{z+1}\text{B} + e^-$ . Assuming conservation of energy and momentum, in a two-body-decay, the momenta of the daughter particle B and the electron e are solely determined by their masses and the “energy content”, as Bohr put it, of the parent particle A [Boh32]. According to Bohr, there was no reason to believe that different nuclei of the same element A should have a different energy content in a  $\beta$  decay. Hence, the continuous nature of the  $\beta$  spectrum could not be explained. As a possible solution, Pauli suggested the  $\beta$  decay to be a three-body decay and postulated an electrically neutral particle that carries part of the decay energy [Pau30].

In 1934, Fermi developed a quantitative theory of  $\beta$  decay that could describe the preceding experimental results [Fer34]. It comprises a four-fermion contact interaction, respectively a

three-body-decay model. It was the first description of the so-called “weak interaction”. Furthermore, Fermi coined the term “neutrino” for the particle postulated by Pauli. Fermi’s theory inspired the idea to use the “inverse  $\beta$  decay” or “neutrino capture” to detect neutrinos, which in today’s nomenclature is written as

$$\bar{\nu}_e + p^+ \rightarrow n + e^+.$$

In 1956, Cowan and Reines published results of a corresponding experiment [Cow+56]. It was conducted using the high neutrino flux of the nuclear reactor of the Savannah River Plant. The neutrinos originating in the reactor passed a tank of water and cadmium chloride triggering the above process. The emerging neutron was captured by the cadmium which emitted a photon in a 3 MeV to 11 MeV range

$$n + {}^{113}\text{Cd} \rightarrow {}^{114}\text{Cd} + \gamma.$$

The emerging positron annihilated with an electron which produced two photons of 0.5 MeV each. A coincidence measurement of the corresponding photons enabled discriminating signal and background events. Based on their results, Cowan and Reines reported the discovery of the free neutrino.

In the same year, 1956, Lee and Yang published an article on parity conservation [LY56]. Parity conservation implies that a mirrored physical process behaves the same as its non-mirrored counterpart. Here, mirroring means a change of sign of the position vector in the applied physical laws. Lee and Yang pointed out that parity conservation might be violated in weak interactions and suggested several probing methods.

In 1957, Wu et al. conducted an experiment that employed one of the corresponding probing methods, based on  $\beta$  decay [Wu+57]. The parity operation, respectively “the mirroring”, corresponded a change of the magnetic field orientation in the experiment. The results showed that parity is violated in weak interactions.

In 1958, Goldhaber et al. measured the helicity  $H$  of the neutrino [Gol+58]. Helicity is defined as  $H = \hat{\sigma} \cdot \hat{\mathbf{p}}$ , where  $\hat{\sigma}$  is the spin unit vector and  $\hat{\mathbf{p}}$  is the momentum unit vector (here: of the neutrino). The experiment found  $H = -1$  which corresponds to maximum parity violation. In other words, only left-handed neutrinos and right-handed antineutrinos participate in weak interactions.

In 1962, Danby et al. reported on a second type of neutrinos [Dan+62]. A beam of pions generated at the Alternating Gradient Synchrotron in Brookhaven decayed according to  $\pi^\pm \rightarrow \mu^\pm + \bar{\nu}$ . The emerging neutrinos penetrated a 13.5-meter iron shield wall and their interactions were detected in a 10-t aluminum spark chamber. The observed interactions were path-like as opposed to shower-like, which implied the production of muons as opposed to electrons. This was marked as the discovery of the muon neutrino.

The attempts to uniformly describe the manifold discoveries in the field of particle physics in a combined theory converged over the course of the second half of the 20th century into what is known today as the Standard Model of Particle Physics.

## 1.2. Neutrinos in the Standard Model of Particle Physics

This section introduces the basic concepts of the Standard Model of Particle Physics (SM) in a condensed manner. It describes general particle properties in section 1.2.1 and relates them to neutrinos in section 1.2.2.

The SM is a gauge quantum field theory exhibiting the gauge symmetry  $SU(3) \times SU(2) \times U(1)$ . As such, it can be formulated using the principle of least action and a Lagrangian density



$\mathcal{L}$  depending on fields and their derivatives [Zee03]. Albeit it can not account for all known subatomic phenomena, within its known boundaries, the SM is a well-tested and established theory, which is evident by e.g. the extensive review of particle properties of the Particle Data Group [Tan+18].

### 1.2.1. General Particle Properties

The gap between fields and particles can be bridged as follows: If “[i]n region 1 in spacetime there exists a source that sends out a ‘disturbance in the field’, which is later absorbed by a sink in region 2 in spacetime[,] experimentalists choose to call this a particle” [Zee03]. Intrinsic particle properties can be derived from the relation of their associated fields to the Lagrangian density. For example, a particle’s mass is encoded by the Yukawa coupling of its field to the higgs doublet through the higgs mechanism and spontaneous symmetry breaking [Hig64]. A further intrinsic property is a particle’s spin, that takes half-integer values in units of  $\hbar$  for fermions or integer values for bosons. A particle’s flavor is its eigenstate with respect to the weak interaction, which is described by the  $SU(2) \times U(1)$  subgroup (Glashow-Weinberg-Salam model [Gla61; Wei67; Sal68]). According to Noether’s theorem, each symmetry conserves an associated charge [Noe18]. In the case of the  $SU(2) \times U(1)$  symmetry, the associated charges are called isospin  $\mathbf{T} = (T_1, T_2, T_3)^T$  and hypercharge  $Y$ . A derivative of these charges is the electric charge  $Q = T_3 + \frac{1}{2}Y$  [Sch19]. In that sense, each particle has an associated antiparticle that carries the opposite electric charge. As mentioned in the historical overview (section 1.1), a theory consistent with experiment must violate parity. Such theories are called chiral [Zee03]. The SM is a chiral theory and thus its fermion fields can be decomposed in left- and right handed components [Sch19].

Figure 1.1 depicts the particles of the SM along with their selected properties mass and electric charge. It also shows a further categorization among the fermions into quarks and leptons.

### 1.2.2. Neutrino Properties

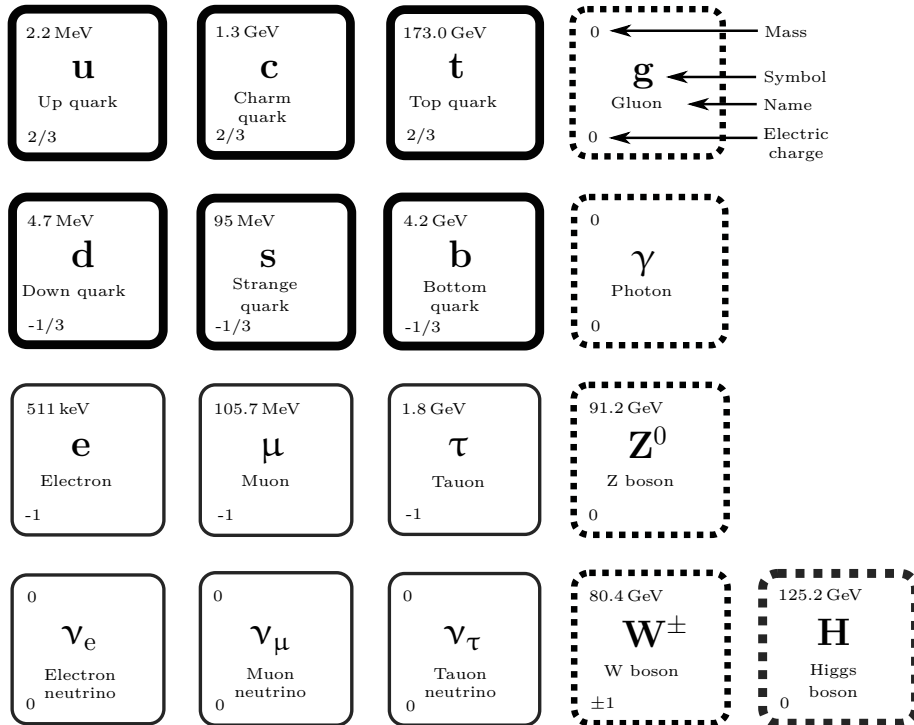
With reference to the particle properties listed in the previous section 1.2.1, a neutrino can be described as follows: A neutrino carries a spin of  $1/2\hbar$ . Thus, it is a fermion. It is categorized as a lepton. It has an electric charge of 0. And there are only left-handed neutrinos and right-handed antineutrinos [Sch19].

The mass of a neutrino will be discussed separately within the following chapters.

Neutrinos come in three flavors, typically denoted as  $\nu_e$ ,  $\nu_\mu$  and  $\nu_\tau$ . Some additional remarks about the neutrino flavors can be made: First, the historical overview (section 1.1) mentions the discovery of the electron and muon flavor, but it was not until 2001 that the tau neutrino was discovered by the DONUT collaboration [Kod+01]. Second, a precision measurement of the width of the  $Z^0$ -boson resonance  $\Gamma_Z = (2.68 \pm 0.15) \text{ GeV}$  at the Large Electron Positron Collider (LEP) in the 1990s yielded a number of active light neutrino flavors consistent with three. In this context, “light” refers to a neutrino mass smaller than half the mass  $M_Z = (91.174 \pm 0.070) \text{ GeV}$  of the  $Z^0$ -boson [Acc+98]. (It should be noted that this refers to active neutrino flavors. Sterile neutrinos as an extension of the SM are not ruled out [OW08].)

## 1.3. Mechanisms that Generate Neutrino Masses

The subsequent section 1.4 lists experiments which proof that neutrinos have mass. However, in the SM, as described in section 1.2, the neutrino masses are assumed to be zero. Nevertheless, an extension of the SM is possible. The corresponding mass terms are introduced in section 1.3.1. Moreover, the neutrino-mass formalism entails neutrino flavor mixing as described in section 1.3.2.



**Figure 1.1.: The Standard Model of Particle Physics.** The diagram illustrates possible categorizations of particles within the SM. The fermions are framed with continuous and the bosons with dotted lines. Among the fermions the quark sector is marked by a thick frame and the lepton sector by a thin one. The first three columns show the fermions; and the fourth and the fifth the bosons. While the bosons in the fourth column carry a spin of 1, the higgs boson, that is depicted in the fifth column and marked with a thicker frame, carries a spin of 0. Also, the particle mass in natural units is shown and the electric charge in units of the absolute electron charge. All quantities along with uncertainties can be found in the Review of Particle Physics [Tan+18]. (Illustration adapted from [Sei19].)

### 1.3.1. Neutrino Mass Terms

For a theory to account for neutrino masses, its Lagrangian density must exhibit corresponding mass terms. According to [Zub11], the formalism can be summarized: The form of a mass term is given by the Dirac equation, which is produced by applying the principle of least action to a suitable Lagrangian density  $\mathcal{L}$ . The mass terms have to be quadratic in the fermion fields  $\psi$  and must leave the Lagrangian density hermitian. Furthermore, a field  $\psi$  must have a left- and right-handed component in order for the mass terms not to vanish. Two possible term forms are named after Dirac and Majorana. Whether one or a mixture of both forms corresponds to the neutrino's reality is an open question.

#### Dirac Masses

A Dirac mass term with mass  $m_D$  split in its chiral components (Weyl spinors)  $\psi_{L,R}$  has the form [Zub11]

$$\mathcal{L}_D = -m_D \bar{\psi} \psi = -m_D (\bar{\psi}_L \psi_R + \bar{\psi}_R \psi_L). \quad (1.1)$$

Applying this to neutrinos requires both a left- and a right-handed Dirac neutrino. Right-handed neutrinos have not yet been observed. If they exist, they do not interact weakly and hence are called sterile.

## Majorana Masses

For Majorana mass terms the CP-conjugate  $\psi^C$  of a fermion spinor  $\psi$  is used. It should be noted, that if  $\psi$  is left-handed,  $\psi^C$  is right-handed and vice versa. Then, a Majorana field  $\phi$  can be defined and a corresponding mass term  $\mathcal{L}_M$  with a mass  $m_M$  can be constructed [Zub11]:

$$\phi = \psi + \psi^C \quad \mathcal{L}_M = -\frac{1}{2}m_M\bar{\phi}\phi. \quad (1.2)$$

As  $\phi^C = \phi$ , the described Majorana particle is its own antiparticle, which, due to charge conservation, is only possible for neutral particles, such as a neutrino.

### 1.3.2. Neutrino Mixing

If neutrinos have mass, their mass eigenstates  $|\nu_i\rangle$  ( $i \in \{1, 2, 3\}$ ) of the free Hamiltonian need not be identical to their flavor eigenstates  $|\nu_\alpha\rangle$  ( $\alpha \in \{e, \mu, \tau\}$ ) of the weak interaction [Zub11]. In case they differ, there must be a basis-change matrix. Such a matrix was introduced by Maki, Nakagawa and Sakata in order to explain the so-called neutrino oscillations (see section 1.4) predicted by Pontecorvo [Pon58; MNS62]. Therefore, the matrix  $U$  for a basis change is called Pontecorvo-Maki-Nakagawa-Sakata matrix (PMNS matrix)

$$|\nu_\alpha\rangle = \sum_i U_{\alpha i} |\nu_i\rangle. \quad (1.3)$$

As a complex unitary matrix,  $U$  can be expressed by six parameters. A possible choice are three angles  $\theta_{12}, \theta_{23}, \theta_{13} \in [0, 2\pi)$ , a phase  $\delta \in [0, 2\pi)$  and two Majorana phases  $\alpha, \beta \in [0, 2\pi)$ :

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}. \quad (1.4)$$

These parameters are called neutrino mixing parameters. It should be noted, that  $\delta$  is also called “ $CP$ -violating” phase. Here,  $P$  stands for parity conjugation as it was explained in section 1.1; and  $C$  for electric charge conjugation that follows the same idea with a sign change of the electric charge instead of the position vector. Why  $\delta \neq 0$  implies  $CP$ -violation is shown in section 1.4.1.

One of the consequences of neutrino mixing, namely neutrino oscillations, is explained in the following section 1.4.

## 1.4. Neutrino Oscillations

The term “neutrino oscillations” refers to the neutrino’s change of flavor after passing a certain propagation distance. In other words, neutrinos might be detected in another flavor than the one they originated in. Section 1.4.1 introduces a demonstrative formalism that aims at showing the link between oscillations and the masses of neutrinos. Neutrino oscillations also depend on the neutrino mixing parameters introduced in section 1.3.2. The accessibility of these mixing parameters and neutrino masses via neutrino oscillation experiments will be evaluated in section 1.4.4. Neutrino oscillation experiments are manifold. As an exemplary case study the so-called “solar neutrino problem” is discussed in section 1.4.3. Finally, the experimental results on neutrino oscillations will be summarized in section 1.4.4.

### 1.4.1. Relation to Neutrino Masses

According to [Zub11], a formula demonstrating neutrino oscillations can be derived: Using the PMNS matrix  $U$  from equation (1.4), the evolution of a neutrino's flavor eigenstate on a one-dimensional path starting at position  $x = 0$  at time  $t = 0$  with momentum  $p_i$  and energy  $E_i$  of its mass eigenstates  $|\nu_i\rangle$  is

$$|\nu_\alpha(x, t)\rangle = \sum_i U_{\alpha i} e^{-i(E_i t - p_i x)} |\nu_i\rangle. \quad (1.5)$$

This leads to the transition amplitudes

$$A(\alpha \rightarrow \beta)(t) = \langle \nu_\beta | \nu_\alpha(x) \rangle = \sum_i U_{\beta i}^* U_{\alpha i} e^{-i(E_i t - p_i x)t} \quad (1.6a)$$

$$A(\bar{\alpha} \rightarrow \bar{\beta})(t) = \langle \bar{\nu}_\beta | \bar{\nu}_\alpha(x) \rangle = \sum_i U_{\beta i} U_{\alpha i}^* e^{-i(E_i t - p_i x)t}. \quad (1.6b)$$

It should be noted that if  $U \neq U^*$  (eq. 1.6) implies  $CP$ -violation. In reference to section 1.3.1, it holds  $U \neq U^* \Leftrightarrow \delta \neq 0$ , justifying that  $\delta$  is called  $CP$ -violating phase.

The following assumptions allow for a simple and demonstrative form of the transition probability:

- The neutrinos are relativistic:
  - Their momentum equals approximately their energy which is by far larger than their mass  $p_i \approx E_i \gg m_i$ . This also implies that the energy can be expanded in the mass-momentum-ratio  $m_i/p_i$ .
  - They travel the distance  $x = L = ct$  at the speed of light  $c$ .
- All neutrino generations have approximately the same momentum  $E \approx p \approx p_i$ .
- The  $CP$ -violating phase vanishes:  $\delta = 0$ . (This assumption is not necessary, but simplifies the expression for the transition probability. See [Zub11] for  $\delta \neq 0$ .)

Then, the transition probability from one flavor  $\alpha$  to another  $\beta$  in dependence of the neutrino masses and mixing parameters is

$$\begin{aligned} P(\alpha \rightarrow \beta)(L) &= |\langle \nu_\beta | \nu_\alpha(L) \rangle|^2 \\ &= \delta_{\alpha\beta} - 4 \sum_i \sum_{j>i} U_{\alpha i} U_{\alpha j} U_{\beta i} U_{\beta j} \sin^2 \left( \frac{(m_i^2 - m_j^2) L}{4 E} \right). \end{aligned} \quad (1.7)$$

Equation (1.7) shows oscillatory behavior if two conditions are fulfilled. First, the mass eigenvalue of at least two mass eigenstates must differ. Second, the product of the PMNS matrix elements in the corresponding summand must not be zero. In other words, neutrino oscillations require at least one neutrino type to have mass and to undergo flavor mixing. Furthermore, neutrino oscillation experiments are sensitive to the difference of squared masses

$$\Delta m_{ij}^2 = |m_i^2 - m_j^2|, \quad (1.8)$$

which only yields two independent observables for three masses. Thus, these experiments cannot be used to determine the absolute mass scale of neutrinos.

### 1.4.2. Experimental Considerations

According to equation (1.7), the ratio  $L/E$  determines the sensitivity of an experiment to the oscillation parameters given by the PMNS matrix  $U$  (mixing parameters) and  $\Delta m_{ij}^2$  (mass ordering).  $L$  can be tuned by placing the detector in a suitable distance from a Earth-based neutrino source.  $E$  can either be tuned by using e.g. particle accelerators as source or it varies naturally, for instance if the source exhibits an energy spectrum like the Sun.

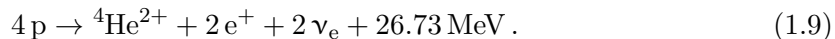
Furthermore, two detection channels can be distinguished. If, on one hand, an experiment is only sensitive to the neutrino flavor which is emitted by the neutrino source, it can detect a weakening of the neutrino flux. This is referred to as disappearance channel. If, on the other hand, an experiment is sensitive to a different neutrino flavor than the one which the neutrino source emits, this is referred to as appearance channel [Zub11].

There are four major classes of neutrino sources that can be used to measure the mixing parameters and the mass ordering. They are listed in table 1.1. For each class, multiple experiments exist [Zub11]. Not all experiments will be discussed here. Instead, the following section 1.4.3 discusses the so-called “solar neutrino problem” as an exemplary case study on solar neutrino experiments.

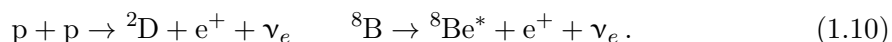
### 1.4.3. Experimental Case Study: The Solar Neutrino Problem

The term “Solar Neutrino Problem” refers to the mismatch of the total neutrino flux arriving on Earth as predicted by the standard solar model (SSM) and measured by, first, the Homestake and, later, other experiments. In this section, the problem is briefly described and its resolution developed.

At the end of the 1930s, Bethe, von Weizsäcker and Critchfield showed that there are two main fusion cycles in the Sun, the so-called CNO and pp cycle. The latter is the primary source of solar neutrinos [Wei38; BC38; Bet39]. Its multi-step reaction can be summarized as [Zub11]



Its initial reaction and the one with the broadest neutrino energy spectrum (from below 0.1 to  $\sim 11$  MeV [Tan+18]) are



It should be noted, that only electron neutrinos are produced in the Sun. Starting from the 1970s, the solar electron neutrino flux was measured; the first time by the Homestake

**Table 1.1.: Neutrino source classes for neutrino oscillation experiments.** Different neutrino source classes are listed and which neutrino flavors they emit [Zub11]. Furthermore, the oscillation parameters they are mainly sensitive to are tabulated along with an example experiment.

| source class         | flavors                                      | sensitive to   | example experiment        |
|----------------------|--|--|---------------------------|
| nuclear power plants | $\bar{\nu}_e$                                | $\sin \theta_{13}$   | Double Chooz [Abe+16]     |
| accelerators         | $\nu_e, \nu_\mu, \bar{\nu}_e, \bar{\nu}_\mu$ | $\sin \theta_{12}, \sin \theta_{23}, \Delta m_{12}^2, \Delta m_{23}^2$ | MiniBooNE [Agu+09]        |
| atmosphere           | $\nu_e, \nu_\mu, \bar{\nu}_e, \bar{\nu}_\mu$ | $\sin \theta_{23}, \Delta m_{23}^2$                                    | Super-KamiokaNDE [Fuk+98] |
| the Sun              | $\nu_e$                                      | $\sin \theta_{12}, \Delta m_{21}^2$                                    | SNO [Aha+13]              |

experiment using the inverse beta decay of  $^{37}\text{Cl}$ . It could detect electron neutrinos with an energy threshold of 813 keV. The measured flux was one third of the prediction by the SSM [Cle+98; BPB01]. This is marked as the beginning of the solar neutrino problem. The experiments GALLEX/GNO and SAGE confirmed the results, where the latter could detect electron neutrinos with an energy threshold of 233 keV [Kir98; Alt+05; Abd+09]. The low energy threshold is of importance because the neutrinos emitted by the initial reaction of the pp cycle (eq. 1.10) exhibit the highest flux, but at the same time an energy spectrum that ends at approximately 400 keV [Bah96].

Starting from 1999, the SNO experiment measured the neutrino flux of all flavors. It used 1000 t of heavy water  $\text{D}_2\text{O}$  to detect electron neutrinos via charged currents as well as all flavors via neutral currents and elastic neutrino-electron scattering. In order to fully explain the flux data, the so-called Mikheyev-Smirnov-Wolfenstein effect (MSW effect) had to be respected [Wol78; MS86]: Electron neutrinos can undergo charged current interactions with surrounding electrons in a coherent forward scattering process, which alters the flavor transition amplitude. This effect is only significant in areas of high electron densities, such as the Sun. Taking these matter-mediated oscillations into account, the measured flux of all flavors of the  $^8\text{B}$  neutrinos (eq. 1.10) was in accordance with the electron neutrino flux predicted by the SSM [Aha+13]. Thus, the solar neutrino problem was resolved after more than three decades.

#### 1.4.4. Summary of Experimental Results

This section summarizes the results obtained from neutrino oscillation experiments. The neutrino oscillation parameters consist of the squared neutrino mass differences (eq. 1.8) and the mixing parameters of the PMNS matrix (eq. 1.4). It should be noted first, that the MSW resonance of solar neutrinos requires  $m_1 < m_2$ , which still allows for two possible mass orderings [Zub11]:

1. normal ordering  $m_1 < m_2 < m_3$  and
2. inverted ordering  $m_3 < m_1 < m_2$ .

For these two cases, a combination of recent experimental results for the neutrino oscillation parameters is given in table 1.2 and illustrated in figure 1.2. All in all, neutrino oscillations are experimentally verified and provide unequivocal proof that neutrinos have mass.

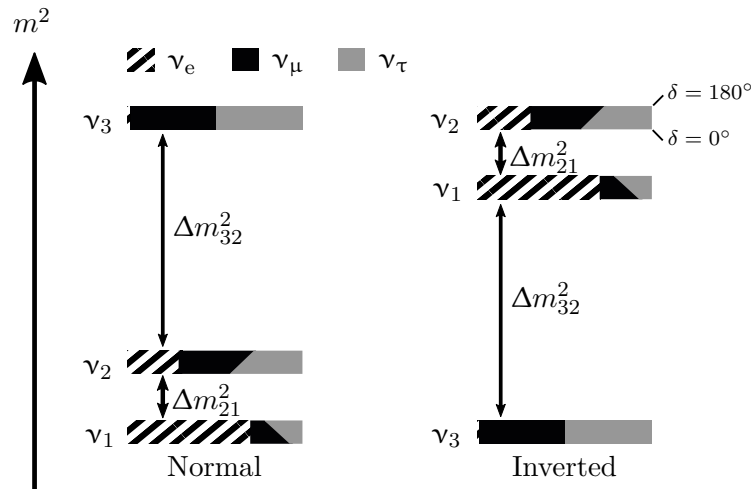
#### 1.5. Absolute Neutrino Mass Measurements

The absolute masses of neutrinos remain unknown and, as shown in section 1.4.1, neutrino oscillations are only sensitive to squared mass differences. This section presents methods to probe the absolute neutrino mass. The corresponding measurements fall into one of three categories [OW08]:

- observational cosmology (section 1.5.1),
- search for neutrinoless double  $\beta$  decay (section 1.5.2) or
- kinematic measurements of weak decays such as  $\beta$  decay and electron capture (section 1.5.3).

##### 1.5.1. Observational Cosmology

In the early Universe, neutral particles such as light neutrinos could escape from areas of high mass density to areas of low mass density. As they carry away mass, the larger the neutrino mass, the stronger is the suppression of density fluctuations on small scales.



**Figure 1.2.: Neutrino mixing parameters and mass ordering.** The chart shows how the mass eigenstates  $\nu_i$  are composed of the flavor eigenstates  $\nu_\alpha$  in the normal and inverted mass ordering. The composition depends on the phase  $\delta$ . The mixing is shown for the two extreme cases  $\delta = 0^\circ$  (baseline) and  $\delta = 180^\circ$  (topline). (Adapted from [Sei19]. Numerical values can be found in [Est+19] and table 1.2)

Corresponding data are obtained, for example, by the Sloan Digital Sky Survey (SDSS). This survey records the sky's optical and infrared spectrum via telescope in order to map the large scale distribution of galaxies and galaxy clusters [Dor+04]. Furthermore, the temperature anisotropies in the cosmic microwave background (CMB) encode information on the Universe's structure. The latest and most precise data are recorded by the Planck satellite [Agh+18]. Under the assumption that all mass states contribute with the same number density, cosmological observations are to first order only sensitive to the sum of all neutrino masses  $\sum_i m_i$ . A combination of the above data sets yields [Yèc+17]

$$\sum_i m_i < 0.14 \text{ eV} \quad (95\% \text{ C.L.}).$$

**Table 1.2.: Overview of neutrino oscillation parameters.** The table lists the observables and their best fit values along with the  $1\sigma$ -uncertainty range for normal and inverted ordering. (From [Est+19].)

| observable                                     | normal ordering                 | inverted ordering               |
|--|---------------------------------|---------------------------------|
| $\sin \theta_{13}$                             | $0.310^{+0.013}_{-0.012}$       | $0.310^{+0.013}_{-0.012}$       |
| $\sin \theta_{23}$                             | $0.580^{+0.017}_{-0.021}$       | $0.584^{+0.016}_{-0.020}$       |
| $\sin \theta_{13}$                             | $0.02241^{+0.00065}_{-0.00065}$ | $0.02264^{+0.00066}_{-0.00066}$ |
| $\delta/^\circ$                                | $215^{+40}_{-29}$               | $248^{+27}_{-29}$               |
| $\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$ | $7.39^{+0.21}_{-0.20}$          | $7.39^{+0.21}_{-0.20}$          |
| $\frac{\Delta m_{31}^2}{10^{-3} \text{ eV}^2}$ | $+2.525^{+0.033}_{-0.032}$      |                                 |
| $\frac{\Delta m_{32}^2}{10^{-3} \text{ eV}^2}$ |                                 | $-2.512^{+0.034}_{-0.032}$      |

### 1.5.2. Search for Neutrinoless Double- $\beta$ Decay

Double- $\beta$  decay ( $2\nu\beta\beta$ ) is described as a nucleus of element  $X(Z, A)$  with  $Z$  protons and  $A - Z$  neutrons that decays to a daughter isotope  $Y(Z + 2, A)$  via two simultaneous  $\beta$  decays

$$X(Z, A) \rightarrow Y(Z + 2, A) + 2e^- + 2\bar{\nu}_e. \quad (1.11)$$

If the neutrino is its own antiparticle, respectively of Majorana type, the neutrino emitted in the first decay can be absorbed in the second decay resulting in a neutrinoless double decay ( $0\nu\beta\beta$ ). This would require the neutrino to have mass. Such a decay would manifest itself in a peak in the  $\beta$  spectrum two neutrino masses above the endpoint of the continuum [Zub11]. It should be noted, that this would violate lepton number conservation. The half-life of such a decay encodes the Majorana mass of the electron neutrino as a coherent sum of all neutrino masses using the PMNS matrix  $U$  (eq. 1.4)

$$m_{\beta\beta}^2 = \left| \sum_i U_{ei}^2 m_i \right|^2. \quad (1.12)$$

As  $U$  contains two unknown Majorana phases, partial cancellation might occur. Hence, it is difficult to compare  $m_{\beta\beta}$  to masses obtained by other methods. The two most stringent upper limits on  $m_{\beta\beta}$  are listed in table 1.3.

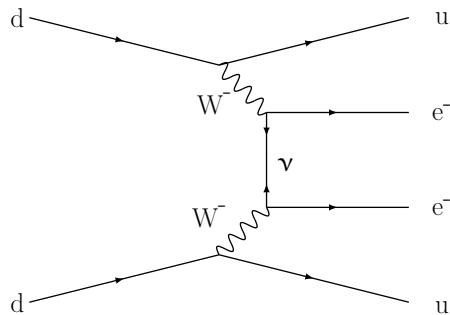
### 1.5.3. Kinematic Measurements of Weak Decays

Several laboratory experiments as well as the supernova event 1987A have provided upper limits of absolute neutrino masses from the analysis of kinematics of weak interactions involving neutrinos or neutrino time-of-flight considerations. Such experiments can not

**Table 1.3.: Constrains on the neutrino mass by double- $\beta$  decay experiments.** The two most stringent limits on  $m_{\beta\beta}$  (eq. 1.12) are given as ranges.

| experiment           | isotope           | upper limit (90 % C.L.) on $m_{\beta\beta}$ (eV) |
|----------------------|-------------------|--|
| GERDA [Ago+18]       | $^{76}\text{Ge}$  | 0.12–0.26  |
| KamLAND-Zen [Gan+16] | $^{136}\text{Xe}$ | 0.05–0.16  |





**Figure 1.3.: Feynman graph of neutrinoless double- $\beta$  decay.** The graph depicts the simultaneous transformation of two neutrons into two protons where the down quarks transform into up quarks, whilst two electrons and two neutrinos are produced. The two emitted neutrinos annihilate in a Majorana transition.

**Table 1.4.: Constraints on the neutrino mass by kinematic measurements.** The table lists upper limits on the absolute neutrino masses for different neutrino flavors.

| flavor        | measurement basis              | upper limit          | reference                    |
|---------------|--------------------------------|----------------------|------------------------------|
| $\nu_e$       | neutrinos from Supernova 1987A | 5.7 eV (95 % C.I.)   | [LL02]                       |
| $\nu_\mu$     | muon decay                     | 17 keV (90 % C.L.)   | [Ass+96]                     |
| $\nu_\tau$    | tau decay                      | 18.2 MeV (95 % C.L.) | [Bar+98]                     |
| $\bar{\nu}_e$ | tritium- $\beta$ decay         | 2 eV (95 % C.L.)     | [Kra+05], [Ase+11], [Tan+18] |

resolve the mass splitting between the squared mass eigenvalues. Therefore, the corresponding observable is an effective mass, respectively a weighted sum of the  $N$  neutrino eigenmasses, where the weights are the elements of the PMNS matrix (eq. 1.4) [OW08]

$$m_{\nu_\alpha}^2 = \sum_i^N |U_{\alpha i}|^2 m_i^2. \quad (1.13)$$

In the scope of this thesis, the measurement of the mass of the electron antineutrino via  $\beta^-$ -decay kinematics is of special interest. Hence, this subject is examined more closely within this section. For completeness, aside from the upper limit on the effective mass of the electron antineutrino, table 1.4 also lists upper limits for other neutrino flavors obtained by kinematic measurements.

### Neutrino Masses from $\beta$ -decay Kinematics

In  $\beta^-$  decay

$$X(Z, A) \rightarrow Y(Z + 1, A) + e^- + \bar{\nu}_e \quad (1.14)$$

part of the released surplus energy generates the neutrino's mass. This leaves a signature in the  $\beta$  spectrum (see section 3.1 for a quantitative description). In a neutrino mass experiment four criteria are important for a suitable  $\beta$  emitter [OW08]:

- The  $\beta$  emitter should have an energy spectrum with a relatively low endpoint because in the uncertainty on the neutrino mass enter the input uncertainties of the neutrino energy and momentum scaled up with the endpoint energy.
- The  $\beta$  emitter must have a sufficiently high activity to provide statistically relevant count rates for quantities of the  $\beta$  emitter that can be handled in the laboratory.
- The  $\beta$  decay should be super-allowed in order for the nuclear matrix element of the decay process to be energy independent.

- The  $\beta$ -emitter molecule should be as simple as possible to allow for a theoretical treatment of its decay kinematics such as the energy state of the daughter molecule after the decay.

Tritium is an ideal candidate with respect to these criteria [OW08]. The corresponding measurement principle will be explained more closely in the following chapters about the KATRIN experiment. However, KATRIN has several predecessor experiments. The most recent two experiments based on tritium- $\beta$  decay in Mainz and Troitsk obtained a combined upper limit on the electron antineutrino mass of [Kra+05; Ase+11; Tan+18]

$$m_{\bar{\nu}_e} < 2 \text{ eV} \quad (95\% \text{ C.L.}).$$

It should be noted, that KATRIN aims for a sensitivity that is better by one order of magnitude.

## 2. The KATRIN Experiment

The Karlsruhe TRItium Neutrino (KATRIN) experiment performs a kinematic measurement of the tritium- $\beta$  spectrum in order to determine the effective mass of the electron antineutrino (from here forth labeled  $m_\nu$  and called neutrino mass) as defined by equation (1.13). In case no neutrino mass signal is observed, KATRIN aims to set an upper limit of

$$m_\nu < 200 \text{ meV} \quad (90 \% \text{ C.L.}),$$

which is one order of magnitude more constraining than the one set by its predecessor experiments. KATRIN recorded the first  $\beta$  spectrum in a commissioning run in May 2018 and started neutrino mass measurements in March 2019.

This chapter provides an overview of the KATRIN apparatus. However, given KATRIN's complexity, it can by no means be exhaustive and for a comprehensive treatment the reader is referred to the KATRIN Design Report [KAT05] supplemented by an up-to-date hardware overview that is in the making at the time of writing this thesis<sup>1</sup>.

### 2.1. Overview of the KATRIN Experimental Setup

The KATRIN experiment comprises a 70-m-long beam line depicted in figure 2.1. It can be divided into two sections:

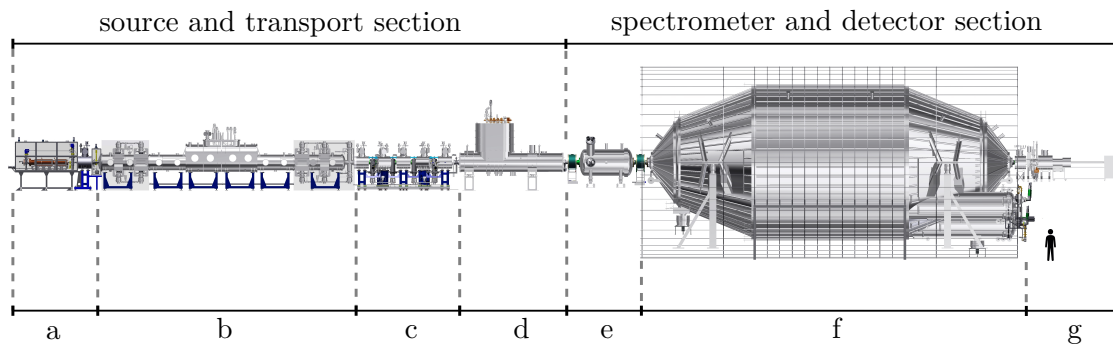
1. The **source and transport section (STS)** comprises i. a. the gaseous tritium source where the tritium decays and the  $\beta$  electrons are magnetically guided along the beam line. Furthermore, the gas flow from the tritium source to the exit of the STS is reduced by at least 14 orders of magnitude.
2. In the **spectrometer and detector section (SDS)** the  $\beta$  electrons are filtered according to their kinetic energy and finally counted at the detector.

A central concept of the KATRIN setup is the magnetic flux tube. The  $\beta$  electrons must be guided from their point of origin to the detector. Therefore, a magnetic field is created by superconducting solenoids surrounding the beam line in the STS as well as coils around, and superconducting solenoids before and after the spectrometer tank in the SDS. The field lines intersperse the beam line over the range of the whole experiment. The volume that is mapped onto the detector by this mechanism is called the flux tube. Within the flux tube, charged particles perform cyclotron motions around the field lines and are adiabatically guided from the STS to the detector. Adiabaticity is guaranteed by avoiding strongly varying field strengths on short distances.

A further central concept of KATRIN is the windowless source. As  $\beta$  electrons must not lose energy before energy analysis takes place, the SDS is windowlessly connected to the STS. However, the spectrometer must be kept practically free of any tritium flow

---

<sup>1</sup>K. Altenmüller et al. (KATRIN collaboration), in prep.



**Figure 2.1.: The KATRIN beamline.** Shown are the main hardware components:

- a) rear section (see section 2.3)
- b) windowless gaseous tritium source (WGTS) (see section 2.2)
- c) differential pumping section (DPS) (see section 2.4)
- d) cryogenic pumping section (CPS) (see section 2.5)
- e) pre spectrometer (see section 2.6)
- f) main spectrometer (see section 2.6)
- g) detector section (see section 2.7)

for safety reasons and to keep the strict background requirements. Therefore, pumping systems reduce the gas inlet pressure of  $1.8 \text{ mbar}\ell/\text{s}$  to the tritium partial pressure of below  $10^{-14} \text{ mbar}\ell/\text{s}$  of the spectrometer.

The following sections step through the various components along the KATRIN beam line describing their functionality and purpose.

## 2.2. Windowless Gaseous Tritium Source

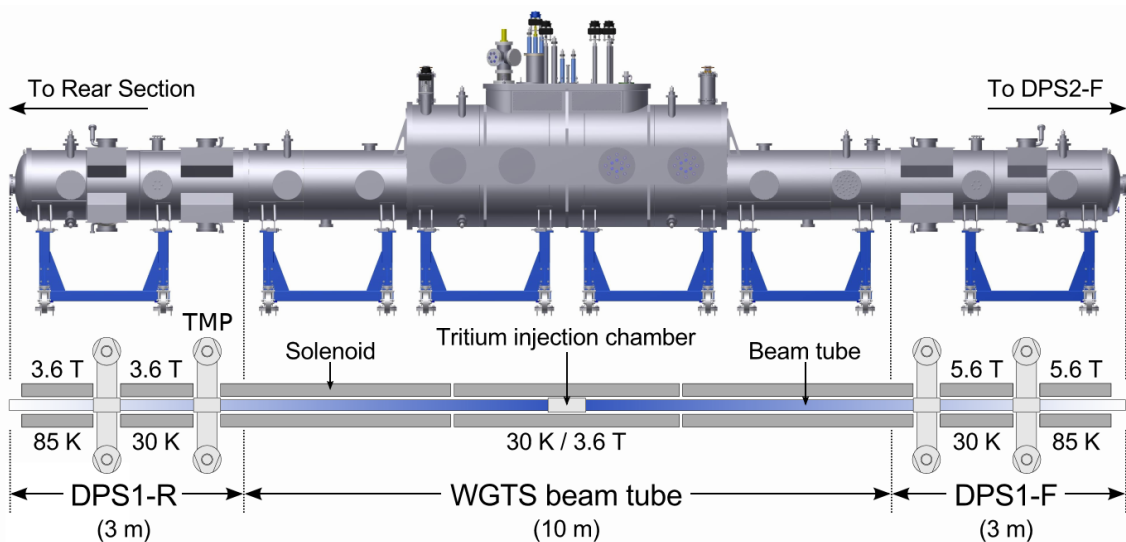
The windowless gaseous tritium source (WGTS) is a 16-m-long, 1.5-m-wide and 4-m-high cryostat. It is depicted in figure 2.2 and a detailed description can e. g. be found in [Gro+08; Bab+12]. In the following, the major features of the WGTS are reviewed:

**Tritium purity:** The molecular tritium ( $\text{T}_2$ ) is injected in the middle of the 10-m-long beam tube of 90 mm diameter, where it decays. The design gas column density is  $\rho d = 5 \times 10^{17} \text{ molecules/cm}^2$  with an isotropic tritium purity of  $\epsilon_{\text{T}} = 95 \%$  [KAT05] or better. At the front and rear of the WGTS, the gas is extracted from the beam tube by turbo molecular pumps in designated differential pumping sections called DPS-1-R (rear) and DPS-1-F (front). The extracted gas is re-injected in the center of the beam tube. The respective circulation system is called the inner loop [PSB15]. The tritium purity  $\epsilon_{\text{T}}$  must be kept stable on a 0.1 % level [KAT05]. Therefore, a permeator is installed that separates impurities (like e. g. helium) and ejects them into the exhaust loop of the Tritium Laboratory Karlsruhe (TLK). Furthermore, the isotopic composition of the gas is monitored by a designated laser Raman system (LARA) [Sch13].

**Injection pressure:** The design injection pressure of the tritium gas is  $1.8 \text{ mbar}\ell/\text{s}$ . It must be kept stable at the 0.1 % level. This is achieved via a pressure- and temperature-controlled buffer vessel within the inner loop [PSB15].

**Magnetic field:** In order to adiabatically guide the  $\beta$  electrons to the spectrometer section, the WGTS is pervaded by a magnetic field parallel to its beam tube axis of up to 5.6 T. It is created by seven superconducting coils, that surround the beam tube. These magnets are kept at a temperature of 4.2 K by liquid helium [Are+18b].

**Temperature:** On the one hand, thermal motion smears the energy spectrum of the  $\beta$  electrons (Doppler effect). On the other hand, at low temperatures, the gas molecules



**Figure 2.2.:** The windowless gaseous tritium source (WGTS). The hull and a sketch of the beam tube are shown. Indicated are the 8 turbo molecular pumps (TMP), the 7 magnets, the design temperatures for tritium operation, the maximum magnetic field strengths and a gradient within the beam tube depicting the decreasing gas density from the center to the sides. (Adapted from [Har15].)

start to cluster.  $T = 30$  K is chosen as a compromise and established by a two-phase neon cooling system. For calibration purposes, it is also possible to operate the WGTS with krypton-83m instead of tritium. This requires a beam tube temperature of  $T = 100$  K in order for the krypton not to freeze. In this operational mode, the neon has to be exchanged for argon [KAT05].

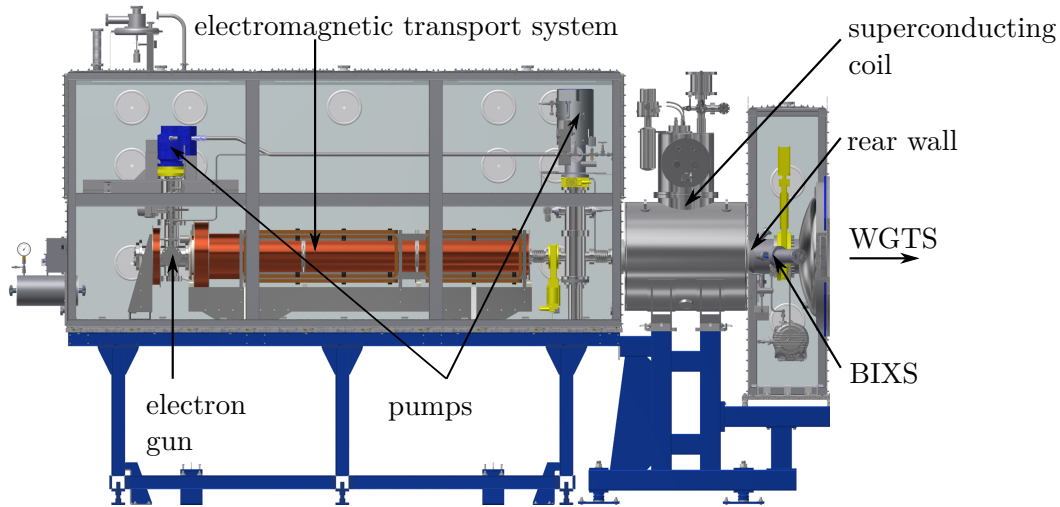
### 2.3. Rear Section

The rear section (RS) terminates the beam line in the upstream direction and houses monitoring, calibration and control devices. It is depicted in figure 2.3 and a detailed description can e. g. be found in [Bab14]. The major features of the RS are reviewed in the following:

**Electron gun:** The RS houses an electron gun in order to measure the response function of the experiment (see section 3.2) via an electron source with a well-defined energy resolution of  $\sim 0.2$  eV and angular resolution of  $\sim 4^\circ$ . The electrons are guided towards and through the rear wall by a designated electromagnetic guidance system. Furthermore, their flight path can be adjusted by dipole magnets mounted in the WGTS which enable a scanning of the full flux tube [Bab14].

**Rear wall and plasma control:** The so-called rear wall is a gold-coated stainless-steel disc with a diameter of 6 inches that terminates the beam tube. It has a hole in the center to let electrons from the electron gun pass through. Its main purpose is the control of plasma effects: Space charges, respectively a plasma, forms within the WGTS due to the tritium decay. Therefore,  $\beta$  electrons may start at different potentials which adds uncertainty to the measured  $\beta$  spectrum. Simulations show that the plasma can be influenced by the rear wall potential, which can be controlled via a voltage supply in the range of  $\pm 10$  V. Moreover, a UV light illumination of the rear wall can provide a large flux of electrons, extracted via the photoelectric effect, that can compensate space charges. Therefore, a homogeneous work function of the rear wall with fluctuations less than 20 meV is required [Kuc+18; Kuc16].

**Activity monitoring:** A superconducting coil designed to create a magnetic field of 4.7 T in the RS ensures that the magnetic flux tube terminates at the rear wall. Hence, per design



**Figure 2.3.:** The rear section (RS) terminates the KATRIN beam line and houses several monitoring and calibration devices that are described in the main text. (Adapted from [Sei19].)

of the magnetic guidance,  $\beta$  electrons either arrive at the detector or hit the rear wall. On hitting the rear wall, they emit bremsstrahlung. Two designated beta-induced X-ray spectroscopy (BIXS) systems measure the corresponding X-ray spectrum to determine the source strength respectively the gas column density [Röl15].

## 2.4. Differential Pumping Section

The differential pumping section (DPS) is composed of five elements. It is depicted in figure 2.4 and a detailed description can e. g. be found in [Kos12]. For orientation, in this section, the elements are labeled 1 to 5 from WGTS to cryogenic pumping section (CPS). The major features of the differential pumping section (DPS) are reviewed in the following:

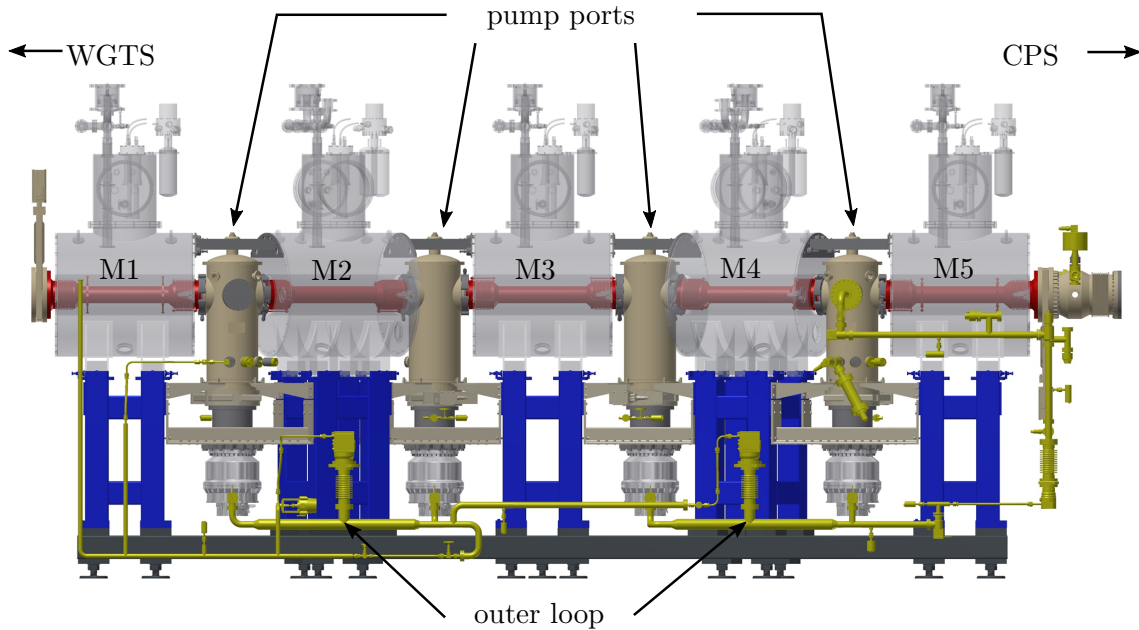
**Reduction of tritium flow:** The five beam tube elements of the differential pumping section (DPS) form a  $20^\circ$  angle to each other and are arranged in a chicane.  $\beta$  electrons are magnetically guided along the chicane by a magnetic field of up to 5.5 T created by five superconducting solenoids. By contrast, the neutral gas molecules scatter off the walls. This reduces the molecular beaming effect and enhances the pumping probability [ZJH12]. Four turbo molecular pumps mounted between the beam tube elements then reduce the gas flow by approximately five orders of magnitude and feed the gas into the so-called outer loop where it is reprocessed [Kos12].

**Ion blocking:** In the WGTS, ions such as  $\text{HeT}^+$ ,  $\text{T}_2^+$ ,  $\text{T}_3^+$ ,  $\text{T}_5^+$  can form. If they were not blocked, they would reach the spectrometer section together with the  $\beta$  electrons and would be even accelerated by the retarding voltage (see section 2.6). This would eventually lead to an increased background rate. A potential barrier created by two ring electrodes in element 5 and the pump port between DPS and CPS set to +100 V avoids such a scenario. The positive ions are deflected, and dipole electrodes in the elements 1 to 4 make them drift out of the flux tube. They hit the wall and get neutralized [Klein19].

**Ion monitoring:** Downstream of the blocking electrodes, the remaining ion flux is measured by a Fourier transform ion cyclotron resonance device (FT-ICR) [Ubi+09].

## 2.5. Cryogenic Pumping Section

The cryogenic pumping section (CPS) is an approximately 7-m-long cryostat. It is depicted in figure 2.5 and a detailed description can e. g. be found in [Jan15]. For orientation, in this



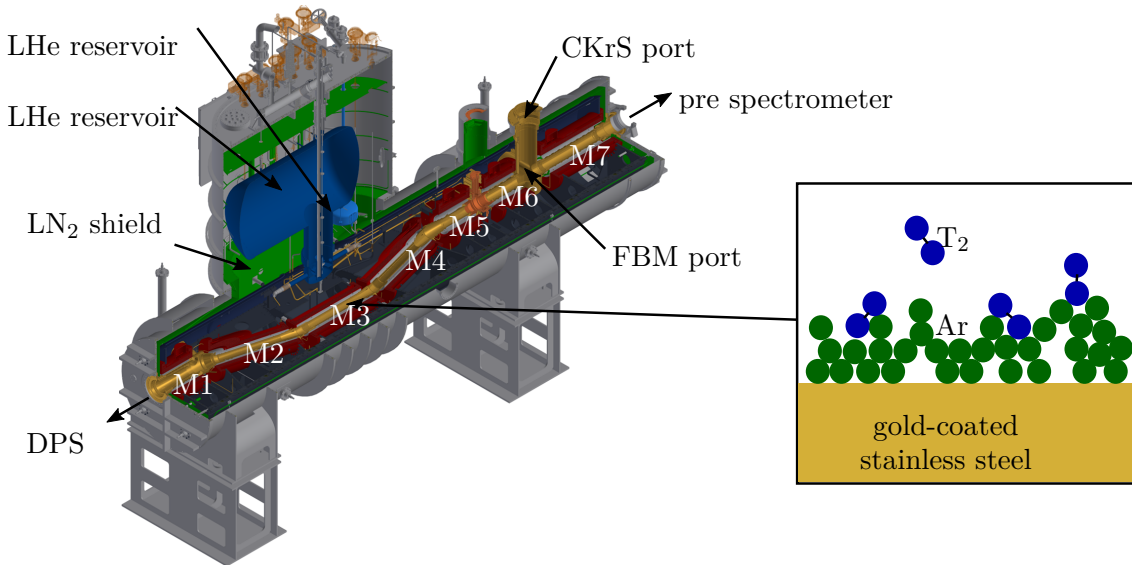
**Figure 2.4.:** The differential pumping section (DPS) reduces the gas flow by five orders of magnitude and blocks tritium ions. Its five elements, each with a separate magnet (M1-M5), and connected by pump ports, are shown. The outer loop connects the turbo molecular pumps to the infrastructure of the TLK. (Adapted from [Sei19].)

section, its seven elements are labeled by 1 to 7 from WGTS to CPS. The major features of the CPS are reviewed in the following:

**Reduction of tritium flow:** The CPS consists of seven beam tube elements, of which the first five are arranged in a chicane forming  $15^\circ$  angles, in a similar manner as the beam tube elements of the DPS. Charged particles are guided along the chicane by a magnetic field of up to 5.6 T created by seven superconducting coils. Neutral molecules hit the walls, that are covered by a frozen argon layer cooled down to 3 K in order to cold-trap particles. These low temperatures are achieved via liquid helium cooling and a heat shield of liquid neon. After the accumulation of about 1 Ci of tritium, the argon frost layer has to be renewed. In order to achieve this, the beam tube is warmed up and the argon is pumped off along with the accumulated tritium. Tests and simulations show a reduction of the tritium flow by approximately 10 orders of magnitude from the entrance to the exit of the CPS [Jan15; Röt19].

**The forward beam monitor (FBM):** The FBM can be moved horizontally into the pump port between beam tube element 6 and 7 of the CPS with a 2-dimensional spatial resolution of 0.1 mm. Two *pin*-diodes measure the  $\beta$ -electron flux and thus the stability of the gas column density in the WGTS. Furthermore, the forward beam monitor (FBM) equips a temperature and a hall sensor. A second detector board holding a Faraday cup for ion measurements is also available [Klein19]. More information about the FBM can e. g. be found in [Ell+17; Ell19].

**The condensed  $^{83m}\text{Kr}$  source (CKrS):** The CKrS is a sub mono-layer of  $^{83m}\text{Kr}$  on a pyrolytic graphite substrate with a diameter of 2 cm. It can be lowered in the pump port of the CPS and moved in a 2-dimensional plane perpendicular to the beam line. This enables the spatial scanning of the properties of the spectrometer using quasi-monoenergetic conversion electron lines of  $^{83m}\text{Kr}$  [Bau14; Dyb19; Are+18a].



**Figure 2.5.:** The cryogenic pumping section (CPS) is the coldest part of the KATRIN experiment. It consists of seven elements, labeled by 1 to 7 from the DPS to the pre spectrometer. Elements 2 to 5 are covered by a frozen argon layer at 3 K in order to cold-trap tritium molecules. The low temperatures are established using liquid helium (LHe) and an insulation of liquid nitrogen ( $\text{LN}_2$ ). Each element is enclosed by a superconducting coil (M1 to M7) for magnetic guidance of the  $\beta$  electrons. For the FBM and the CKrS the reader is referred to the main text. (Adapted from [Sei19].)

## 2.6. Pre and Main Spectrometer

The pre and main spectrometer are vacuum vessels designed to filter passing electrons according to their kinetic energy. The pre spectrometer has a length of 3.4 m and a diameter of 1.7 m. Details on its design can e. g. be found in [Val09; Frä10]. The main spectrometer has a length of 23 m and a diameter of 10 m. Details on its design can e. g. be found in [Val09; Val04]. The functionality and purpose of spectrometer related aspects are reviewed in this section. The content is divided into two parts: Section 2.6.1 explains the so-called MAC-E filter principle and section 2.6.2 list several measures to keep the strict KATRIN background requirements.

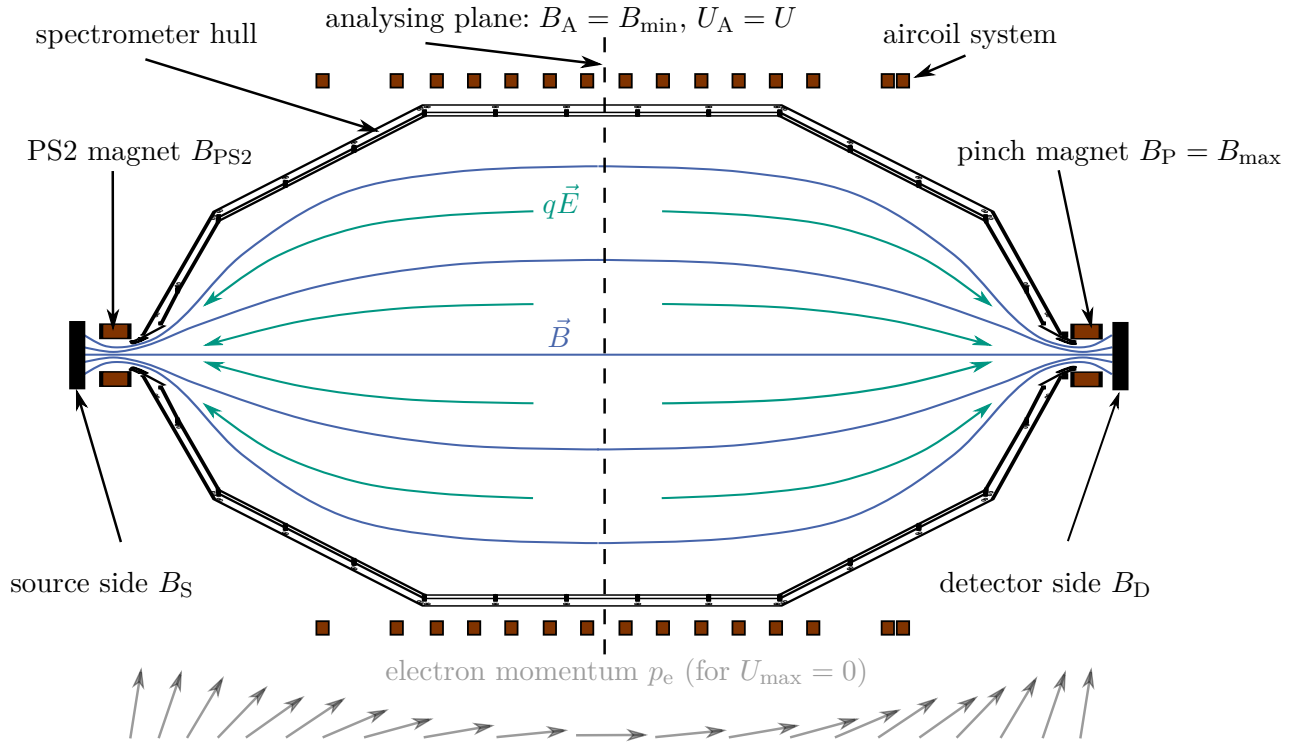
### 2.6.1. MAC-E-Filter Principle

The pre and main spectrometer are based on the principle of the so-called magnetic adiabatic collimation with electrostatic filtering (MAC-E) [BPT80]. It enables the filtering of electrons according to their kinetic energy. Figure 2.6 sketches the MAC-E filter of KATRIN. The following paragraphs outline the basic concepts and their experimental implementation. As the principle is of key importance for the KATRIN experiment, it is additionally treated in a mathematical way in the subsequent section 3.2.3.

**Electrostatic filtering:** A retarding voltage barrier is applied along the beam axis within the spectrometer, reaching its maximum  $U$  at the so-called analyzing plane in the center and dropping off towards the source section and the detector. The retarding voltage barrier deflects electrons with kinetic energies below  $eU$ . For higher energies, whether an electron can pass the spectrometer or not depends on the angle between its direction of motion and the magnetic field lines (see eq. 3.13 for a quantitative description).

**Magnetic collimation:** The electric field gradient of the retarding voltage barrier is parallel to the beam line, but  $\beta$  electrons are emitted in an arbitrary angle with respect





**Figure 2.6.:** Scheme of the KATRIN main spectrometer and the magnetic adiabatic collimation with electrostatic filtering (MAC-E) filter principle. The KATRIN design magnetic field settings are  $B_{PS2} = 4.5$  T,  $B_S = 3.6$  T,  $B_{max} = 6.0$  T,  $B_D = 3.6$  T,  $B_A \approx 3 \times 10^{-4}$  T.  $\vec{E}$  denotes the magnetic field regulated by the retarding potential  $U$  that reaches its maximum  $U_a = U$  at the analysing plane. (Adapted from [Sei19].)

to the magnetic field lines. In order to analyze their full kinetic energy, they have to be collimated. This is achieved by a magnetic field gradient that drops from  $B_S = 3.6$  T in the STS to  $B_A \approx 3 \times 10^{-4}$  T in the analysing plane. In the following, a plausibility argument for the momentum collimation due to the field gradient, according to [KAT05], is given: Electrons entering the spectrometer vessel perform cyclotron motions around the magnetic field lines. Their total kinetic energy  $E_S$  is split into a longitudinal component  $E_{\parallel}$  along the beam axis and a transverse component  $E_{\perp}$

$$E_S = E_{\parallel} + E_{\perp}. \quad (2.1)$$

In the non-relativistic and adiabatic approximation, the transverse component can be expressed by the magnetic field strength  $B$  and the electron's magnetic moment  $\mu$  respectively its charge  $q = e$ , its mass  $m_e$  and angular momentum  $L$  [Jac75]

$$E_{\perp} = -\mu B = \frac{e}{2m_e} L B. \quad (2.2)$$

Adiabaticity conserves the angular momentum  $L$  and the total energy of the electron  $E_S$  along its trajectory. Hence, when the magnetic field strength  $B$  decreases to  $B_A = B_{min}$  in the analysing plane, the transverse component of the electron's energy  $E_{\perp}$  decreases likewise and transforms to longitudinal energy  $E_{\parallel}$ .

**Magnetic Bottle effect:** As the source is placed in a lower magnetic field  $B_S$  compared to the maximum field strength along the beam line  $B_{max}$  at the detector side,  $\beta$  electrons traveling downstream to the detector are subject to the magnetic bottle effect [KAT05].

They get reflected and travel upstream to the rear wall if their starting angle  $\theta_S$  with respect to the beam line axis surpasses  $\theta_{\max}$  with

$$\sin \theta_{\max} = \sqrt{\frac{B_S}{B_{\max}}}. \quad (2.3)$$

For the KATRIN design values  $B_{\max} = 6 \text{ T}$  and  $B_S = 3.6 \text{ T}$  one obtains  $\theta_{\max} \approx 51^\circ$ . A cutting angle  $\theta_{\max}$  is beneficial because the greater the emission angle of a  $\beta$  electron the larger the distance it travels in the WGTS and the more it is subject to energy losses such as scattering or synchrotron radiation [KAT05].

**MAC-E-filter width:** Electrons with a kinetic energy below  $qU$  cannot pass the spectrometer. Electrons with a kinetic energy above  $qU + \Delta E$  do pass the spectrometer. Here,  $\Delta E$  denotes the filter width [KAT05]

$$\Delta E = \frac{B_A}{B_{\max}} E. \quad (2.4)$$

Electrons with an energy between  $eU$  and  $eU + \Delta E$  pass the potential barrier only with a certain probability. A quantitative description of this so-called transmission probability is given in the subsequent section 3.2.3. However, it can already be deduced, that a larger  $\Delta E$  adds a greater uncertainty to the measurement and thus it should be kept as low as possible.  $\Delta E$  depends on the maximum magnetic field strength along the beam line  $B_{\max} = 6 \text{ T}$ , the kinetic energy of  $\beta$  electrons  $E \approx 18.6 \text{ keV}$ , and the magnetic field in the analyzing plane  $B_A \approx 3 \times 10^{-4} \text{ T}$ . Hence, its KATRIN design value is  $\Delta E \approx 0.93 \text{ eV}$ .

**Dimensions of the KATRIN main spectrometer:** This paragraph outlines, why the diameter of the KATRIN main spectrometer is 10 m, while the one of its predecessor experiment in Mainz was only 1 m [Kra+05]. KATRIN's envisaged sensitivity requires a relative MAC-E-filter width of at least  $\Delta E/E = 1/20000$ , which directly corresponds the ratio of the magnetic fields  $B_{\max}/B_A$  (see eq. 2.4). For a smaller  $\Delta E$ ,  $B_A$  should be chosen as low as possible. However, the lower  $B_A$ , the wider the flux tube that must be governed by the spectrometer vessel. Also, the magnetic field must decrease at a sufficiently slow rate from the spectrometer's entrance to the analyzing plane in order to guarantee adiabaticity, which requires a certain spectrometer length. Dimensions that meet the demands and are feasible for the main spectrometer were found to be a radius of 10 m and a length of 23 m [KAT05; Val04].

Given the requirements on the magnetic and electrostatic fields, the following two paragraphs review their technical implementation:

**Magnetic field:** The main spectrometer is surrounded by a system of coils that shapes the MAC-E filter's magnetic field. Upstream, there is the PS2 magnet ( $B_{\max} = 4.5 \text{ T}$ ); downstream the pinch ( $B_P = B_{\max} = 6.0 \text{ T}$ ) as well as the detector magnet ( $B_D = 3.6 \text{ T}$ ), which are superconducting solenoids. The field is fine-tuned by a system of air coils around the spectrometer hull: There is the Earth magnetic field compensation system (EMCS) with 26 current loops parallel to the beam line axis. Furthermore, there is the low-field correction system (LFCS) with 14 air coils perpendicular to the beam line axis. The combined system constrains the electrons' flux tube to the spectrometer vessel and compensates the Earth's magnetic field as well as effects from ferromagnetic materials in the spectrometer's surroundings [Erh+18]. Additionally, a vertical and radial magnetic measuring system (VMMS and RMMS) are installed outside the spectrometer vessel. The field inside the spectrometer vessel is assessed via samples of these measuring systems combined with simulations [Let+18].

**Electrostatic field:** A high-voltage system establishes the MAC-E filter's retarding potential. The fluctuation of the retarding voltages must have a standard deviation smaller

than 60 mV for the envisaged sensitivity to the neutrino mass [KAT05]. The antenna-like beam line setup is sensitive to electromagnetic fluctuations of any source, which is why an active post-regulation system for the voltage is deployed. It monitors the retarding potential and regulates it with the required precision. For the monitoring the monitor spectrometer and a voltage divider are deployed. For details on the later systems the reader is referred to [Thü+09; Erh+14; Zbo11].

### 2.6.2. Background Mitigation Strategies

The KATRIN sensitivity goal requires a background rate of less than 10 mcps [KAT05]. Several background-related aspects with respect to the spectrometer tanks are listed below:

**Vacuum:** The spectrometers are operated at a pressure on the order of  $1 \times 10^{-11}$  mbar. This prevents electron scattering on residual gas and minimizes background effects by ionization. Correspondingly, turbo molecular and getter pumps are installed at three pump ports of the spectrometer vessels. Furthermore, the spectrometers can be baked out at up to 350 °C [Are+16].

**Wire electrodes:** The inner walls of the spectrometer vessels are lined by wire electrodes. Their potential is a few hundred volts more negative than the spectrometer hull reflecting electrons coming from the vessel walls. Such electrons may be induced by cosmic rays [Val09].

**Ion blocking:** Analogously to the ones in the CPS (section 2.5), three blocking electrodes are installed; one between the CPS and the pre spectrometer, one between the pre and main spectrometer; and one between the main spectrometer and the detector [Klein19].

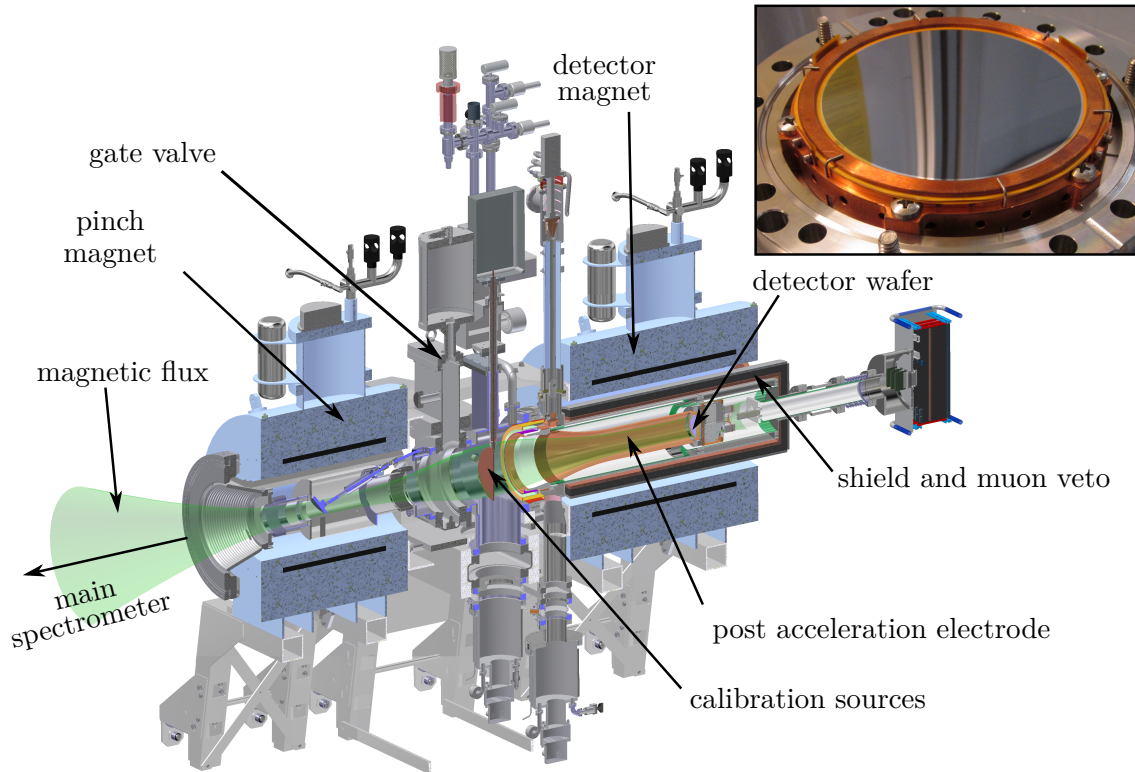
**Tandem setup:**  $\beta$  electrons may scatter on residual gas. This can either directly lead to secondary electrons or create positive ions that travel down the beam line. The positive ions in turn may again yield secondary electrons through scattering. The more  $\beta$  electrons enter the main spectrometer, the higher is the probability to create secondary electrons. In order to reduce the flux of  $\beta$  electrons into the main spectrometer, the retarding potential of the pre spectrometer is set to a few hundred volts more positive than the one of the main spectrometer. On the one hand, this is a countermeasure against background events, but, on the other hand, charged particles can be trapped between the two spectrometers due to the electromagnetic setup (Penning trap). A sudden discharge may harm the hardware, especially the detector. Therefore, it is possible to sweep a charged wire through the volume in order to collect the trapped particles and avoid this “Penning discharges” [Val09].

## 2.7. Detector Section

The detector section terminates the beam line in downstream direction. It can be separated from the spectrometer section by closing a gate valve. The detector section is depicted in figure 2.7 and a detailed description can e.g. be found in [Ams+15]. The major features of the detector section are reviewed in the following:

**Focal plane detector (FPD):** The FPD counts the  $\beta$  electrons that pass the spectrometer section. It is a *pin*-silicon detector with a sensitive area of 9 cm diameter. It is subdivided in 148 pixels of the same area arranged in 12 rings of 12 pixels each and the so-called bull’s eye of 4 pixels in the center. This arrangement allows for later correction for radial electrical, magnetic and gas dynamical inhomogeneities in the beam line [Ams+15].

**Shield and veto system:** The radiation shield of the FPD system consists of two nested cylindrical shells: an outer lead shell of 3 cm, that reduces photon background and an inner copper shell of 1.27 cm, that blocks X-rays originating from the outer lead shell. The shield is surrounded by a veto system to tag incoming muons. Such a system is necessary to keep the strict background requirements [Ams+15].



**Figure 2.7.:** The detector section terminates the KATRIN beam line. Among other instruments, it houses the focal plane detector (FPD) for  $\beta$  electrons with the detector wafer at its core. For an explanation of the other components the reader is referred to the main text. (Adapted from [Sei19].)

**Calibration:** Photoelectron sources can be lowered in the line of sight of the detector. The corresponding photocurrent can be measured with the precision ultra-low current integrating normalization electrometer for low-level analysis (PULCINELLA) system. A comparison of PULCINELLA and the FPD yields the FPD’s detection efficiency. It was determined to be  $\epsilon_{\text{det}} = 95 \pm 1.8 \pm 2.2\%$  [Ams+15].

**Detector magnet:** The detector magnet ( $B_D = 3.6\text{ T}$ ) allows to form the flux tube near the detector independently of the main spectrometer magnetic field setting. It especially allows for the mapping of the flux tube onto the detector [Ams+15].

**Post-acceleration electrode:** The post-acceleration potential shifts the electrons arriving from the main spectrometer to a more favorable energy region. This increases the detector efficiency and, additionally,  $\beta$  electrons can be distinguished from noise originating in the detector by an energy region of interest cut. An appropriate setting was found to be  $\sim 10\text{ keV}$  [Ams+15].

### 3. Mathematical Formalism of a KATRIN Measurement

For neutrino mass inference from simulated or real data, a mathematical model of a KATRIN neutrino mass measurement is required. As parameter inference is of relevance for the topic of this thesis, a mathematical formalism describing a KATRIN neutrino mass measurement is outlined within this chapter. Section 3.1 gives an expression for the tritium  $\beta$ -decay rate. Section 3.2 describes the KATRIN response function, respectively the mathematical modeling of the KATRIN apparatus. Sections 3.3 combines the concepts of the two preceding sections into the  $\beta$ -electron rate at the KATRIN detector. Section 3.4 incorporates the detector efficiency, the background rate and the measurement time in order to translate the rates into an expression for the electron counts measured by the KATRIN detector for a fixed retarding potential. Section 3.6 introduces the concept of a measurement time distribution over retarding potentials. Finally, section 3.7 shows a full simulated KATRIN neutrino mass measurement as obtained by the presented formalism.

#### 3.1. Differential Tritium- $\beta$ -Decay Spectrum

This section presents a quantitative expression for the  $\beta$ -decay rate within a tritium molecule normalized to a tritium nuclei in dependence on the kinetic energy of the emitted  $\beta$  electron (differential rate). First the whole mathematical description is denoted, then its components are explained.

Using Fermi theory and Fermi's golden rule the decay rate of a tritium molecule is [Kle+19; OW08]

$$\frac{d\Gamma(E_S)}{dE_S} = \frac{G_F^2 |V_{ud}|^2}{2\pi^3} |M_{\text{nuc}}|^2 \cdot F(Z, E_S) \cdot p(E_S + m_e) \cdot \sum_f P_f \cdot \epsilon_f \cdot \sqrt{\epsilon_f^2 - m_\nu^2} \cdot \Theta(\epsilon_f - m_\nu). \quad (3.1)$$

Its constituents are the kinetic electron energy  $E_S$ ; the effective squared electron-antineutrino mass  $m_\nu^2$  defined via the PMNS matrix  $U$  (eq. 1.4),

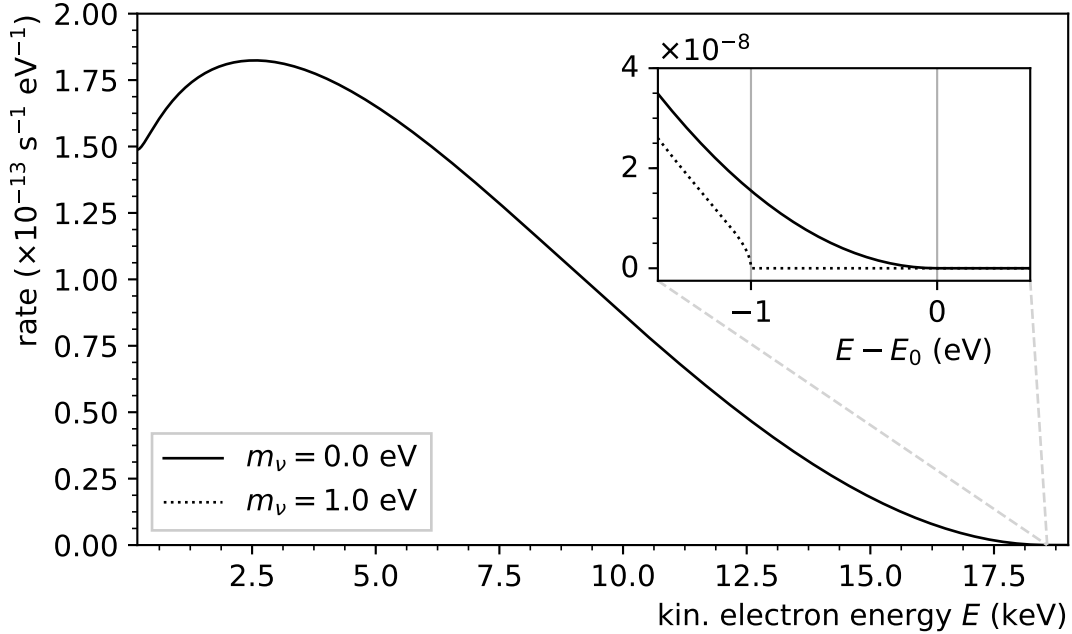
$$m_\nu^2 = \sum_i |U_{ei}|^2 m_i^2; \quad (3.2)$$

the Fermi constant  $G_F$ ; the up-down-quark-coupling given by the Cabibbo angle  $\theta_C$  [Tan+18]

$$V_{ud} = \cos \theta_C = 0.97425 \pm 0.00022; \quad (3.3)$$

and the nuclear transition matrix element [Tan+18]

$$|M_{\text{nuc}}|^2 = g_V^2 + 3g_A^2 \quad \text{with } g_V = 1 \quad \text{and} \quad g_A/g_V = -1.2646 \pm 0.0035 \quad (3.4)$$



**Figure 3.1.: Tritium- $\beta$  spectrum for a vanishing and non-vanishing neutrino mass.** The graph shows the differential rate as described by equation (3.1) for a vanishing and non-vanishing neutrino mass. The inset zooms into the endpoint region where a non-vanishing mass causes a shift and a distortion of the spectrum. (Calculated with [S-SSC].)

which is independent of the electron's kinetic energy (as the decay is super-allowed) and given by the vector  $g_V$  and axial vector  $g_A$  coupling.

Furthermore, the Fermi function  $F(Z, E_S)$  accounts for the Coulomb interaction between the outgoing electron and the daughter nucleus with atomic charge  $Z = 2$ , which in its relativistic version can be approximated as [Sim81]

$$F(Z, E_S) \approx \frac{2\pi\eta}{1 - \exp 2\pi\eta} \cdot R, \quad (3.5)$$

with Sommerfeld parameter  $\eta = \alpha Z/\beta$ , fine structure constant  $\alpha$ , relativistic velocity  $\beta$  and a relativistic correction factor  $R = 1.002037 - 0.001427\beta$ .

The phase-space factor of the outgoing electron with momentum  $p$  and mass  $m_e$  is given by the factor  $p(E_S + m_e)$ .

The phase space factor of the emitted neutrino depends on multiple quantities: First, there is the  $\beta$ -spectrum endpoint of molecular tritium  $E_0 = (18\,574.00 \pm 0.07)$  eV [Mye+15; OW08]. Second, there is the final state energy of the molecule  $V_f$ . The excited energy state  $f$  is caused by vibration, rotation or electronic excitation of the decaying molecule. A review on tritium molecular final states and tabulated values can e. g. be found in [BPR15] and references therein. The probability that the molecule is in a final state of energy  $V_f$  after the decay is denoted by  $P_f$ . Then the energy of the neutrino reads

$$\epsilon_f = E_0 - E_S - V_f. \quad (3.6)$$

Third, there is the neutrino's momentum  $\sqrt{\epsilon_f^2 - m_\nu^2}$ . Then, the complete phase space factor of the neutrino is a sum over all possible molecular final states labeled  $f$ .

Lastly, the Heavyside step function  $\Theta$  ensures a positive kinetic energy of the neutrino.

The differential rate is depicted in figure 3.1 for a vanishing and non-vanishing effective electron-antineutrino mass. The difference of the two  $\beta$  spectra forms the foundation for neutrino mass inference at KATRIN.

### 3.2. Response Function

The aim of this chapter is an introduction to the mathematical formalism for the electron rate at the KATRIN detector. The previous section 3.1 gives an expression for the differential  $\beta$ -electron rate. The next step is the inclusion of the characteristics of the KATRIN experimental setup. This can be accomplished by introducing the KATRIN response function. In the outlined formalism, it reflects the probability of an electron emitted in the WGTS to reach the KATRIN detector [Gro15] as a function of experimental settings. Which settings are incorporated in particular is developed throughout this chapter.

First, central concepts, the used nomenclature and an overview of the respected experimental settings are presented in section 3.2.1. Then, the components of the response function are introduced:

- The gas dynamics within the WGTS needs to be simulated. See section 3.2.2.
- The characteristics of the KATRIN spectrometer can be summarized in the transmission function. See section 3.2.3.
- The passage of electrons through the WGTS is influenced by scattering off gas molecules. The probability for such scattering is discussed in section 3.2.4. Furthermore, the amount of energy an electron loses when scattering is considered in section 3.2.5.

Finally, the described components will be assembled to the KATRIN response function in section 3.2.6 and section 3.2.7 summarizes the obtained results.

#### 3.2.1. Concepts and Nomenclature

Before the formalism for the KATRIN response function is developed, this section introduces naming conventions and useful concepts.

##### Coordinate System

This chapter focuses on a one-dimensional description of the KATRIN response function. The position along the beam line is denoted with  $z$ . The origin of the coordinate system is the center of the WGTS as already chosen in previous works, e. g. [Gro15; Kle14]. In this sense, the rear and the front of the WGTS of length  $d$  have the coordinates  $\mp d/2$ .

##### Pitch Angle

In this chapter, the angle between an electron's direction of motion and the magnetic field along the beam line axis, the so-called pitch angle, is denoted by  $\theta$ .

##### Parameter Indices

Whether an electron reaches the KATRIN detector depends i. a. on its starting parameters when originating in the WGTS. In this chapter, these starting parameters are denoted with a lower index  $S$ . The three decisive starting parameters are the following:

1. the starting kinetic energy  $E_S$  as discussed within the description of the differential rate in equation (3.1),
2. the starting position  $z_S$  within the WGTS and

3. the starting pitch angle  $\theta_S$  within the WGTS.

Parameters that denote quantities in the analyzing plane (see section 2.6) are denoted with a lower index A.

### Probabilistic Treatment of the Starting Pitch Angle

It should be noted, that the three listed starting parameters are not known for a single  $\beta$  electron, which suggests a probabilistic treatment. Within the scope of this thesis, this is of importance with respect to the starting pitch angle. Therefore, the concept is explained in the following:

Given the distribution  $\omega(\theta_S)$  of starting pitch angles, the mean value of any function  $g(\theta_S)$  depending on a fixed starting pitch angle  $\theta_S$  can be calculated within an interval  $[0, \theta_{\max}]$  by applying the definition of the mean value

$$\langle g(\theta_S) \rangle = \frac{\int_0^{\theta_{\max}} \omega(\theta_S) g(\theta_S) d\theta_S}{\int_0^{\theta_{\max}} \omega(\theta_S) d\theta_S}. \quad (3.7)$$

An isotropic  $\beta$ -electron emission by a tritium molecule into the unit sphere, meaning all combinations of spherical emission angles ( $\varphi, \vartheta = \theta_S$ ) are equally likely, yields as distribution for the starting pitch angles of [KAT05]

$$\omega(\theta_S) = \sin \theta_S \quad (3.8)$$

with normalization

$$\int_0^{\theta_{\max}} \omega(\theta_S) d\theta_S = \frac{1}{1 - \cos \theta_{\max}}. \quad (3.9)$$

In this chapter,  $\theta_{\max}$  denotes the maximum acceptance angle due to the magnetic bottle effect as explained in section 2.6.1 with a design value of  $\theta_{\max} \approx 51^\circ$  [KAT05]. This calculation of the mean value is applied multiple times throughout this thesis.

### Experimental Settings

As the response function models the characteristics of the KATRIN apparatus, it naturally depends on the experimental settings. The quantities used within this chapter are listed in the following:

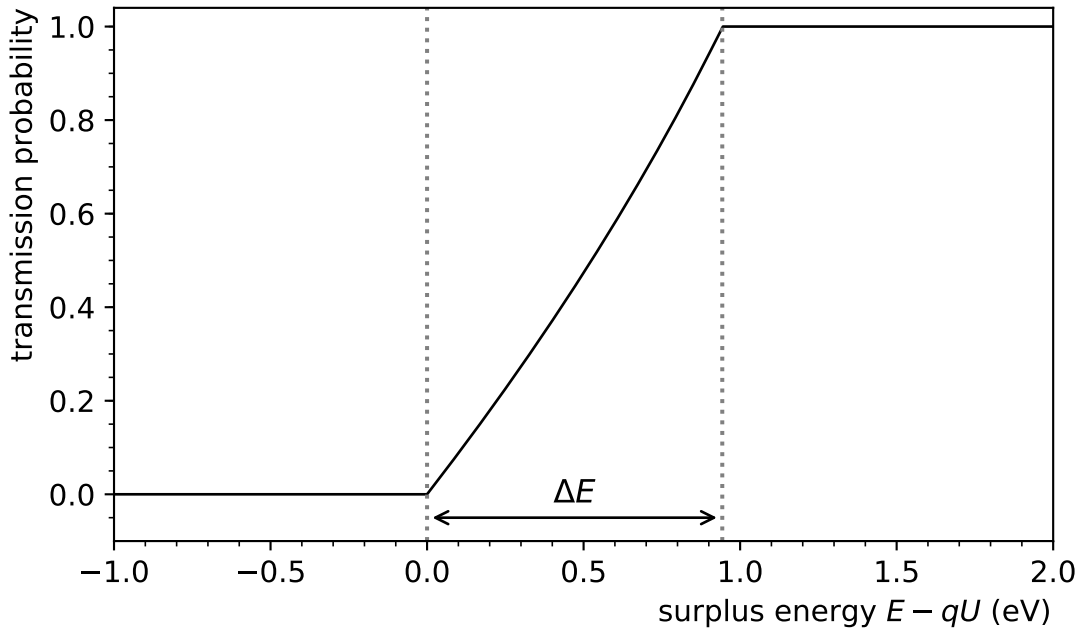
- the magnetic field  $B_S$  at the place of origin of a  $\beta$  electron within the WGTS,
- the magnetic field  $B_A$  within the analyzing plane,
- the maximum magnetic field  $B_{\max}$  along the beam line axis,
- the retarding voltage  $U$  and the retarding energy  $qU$  and
- the starting potential  $U_S$  of a  $\beta$  electron within the WGTS.

For the detailed meaning of these parameters and their KATRIN design values, see section 2.6. It should be noted, that none of these quantities are constant, but that they exhibit a spatial, especially a radial, dependency [KAT05]. For ease of notation, the spatial dependency is left implicit within this chapter. Also, the detector efficiency is treated separately from the response function (see subsequent section 3.4).

#### 3.2.2. Gas Dynamics

The gas dynamics within the STS has to be simulated. This topic is not treated in detail here. The reader is referred to [Höt12; Hei18; Kuc+18; Kuc16]. In short, in a one-dimensional description, the result of such a gas dynamic simulation is the gas molecule density  $\rho(z)$ . For nominal settings, averaging  $\rho(z)$  along the beam line axis and multiplication by the length  $d$  of the WGTS yields the design column density  $\rho d = 5 \times 10^{17} \text{ cm}^{-2}$  [KAT05].





**Figure 3.2.:** The KATRIN transmission function as described by equation (3.13). It denotes the probability for an electron with a kinetic energy  $E$  to pass through the spectrometer set to a retarding energy of  $qU$ . The probabilistic treatment of the starting pitch angles of electrons leads to the MAC-E-filter width  $\Delta E$  with the nominal value of 0.93 eV [KAT05]. (Calculated with [S-SSC].)

### 3.2.3. Transmission Function

The transmission function denotes the probability of an electron to pass the MAC-E filter. It can be characterized by the transmission energy [Gro15]

$$E_{\text{tr}}(qU, E, \theta_S) = \frac{q(U - U_S)}{1 - \sin^2 \theta_S \frac{B_A}{B_S} \frac{\gamma(E)+1}{\gamma_A+1}}. \quad (3.10)$$

where  $\gamma(E)$  and  $\gamma_A$  denote the relativistic Lorentz factor of the  $\beta$  electrons with energy  $E$  and in the analyzing plane. As the electrons are slowed down substantially by the retarding potential in the spectrometer,  $\gamma_A \approx 1$  holds to a good approximation. In the following, for ease of notation, also  $U_S = 0$  and  $\gamma(E) = 1$  is assumed (the form of the following equations allow to easily identify that the correct values would have to be re-substituted into the retarding voltage or in the fraction of the magnetic fields, which justifies this simplified notation).

Electrons pass the MAC-E filter if their energy  $E$  when arriving at the spectrometer surpasses the transmission energy  $E_{\text{tr}}$  (eq. 3.10). This condition can be resolved for the starting pitch angle [Gro15]

$$\begin{aligned} E &> E_{\text{tr}}(qU, E, \theta_S) \\ \Leftrightarrow \theta_S &< \theta_{\text{tr}}(E, qU) := \arcsin \left( \sqrt{\frac{E - qU}{E} \frac{B_A}{B_S}} \right). \end{aligned} \quad (3.11)$$

Using equation (3.11), the transmission function depending on the starting pitch angle and the starting energy of electrons can be formulated as a step function

$$\mathcal{T}(E, qU, \theta_S) = \begin{cases} 1 & \text{if } \theta_S < \theta_{\text{tr}}(E, qU) \\ 0 & \text{otherwise} \end{cases}. \quad (3.12)$$

Calculating the mean value of this step function with respect to the probabilistic distributed starting pitch angles of  $\beta$  electrons as described in section 3.2.1 yields the KATRIN transmission function [KAT05]

$$T(E, qU) = \langle \mathcal{T}(E, qU, \theta_S) \rangle = \begin{cases} 0 & \text{if } E < qU \\ \frac{1 - \sqrt{1 - \frac{E - qU}{E} \frac{B_S}{B_A}}}{1 - \sqrt{1 - \frac{\Delta E}{E} \frac{B_S}{B_A}}} & \text{if } qU < E < qU + \Delta E, \\ 1 & \text{if } qU + \Delta E < E \end{cases} \quad (3.13)$$

where

$$\Delta E = E \cdot B_A / B_{\max} \quad (3.14)$$

is the MAC-E-filter width as explained in section 2.6. The transmission function is depicted in figure 3.2 for the KATRIN design values.

### 3.2.4. Probability for Electron-Scattering within the WGTS

This section derives an expression for the probability  $P_l$  for an electron to scatter  $l$  times in the WGTS.

The electron moves on a spiral track due to its cyclotron motion in the magnetic field in the WGTS. Therefore, when traveling an infinitesimal distance  $dz$  in  $z$ -direction, it travels a total distance of

$$ds = \frac{1}{\cos \theta_S} dz. \quad (3.15)$$

Remarkably, this expression is independent of the electron energy and the magnetic field strength in the WGTS. The effective column density can then be expressed as a line integral along the electron's path  $\varphi$  over the gas density  $\rho(z)$  from the starting position of the electron to the point where it leaves the WGTS

$$\lambda(z_S, \theta_S) = \int_{\varphi} \rho(\vec{r}) ds = \frac{1}{\cos \theta_S} \int_{z_S}^{d/2} \rho(z) dz. \quad (3.16)$$

Then, the expected scattering count is the product of the effective column density  $\lambda(z_S, \theta_S)$  and the scattering cross section  $\sigma$  [Gro15]

$$\mu(z_S, \theta_S) = \lambda(z_S, \theta_S) \sigma. \quad (3.17)$$

Using equation (3.17), the probability for  $l$ -fold scattering can be expressed as a Poisson distribution [Gro15]

$$P_l(z_S, \theta_S) = \frac{\mu(z_S, \theta_S)^l}{l!} e^{-\mu(z_S, \theta_S)}. \quad (3.18)$$

The mean value with respect to the starting positions and the starting pitch angles can be calculated [Gro15]

$$\bar{P}_l = \frac{1}{d} \int_{-d/2}^{d/2} \frac{1}{1 - \cos \theta_{\max}} \int_0^{\theta_{\max}} \sin \theta_S P_l(z_S, \theta_S) d\theta_S dz_S. \quad (3.19)$$

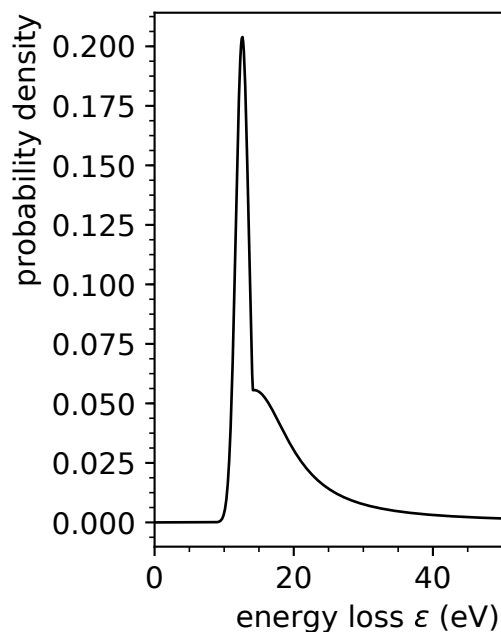
Table 3.1 lists the numerical evaluation of these averaged scattering probabilities. With these probabilities at hand, the next step is the derivation of the energy an electron loses when scattering.

**Table 3.1.: Probability for severalfold electron-scattering in the WGTS.** Listed are the evaluations of equation (3.19) for the following input parameters: A scattering cross section of  $\sigma = 3.456 \times 10^{-22} \text{ m}^2$  [KAT05], a constant gas column density  $\rho d = 5 \times 10^{17} \text{ cm}^{-2}$ , a WGTS beam tube length of  $d = 10.0820 \text{ m}$  and a maximum acceptance angle of  $\theta_{\text{max}} = 50.7685^\circ$ . The same values can be found in [Gro15; Kle14]. The reader is also referred to a review of this values in section 5.2.

| scattering count $l$ | scattering probability<br>according to equation (3.19) |
|----------------------|--|
| 0                    | 41.33 %  |
| 1                    | 29.27 %  |
| 2                    | 16.73 %  |
| 3                    | 7.91 %   |
| 4                    | 3.18 %   |

**Figure 3.3.: The probability density for the energy loss of electrons due to scattering off tritium molecules.** The energy loss function is shown as per equation (3.21) and as determined at the Troitsk experiment [Ase+00]. The table below lists the corresponding parameters, where  $\epsilon_1$  was fixed and  $\epsilon_c$  was chosen to make the piecewise defined function continuous.

| parameter    | value                         |
|--------------|-------------------------------|
| $A_1$        | $0.204 \pm 0.001$             |
| $A_2$        | $0.0556 \pm 0.0003$           |
| $\omega_1$   | $(1.85 \pm 0.02) \text{ eV}$  |
| $\omega_2$   | $(12.5 \pm 0.1) \text{ eV}$   |
| $\epsilon_1$ | 12.6 eV                       |
| $\epsilon_2$ | $(14.30 \pm 0.02) \text{ eV}$ |
| $\epsilon_c$ | 14.09 eV                      |



### 3.2.5. Energy Loss of Electrons due to Scattering

This section describes the “energy loss function”  $f_l(\epsilon)$ . It denotes the probability density for an electron to lose a specific amount of energy  $\epsilon$  when scattering  $l$  times. Only the case of inelastic scattering is treated here. For an additional treatment of elastic scattering, which is less likely by one order of magnitude<sup>1</sup>, the reader is referred to [Kle+19].

The energy loss function for no scattering is the Dirac delta function [Kle+19]

$$f_0(\epsilon) = \delta(\epsilon). \quad (3.20)$$

A phenomenological description for 1-fold scattering of electrons off hydrogen isotopologues

<sup>1</sup>Compare the cross sections  $\sigma_{\text{inel}} = 3.456 \times 10^{-22} \text{ m}^2$  [KAT05] and  $\sigma_{\text{el}} = 0.29 \times 10^{-22} \text{ m}^2$  [Kle+19] for kinetic electron energies of  $\sim 18.6 \text{ keV}$  that are relevant in regard to the KATRIN experiment.

was derived from data at the Troitsk experiment [Ase+00; Abd+17]

$$f_1(\epsilon) = \begin{cases} 0 & \text{if } \epsilon < 0 \\ A_1 \cdot e^{-2\left(\frac{\epsilon-\epsilon_1}{\omega_1}\right)^2} & \text{if } 0 \leq \epsilon < \epsilon_c \\ A_2 \cdot \frac{\omega_2^2}{\omega_2^2 + 4(\epsilon-\epsilon_2)^2} & \text{if } \epsilon \geq \epsilon_c \end{cases} \quad (3.21)$$

Figure 3.3 depicts this energy loss function for scattering off tritium molecules. It should be noted, that a more recent, preliminary energy loss model derived from a designated subgroup of the KATRIN collaboration is investigated in chapter 6.

For severalfold scattering, the above function  $f_1$  has to be convoluted with itself and the energy loss function becomes [Kle+19]

$$f_l(\epsilon) = \bigotimes_{i=0}^l f_1(\epsilon) \quad (3.22)$$

where  $\otimes$  denotes the convolution

$$(f \otimes f)(\epsilon) = \int_{-\infty}^{\infty} f(\epsilon - \epsilon') f(\epsilon') d\epsilon'. \quad (3.23)$$

### 3.2.6. Assembly of the Response Function

This section gives an expression for the KATRIN response function. It should be noted, that the notation chosen here differs slightly from the ones used in the works [Gro15; Kle+19], that this derivation is largely based on. The latter make approximations of the transmission properties and introduce more involved concepts where the approximations do not hold. Here, this approach is inverted: First, the involved concepts are applied and the approximations are introduced in a second step. However, the final results are in agreement.

The KATRIN response function in dependence on the starting position and pitch angle of an electron reads

$$\mathcal{R}(E_S, qU, z_S, \theta_S) = \sum_l \int_{-\infty}^{\infty} \mathcal{T}(E_S - \epsilon, qU, \theta_S) \cdot P_l(z_S, \theta_S) \cdot f_l(\epsilon) d\epsilon \quad (3.24)$$

$$= \sum_l \int_0^{\infty} \mathcal{T}(E_S - \epsilon, qU, \theta_S) \cdot P_l(z_S, \theta_S) \cdot f_l(\epsilon) d\epsilon. \quad (3.25)$$

where the integral goes over the energy losses, the sum goes over the scattering count,  $\mathcal{T}$  denotes the non-averaged transmission function (3.13),  $P_l$  the non-averaged scattering probabilities (3.18) and  $f_l$  the energy loss function (3.22). The cut of the lower integral limit is caused by the vanishing energy loss function ( $f_l(\epsilon) = 0$  if  $\epsilon < 0$ ). In words, the transmission function is smeared using the energy loss function as a smearing kernel and then a weighted sum is formed over generations of  $l$ -fold scattered electrons where the weight is the probability to scatter  $l$  times.

The mean value of equation (3.24) with respect to the starting pitch angle can be calculated as described in section 3.2.1. Also, the corresponding integral is swapped with the integral over the energy loss and the sum over the scattering count

$$R(E_S, qU, z_S) = \langle \mathcal{R}(E_S, qU, z_S, \theta_S) \rangle \\ = \sum_l \int_0^{\infty} \int_0^{\theta_{\max}} \frac{\sin \theta_S \cdot \mathcal{T}(E_S - \epsilon, qU, \theta_S) \cdot P_l(z_S, \theta_S)}{1 - \cos \theta_{\max}} d\theta_S \cdot f_l(\epsilon) d\epsilon. \quad (3.26)$$

This expression can be reformulated to have the same form as the non-averaged response function (3.24). This means, the product “transmission function times scattering probability times energy loss function” can be reestablished, which also reconciles the notation with the expression given in [Gro15]. Therefore, the factor  $1 = \bar{P}_l/\bar{P}_l$  with the averaged scattering probabilities from equation (3.19) is introduced into equation (3.26) and the “detailed transmission function”  $T_l^*$  is defined

$$\begin{aligned} R(E_S, qU, z_S) &= \sum_l \int_0^\infty \underbrace{\int_0^{\theta_{\max}} \frac{\sin \theta_S \cdot \mathcal{T}(E_S - \epsilon, qU, \theta_S) \cdot P_l(z_S, \theta_S)}{(1 - \cos \theta_{\max}) \cdot \bar{P}_l} d\theta_S}_{T_l^*(E_S - \epsilon, qU, z_S)} \cdot \bar{P}_l \cdot f_l(\epsilon) d\epsilon \\ &= \sum_l \int_0^{E_S - qU} T_l^*(E_S - \epsilon, qU, z_S) \cdot \bar{P}_l \cdot f_l(\epsilon) d\epsilon, \end{aligned} \quad (3.27)$$

where the cut on the upper integral limit from  $\infty$  to  $E_S - qU$  is justified below.

The non-averaged transmission function  $\mathcal{T}$  (3.11) within  $T_l^*$  is a step function with respect to the starting pitch angle  $\theta_S$  of an electron. This cuts the upper integral limit from  $\theta_{\max}$  to  $\theta_{\text{tr}}$  when integrating over  $\theta_S$ . Furthermore, in analogy to the KATRIN transmission function from equation (3.13), a distinction of cases avoids imaginary square roots. One obtains the detailed transmission function as given in [Gro15; Kle+19]

$$T_l^*(E, qU, z_S) = \begin{cases} 0 & \text{if } E < qU \\ \int_0^{\theta_{\text{tr}}(E, qU)} \frac{\sin \theta_S \cdot P_l(z_S, \theta_S)}{(1 - \cos \theta_{\max}) \cdot \bar{P}_l} d\theta_S & \text{if } qU < E < qU + \Delta E \\ 1 & \text{if } qU + \Delta E < E \end{cases}, \quad (3.28)$$

where  $\theta_{\text{tr}}$  denotes the transmission-pitch angle (3.11) and  $\Delta E$  the MAC-E-filter width (3.14). As  $T_l^*$  vanishes for  $E < qU$  the upper integral limit over energy losses in the response function (3.27) can be cut to  $E_S - qU$ . Furthermore, it was found, that for  $l > 3$  scatterings, the detailed transmission function  $T_l^*$  can be exchanged for the KATRIN transmission function (3.13) without introducing a significant error [Gro15].

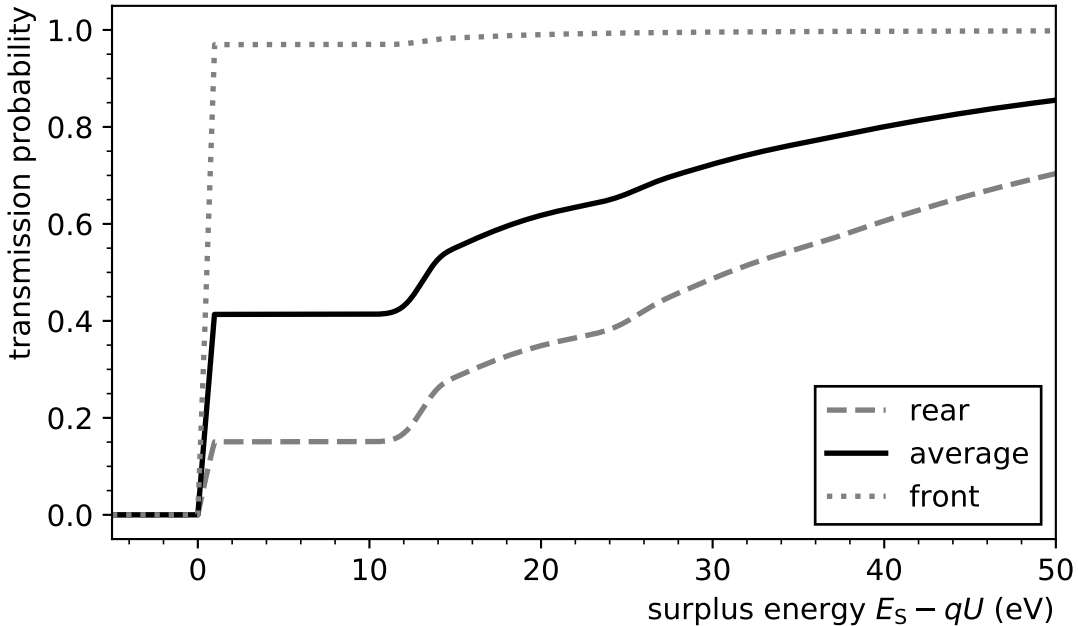
### 3.2.7. Discussion of the Response Function

In equation (3.27), the KATRIN response function was derived, which reconciles with the expressions given in [Gro15; Kle+19]

$$R(E_S, qU, z_S) = \sum_l \int_0^{E_S - qU} T_l^*(E_S - \epsilon, qU, z_S) \cdot \bar{P}_l \cdot f_l(\epsilon) d\epsilon, \quad (3.29)$$

where  $T_l^*$  is the detailed transmission function (eq. 3.28),  $\bar{P}_l$  are the averaged scattering probabilities (eq. 3.19), and  $f_l$  is the energy loss function (eq. 3.22). The response function denotes the mean (with respect to the pitch angle distribution) probability of an electron starting with an energy  $E_S$  at a position  $z_S$  to overcome the retarding energy  $qU$  and reach the detector.

Figure 3.4 shows the response function for two different starting positions of electrons as well as averaged over all starting positions. It exhibits many features: For unscattered electrons, the response function resembles the transmission function. This causes the steep rise at  $E_S - qU = 0$  eV within the interval of the MAC-E-filter width  $\Delta E \approx 0.93$  eV. As the transmission probability is weighted by the probability for no scattering, the plateaus resemble the corresponding probabilities (eq. 3.18 averaged over starting pitch angles):  $\sim 12\%$  (rear),  $41.33\%$  (average, see table 3.1),  $\sim 98\%$  (front). Furthermore, the discontinuity in the first derivative of the energy loss function at  $\epsilon_c = 14.09$  eV causes



**Figure 3.4.: The KATRIN response function** at a retarding energy of  $qU = 18\,545$  V. It is depicted for three cases: for electrons starting  $\sim 9$  mm from the rear and front of the WGTS and averaged over all starting positions. For a description of the manifold features, the reader is referred to the main text. (Calculated with [S-SSC].)

kinks in the response function. As the energy loss function has an onset at  $\epsilon_0 \approx 10$  eV, the corresponding kinks are at  $n \cdot \epsilon_0 + \epsilon_c$  ( $n \in \{0, 1, \dots\}$ ) and are increasingly smoothed for higher  $n$ . (Also see figure 3.3 for the energy loss function.) The response function (3.29) can be understood as a weighted sum of smeared transmission functions approximately shifted by the onset of the energy loss function. Electrons starting from the front of the WGTS are unlikely to scatter, which is why the response function almost resembles the transmission function. Electrons starting from the rear are likely to scatter. Thus, the corresponding response function shows the features of multiple scatterings. For multiple scatterings, the sharp edges of the transmission function are smoothed by the energy loss, which is why only one sharp edge and one plateau, namely for the no-scattering case, is apparent.

### 3.3. Integral Rate

This section gives an expression for the integral  $\beta$ -electron rate at the KATRIN detector.

As already mentioned, the response function (3.29) depends on the starting position of the electrons. To account for this, the WGTS can be thought of being divided into  $n$  slices of width  $w = d/n$  and an averaged response function for the  $j$ th ( $j \in \{0, 1, \dots, n-1\}$ ) slice can be given:

$$R(E_S, qU, z_S) \rightarrow R_j(E_S, qU) = \frac{1}{w} \int_{-d/2+jw}^{-d/2+(j+1)w} R(E_S, qU, z_S) dz_S. \quad (3.30)$$

As can be seen from equations (3.29) and (3.28), this averaging integral can be propagated to the scattering probabilities in the numerator of the detailed transmission function.

The integral rate then reads [Kle+19]

$$\Gamma(qU) = \frac{1}{2} \sum_{j=0}^n N_{j,T} \cdot \int_{qU}^{E_0} \left( \frac{d\Gamma(E_S)}{dE_S} \right) \cdot R_j(E_S, qU) dE_S. \quad (3.31)$$

Here, the integral goes over all starting energies that enable electrons to overcome the retarding potential. The transmission probability vanishes for starting energies smaller than  $qU$  and the differential rate vanishes for energies above the  $\beta$ -spectrum endpoint  $E_0$  (see equation 3.1), which yields the two integral limits. The sum goes over all slices of the WGTS.  $N_{j,T}$  is the number of tritium nuclei in the  $j$ th slice of the WGTS. And the factor  $1/2$  accounts for the fact that, on average, only half the  $\beta$  electrons are emitted towards the detector.

### 3.4. Detector Counts

This section gives an expression for the electron counts measured by the KATRIN detector.

Therefore, the detector efficiency  $\epsilon_{\text{det}} \in [0, 1]$  has to be taken into account (for a description, its determination and value see section 2.7). Furthermore, the background rate  $\Gamma_{\text{bg}}$  (with a nominal value of 10 mcps [KAT05]) has to be considered. Also a relative rate factor  $A_S = 1$  between the background and the  $\beta$ -electron rate is introduced as it can be used in fitting procedures (see section 4.3). Assuming a measurement time of  $t(qU)$  attributed to a retarding energy  $qU$ , the detector counts are [Kle14]

$$N(qU) = t(qU) \cdot \epsilon_{\text{det}} \cdot (A_S \cdot \Gamma(qU) + \Gamma_{\text{bg}}) , \quad (3.32)$$

where  $\Gamma(qU)$  denotes the integral rate (3.31).

### 3.5. Model Amendments

The outlined formalism that lead to the expression for the detector counts (3.32) forms a scaffold for the mathematical formalism that describes a KATRIN measurement. Modifications of isolated terms can incorporate further effects. For a comprehensive list, the reader is referred to [Kle+19]. Selected examples are listed below:

- **Doppler effect:** Gas flow and temperature move the tritium molecules and hence smear the kinetic energy distribution of  $\beta$  electrons (see section 2.2). This can be modeled by convolving the differential rate (3.1) with a Maxwellian distribution or by applying corrections to the final energy states of the decaying molecules.
- **Plasma potential:** Space charges, respectively a plasma, forms within the WGTS due to the tritium decay (see section 2.2).  $\beta$  electrons may originate at higher/lower potentials due to space charges. This can be modeled by an adaption of the starting potential  $U_S$  in the transmission energy (3.10).
- **3-dimensional description:** This chapter focuses on a 1-dimensional formalism. However, as noted, input parameters such as the magnetic fields are not solely  $z$ -dependent, which requires a 3-dimensional approach and an incorporation of the segmentation of the detector. This can be accomplished by calculating the detector counts (3.32) separately for each detector pixel exploiting that the magnetic flux tube maps specific volumes of the WGTS onto specific areas of the analyzing plane and detector pixels.

### 3.6. Measurement Time Distribution

KATRIN measures electron counts as described in equation (3.32) at a set of retarding energies  $\{qU_i\}$ . How much measurement time  $t(qU_i)$  is attributed to a certain retarding energy is specified in a measurement time distribution (MTD). The MTD influences the experiment's sensitivity to the neutrino mass. An optimal MTD balances the following aspects:

1. Some measurement time has to be attributed to retarding energies beyond the endpoint of the integral tritium- $\beta$  spectrum to determine the background rate. The optimal duration depends on the background rate, but can generally take up a sizable fraction (of order 30 percent) of the overall measurement time. [KAT05; Kle14].
2. Near its endpoint, the shape of the integral tritium- $\beta$  spectrum depends most strongly on the neutrino mass. Hence, most measurement time should be attributed to this region [KAT05; Kle14].
3. Retarding voltage bins deeper into the spectrum increase the count rate and hence, lower the statistical uncertainty due to Poisson statistics. These measurements mainly determine the slope of the integral  $\beta$  spectrum [KAT05; Kle14].
4. The theoretical description of the integral tritium- $\beta$  spectrum is optimized for the endpoint region. For example the molecular final states for  $\beta$ -electron energies 40 eV below the endpoint would need further investigation [Dos+06]. Hence, deeper scans introduce modeling uncertainties. However, it is expected that continuous modeling efforts decrease these uncertainties as needed.

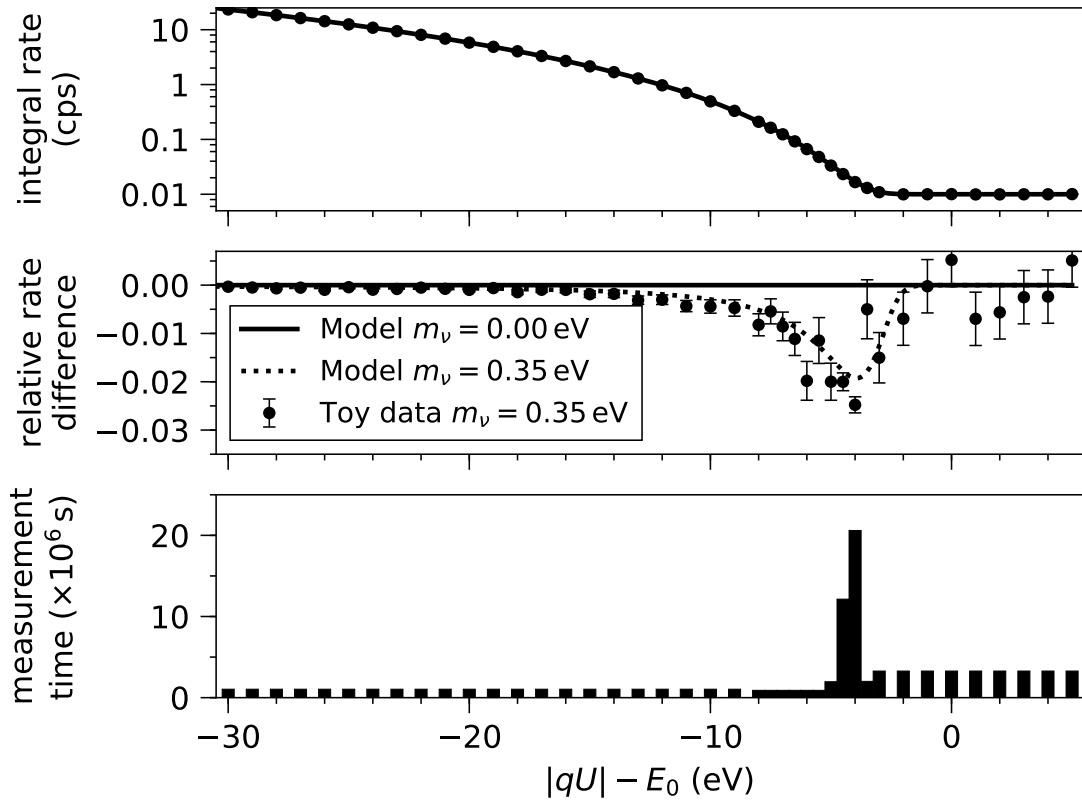
The KATRIN Design Report [KAT05] suggests five three-year-long MTDs for different measurement ranges  $[E_0 - \alpha \text{ eV}, E_0 + 5 \text{ eV}]$  with  $\alpha \in \{20, 25, 30, 40, 50\}$  and comes to the conclusion that  $\alpha = 30$  yields the best sensitivity to the neutrino mass.

As an energy-dependent effect is investigated within this thesis, it should be noted that scans beyond the 50 eV range have already been performed and may also be performed again in the future. For example, searches for sterile neutrinos at the keV-scale would require deeper scans [Mer+19]. On top of that, within several measurement campaigns, deeper scans were conducted: The First Tritium (FT) commissioning campaign successfully proved the apparatus functioning. The corresponding MTD covered a range starting at  $\sim E_0 - 1.6 \text{ keV}$ . The KATRIN neutrino mass measurement campaign 1 (KNM1) is being evaluated during the writing of this thesis. It set out to establish an unprecedented limit on the neutrino mass by  $\beta$ -decay measurements. Its MTD starts at  $\sim E_0 - 90 \text{ eV}$ , but the analysis range for neutrino mass inference remains still to be determined.

### 3.7. A Simulated KATRIN Neutrino Mass Measurement

In summary, a KATRIN measurement yields a set of electron counts  $\{N(qU_i)\}$  (eq. 3.32) distributed over retarding voltage bins  $\{qU_i\}$ , where the counts fluctuate statistically [KAT05]. A possible model for the fluctuations is a Poissonian distribution [Kle14]. Figure 3.5 shows a KATRIN measurement for an MTD starting at  $E_0 - 30 \text{ eV}$  and a total measurement time of three years. The distortion of the measured integral rate by a non-vanishing neutrino mass can be seen approximately 4 eV below the endpoint  $E_0$ . This distortion can be used to infer the squared effective electron antineutrino mass from a KATRIN neutrino mass measurement. Chapter 4 presents corresponding statistical methods.





**Figure 3.5.: Simulated KATRIN neutrino mass measurement for a non-vanishing neutrino mass.** The total measurement time is three years. The top panel shows the measured integral rate  $\Gamma$  in dependence of the retarding energy. The center panel shows the relative rate difference for a non-vanishing neutrino mass  $\Gamma(m_\nu = 0.35 \text{ eV})/\Gamma(m_\nu = 0 \text{ eV}) - 1$ . The difference is  $\sim 2\%$  at a retarding energy approximately 4 eV below the endpoint (simulated as  $E_0 = 18575 \text{ eV}$ ). The bottom panel shows the MTD where most measurement time is attributed to the most sensitive region. This is also reflected by the uncertainty bars of the toy data. (Adapted from [Sei19].)



## 4. Statistical Methods and Neutrino Mass Inference at KATRIN

An estimator for the neutrino mass alongside with an uncertainty or an upper limit will be retrieved by comparing the output of the KATRIN measurement with theoretical predictions within the process of parameter inference. This chapter reviews a selection of statistical approaches suitable in relation to the KATRIN experiment.

Section 4.1 outlines the principle of the maximum likelihood estimator (MLE). Section 4.2 relates the principle of the MLE to a KATRIN measurement and neutrino mass inference. Section 4.3 introduces the formalism of a nominal neutrino mass fit at KATRIN. Section 4.4 reviews the concept of uncertainty intervals and how confidence intervals can be extracted from the likelihood. Section 4.5 introduces the software framework that was used for simulated neutrino mass inference within this thesis. And section 4.6 explains the origin of the often quoted 200 meV (90 % C.L.) KATRIN sensitivity.

### 4.1. Maximum Likelihood Estimation

The likelihood is the probability of a measurement outcome given a hypothesis. A hypothesis depending on a parameter vector  $\boldsymbol{\theta}$  is called a composite hypothesis. A measurement outcome can be quantified by a vector of observed values  $\mathbf{x}$ . The probability  $P$  of  $\mathbf{x}$  given a hypothesis in dependence of  $\boldsymbol{\theta}$  is called the likelihood function [Tan+18]

$$L(\boldsymbol{\theta}) = P(\mathbf{x} | \boldsymbol{\theta}). \quad (4.1)$$

If  $p$  denotes the probability of the independent and identically distributed observed values  $x_i$  in  $\mathbf{x}$ , then the likelihood function can be written as a product [Tan+18]

$$L(\boldsymbol{\theta}) = \prod_i p(x_i | \boldsymbol{\theta}). \quad (4.2)$$

The parameter vector  $\hat{\boldsymbol{\theta}}$  that maximizes the likelihood function is called the maximum likelihood estimator (MLE) for the true values of  $\boldsymbol{\theta}$ .

### 4.2. The Likelihood of a KATRIN Measurement

The MLE-method can be applied to a KATRIN measurement as follows: The data vector is given by a set of  $n$  electron counts  $\{N_{\text{obs},i}\}$  measured at retarding energies  $\{qU_i\}$ . The following hypothesis can be formulated [Kle14]:

$\{N_{\text{obs},i}\}$  follow a Poisson distribution with predicted, expected electron counts  $\{N_{\text{theo},i}(\boldsymbol{\theta})\}$  as per equation (3.32).

The parameters  $\boldsymbol{\theta}$  of this composite hypothesis are discussed in the subsequent section 4.3. For sufficiently high counts ( $\gtrsim 25$  [Kle+19]), the Poisson distribution can be approximated by a Gaussian distribution  $\mathcal{N}(x, \mu, \sigma)$  with mean  $\mu = N_{\text{theo},i}(\boldsymbol{\theta})$  and standard deviation  $\sigma = \sqrt{N_{\text{obs},i}}$ . The likelihood function then reads [Kle14]

$$L(\boldsymbol{\theta}) = \prod_i^n \mathcal{N}\left(x = N_{\text{obs},i}, \mu = N_{\text{theo},i}(\boldsymbol{\theta}), \sigma = \sqrt{N_{\text{obs},i}}\right). \quad (4.3)$$

Commonly, instead of maximizing the likelihood function, its negative logarithm is minimized and a factor 2 is introduced [Tan+18]. This yields

$$-2 \ln L(\boldsymbol{\theta}) = \chi^2(\boldsymbol{\theta}) = \sum_i^n \left( \frac{N_{\text{obs},i} - N_{\text{theo},i}(\boldsymbol{\theta})}{\sqrt{N_{\text{obs},i}}} \right)^2 + \text{constants}. \quad (4.4)$$

The minimization of equation (4.4) yields the MLE estimator  $\hat{\boldsymbol{\theta}}$  for  $\boldsymbol{\theta}$ .

Equation (4.4) is a sum of  $n$  random variables following a standard normal distribution. Hence, evaluated at the MLE, this chi-square expression  $\chi^2(\hat{\boldsymbol{\theta}})$  follows the so-called Pearson's chi-square statistic with  $n - \dim(\boldsymbol{\theta})$  degrees of freedom. Accordingly, the value  $\chi^2(\hat{\boldsymbol{\theta}})$  is a measure for the goodness-of-fit [Tan+18]. In other words,  $\chi^2(\hat{\boldsymbol{\theta}})$  is a measure on how likely the data is under the hypothesis stated at the beginning of this section and can be used to reject it if the data is “not a good fit for the hypothesis”.

In conclusion, equation (4.4) can be used for parameter inference via the maximum likelihood method at KATRIN.

### 4.3. A Nominal KATRIN Neutrino-Mass Fit

With regard to a KATRIN neutrino mass measurement, the parameter of interest in the parameter vector  $\boldsymbol{\theta}$  is the squared neutrino mass  $m_\nu^2$ . Furthermore,  $\boldsymbol{\theta}$  typically comprises: the endpoint of the tritium- $\beta$  spectrum  $E_0$  (eq. 3.6), an overall normalization factor for the  $\beta$ -electron counts  $A_S$  (eq. 3.32) and the background rate  $\Gamma_{\text{bg}}$  (eq. 3.32) [Kle14; KAT05]. Hence, in order to infer the squared neutrino mass, the following four-dimensional likelihood has to be minimized:

$$\chi^2(\boldsymbol{\theta}) = \sum_i^n \left( \frac{N_{\text{obs},i} - N_{\text{theo},i}(m_\nu^2, E_0, A_S, \Gamma_{\text{bg}})}{\sqrt{N_{\text{obs},i}}} \right)^2. \quad (4.5)$$

Applying this procedure with simulated data enables the determination of KATRIN's sensitivity to the neutrino mass (see subsequent section 4.6).

### 4.4. Uncertainty Intervals

The presented maximum likelihood method (section 4.1) provides point estimates  $\hat{\boldsymbol{\theta}}$ . However, additional information can be provided by interval estimates. There are two main approaches to statistical inference, which may be called Bayesian and frequentist [Tan+18]. They differ in their interpretation of probability, which becomes especially evident by the interval estimates associated with the two approaches: Credible and confidence intervals. The following two sections 4.4.1 and 4.4.2 explain the matter in more detail.

#### 4.4.1. Bayesian Credible Intervals

The likelihood  $L(\mathbf{x} | \boldsymbol{\theta})$  is a probability distribution for the data  $\mathbf{x}$  given the parameters  $\boldsymbol{\theta}$ . It can be transformed into a probability density for the parameters  $\boldsymbol{\theta}$  by multiplication with a prior distribution  $\pi(\boldsymbol{\theta})$  and normalization to one using Bayes' theorem. The result is the posterior distribution [Tan+18]

$$P(\boldsymbol{\theta} | \mathbf{x}) = \frac{L(\mathbf{x} | \boldsymbol{\theta}) \cdot \pi(\boldsymbol{\theta})}{\int L(\mathbf{x} | \boldsymbol{\theta}') \cdot \pi(\boldsymbol{\theta}') d\boldsymbol{\theta}'}. \quad (4.6)$$

Bayesian statistics provide no fundamental rule for obtaining the prior probability [Tan+18]. This freedom of choice may reflect subjectivity and thus causes controversy.

The posterior is a multi-dimensional probability distribution. With regard to the KATRIN experiment, there is one parameter of interest, namely, the squared neutrino mass  $m_\nu^2$ . The other dimensions in the parameter space, denoted by  $\boldsymbol{\nu}$ , can be marginalized [Tan+18]

$$p(m_\nu^2 | \mathbf{x}) = \int P(\boldsymbol{\theta} \equiv (m_\nu^2, \boldsymbol{\nu}) | \mathbf{x}) d\boldsymbol{\nu}. \quad (4.7)$$

This yields a one-dimensional probability distribution for the squared neutrino mass, from which a credible interval can be extracted.

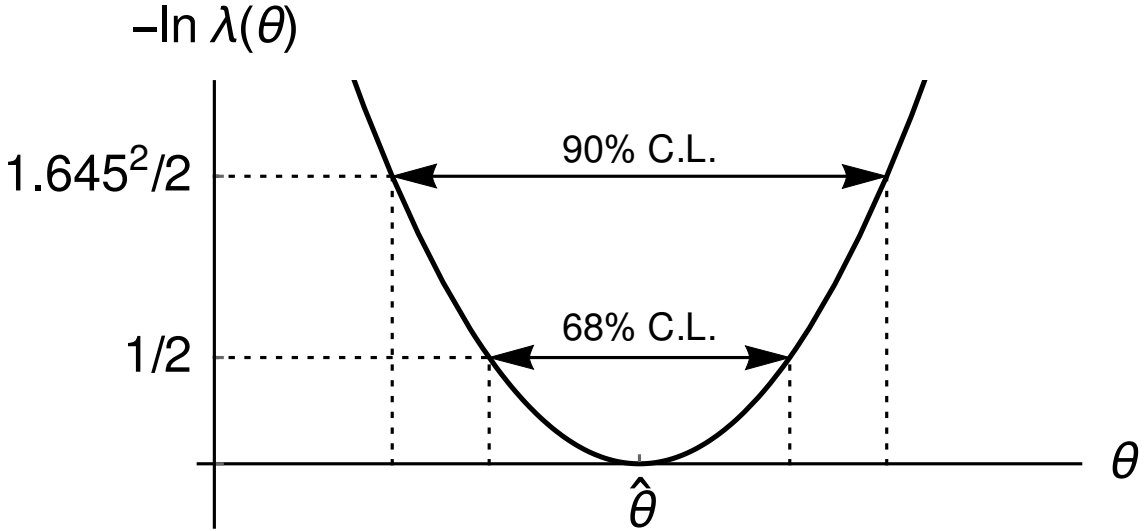
The prior  $\pi(\boldsymbol{\theta} \equiv (m_\nu^2, \boldsymbol{\nu}))$  can be chosen such that it vanishes for  $m_\nu^2 < 0$ . Equation (4.6) shows that then also the posterior vanishes for  $m_\nu^2 < 0$  independently of the likelihood. In other words, the likelihood does not have to be evaluated for negative, unphysical squared neutrino masses.

#### 4.4.2. Frequentist Confidence Intervals

In chapter 6 of this thesis, a sensitivity study is presented that is based on the idea of confidence intervals. Therefore, in this section, the topic is treated in some detail. First, definitions are given for the terms “confidence interval”, “coverage probability” and “confidence level”. Then, it is explained, how confidence intervals can be constructed using the likelihood.

The unqualified phrase “confidence interval” [Tan+18] refers to frequentist intervals obtained with a procedure introduced by Neyman [Ney37] (this Neyman construction is not shown in this thesis, instead another construction is given below, that finds application in chapter 6). Confidence intervals are given with reference to a parameter  $\theta$  that is to be estimated from the data. Within the Neyman construction, the boundary of a confidence interval is given by a function of the data [Tan+18]. This means, the boundary of a confidence interval would fluctuate if one were to repeat the experiment many times (because each time, the obtained data would be different). An ensemble of confidence intervals would be obtained. The coverage probability  $\alpha$  refers to the fraction of confidence intervals in such a hypothetical ensemble that contains the true parameter value of  $\theta$  [Tan+18]. If, in the limit of a large ensemble size, the confidence intervals cover the true parameter value at least a fraction of  $\alpha$  times, the confidence interval obtained from only one data set is understood to have a confidence level of  $\alpha$  [Tan+18]. (An illustration of this abstract definition is developed throughout the subsequent section 6.3 in the context of a sensitivity study. Furthermore, a simulated ensemble of confidence intervals with the corresponding coverage probability is shown in figure 6.3.)

In a practical context, a prescription is required on how to construct confidence intervals of an eligible confidence level. The Neyman construction [Ney37], as mentioned, or the unified approach by Feldman and Cousins [FC98] are such prescriptions. The unified approach is



**Figure 4.1.: Illustration of a confidence interval obtained via the likelihood ratio.** The graph illustrates equation (4.8) on how to extract a confidence interval from a one-dimensional likelihood of Gaussian shape. The negative logarithm of the likelihood ratio (eq. 4.8) is shown, which is zero at the MLE  $\hat{\theta}$  and increases in  $s^2/2$ -steps corresponding to confidence intervals of  $s$ - $\sigma$  C.L. (also refer to the main text).

based on the likelihood ratio. An ordering principle enables the construction of a confidence interval that may either represent an upper limit or and upper as well as a lower limit. How a confidence interval can be obtained via the likelihood ratio is shown below. However, the construction shown below yields a so-called classical confidence interval, that always has an upper and a lower limit and is centrally located around the MLE.

### Construction of a Classical Confidence Interval using the Likelihood Ratio

In this paragraph, the construction of a confidence region is discussed, which is the multidimensional generalization of a confidence interval for the case that the likelihood depends on more than one parameter (which is the case for the KATRIN likelihood). A method to construct a confidence region is to consider a test of the hypothesis that the parameter values  $\theta$  have the true values  $\theta_T$  [Tan+18]. All those parameter values  $\theta$  that are rejected by a test of size  $\alpha$  are excluded from the confidence region of confidence level  $\alpha$  [Tan+18]. A possible test statistic for  $\theta$  is the likelihood ratio of the likelihood at the MLE  $\hat{\theta}$  and  $\theta$  [Tan+18]

$$\lambda(\theta) = \frac{L(\theta)}{L(\hat{\theta})}. \quad (4.8)$$

If the likelihood follows the form of a multivariate Gaussian distribution in  $\theta$ , then the above test statistic (4.8) can be evaluated and the hyper surface defined by

$$-\ln \lambda(\theta) = \frac{s^2}{2} \quad (4.9)$$

approximately encloses a  $s$ - $\sigma$  confidence region for  $\theta$  [Tan+18]. Here,  $s$ - $\sigma$  denotes the corresponding quantile of a Gaussian distribution (for example,  $s \approx 1$  for 68 % C.L. or  $s \approx 1.645$  for 90 % C.L.).

Equation (4.9) is illustrated in figure 4.1 for the case that  $\theta$  is one-dimensional. However, in a realistic scenario,  $\theta$  comprises more than one dimension (at least four in a nominal KATRIN neutrino mass fit, see previous section 4.3). Extracting a confidence interval only for the

squared neutrino mass requires further steps. The introduction of a formalism (profile-likelihood method) that handles the additional dimensions of the nuisance parameters is postponed to the subsequent section 6.3.2 where it is presented in the context of a practical example.

### Negative Squared Neutrino Masses

It should be noted, that the extraction of confidence intervals from the KATRIN likelihood requires its extrapolation to nonphysical, negative squared neutrino masses - a complication that can be avoided when using Bayesian methods as described in the previous section 4.4.1. In this thesis, the strategy proposed by [Wei+93] (and implemented in the scope of [Kle14]) is used and the formula for the phase-space factor of the neutrino in the differential  $\beta$  decay rate (3.1) is adapted for negative squared neutrino masses

$$\sum_f P_f \cdot \epsilon_f \cdot \sqrt{\epsilon_f^2 - m_\nu^2} \cdot \Theta(\epsilon_f - m_\nu) \rightarrow \sum_f P_f \cdot (\epsilon_f + m'_\nu \cdot e^{-(\frac{\epsilon_f}{m'_\nu} + 1)}) \cdot \sqrt{\epsilon_f^2 - m_\nu^2} \quad (4.10)$$

with  $m'_\nu = \mu \cdot \sqrt{-m_\nu^2}$ .

The factor  $\mu$  can be adapted in order to provide an approximately symmetric log-likelihood function around  $m_\nu^2 = 0$ . In this thesis,  $\mu = 0.717$  was used. An approximate symmetry was achieved (see for example table 6.1 and figure 6.2). In order to achieve a perfect symmetry, the factor might need adaption depending on the actual study and the model of the detector counts as per equation (3.32) (for example, [Wei+93] suggests  $\mu = 0.76$ ).

In conclusion, in this section, an outline of confidence intervals has been given. It has also been shown how a confidence region can be constructed using the KATRIN likelihood.

### 4.5. The KaFit and SSC Software Frameworks

With respect to neutrino mass inference at KATRIN, two components have been presented: a model for a KATRIN neutrino mass measurement in chapter 3 and a statistical framework for parameter inference within the current chapter 4.

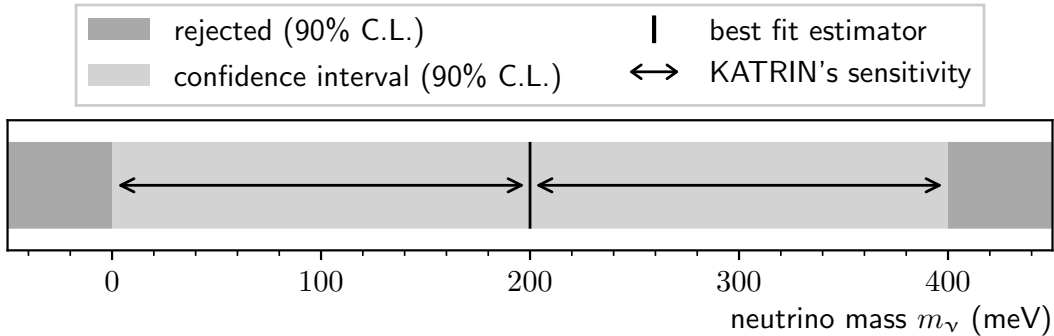
The two components are implemented within two modules of the KATRIN Analysis and Simulations Package (KASPER) [S-KAS]:

1. The **source and spectrum calculation (SSC)** [S-SSC] module implements the formulas as per chapter 3. Additionally, it also includes aspects beyond the given description, for example the gas dynamics within the WGTS. It is based on several previous works, such as [Höt12; Gro15; Kle14; Käf12; Hei18; Kuc16].
2. The **KATRIN Fitter (KaFit)** [S-KF] module translates the  $\beta$  spectrum calculated by SSC into predicted detector counts [Kle14]. Furthermore, KaFit implements several statistic tools tailored to the KATRIN experiment, one of which is the extraction of confidence intervals according to the profile-likelihood method (see subsequent section 6.3.2), which is of importance in the scope of this thesis. The actual minimization and profiling are done by the interfaced MINUIT2 [S-MIN] and MINOS package from the ROOT [Ant+09] analysis framework.

Both modules were used and extended to allow for the analysis done in the scope of this thesis as is explained in the subsequent chapters 5 and 6.

### 4.6. KATRIN's Sensitivity to the Neutrino Mass

This section explains the origin of the often quoted KATRIN design sensitivity to the neutrino mass of 200 meV (90% C.L.). First, a definition of the sensitivity is given in section 4.6.1. Subsequently, section 4.6.2 and 4.6.3 show how this definition was applied by different works to deduce KATRIN's sensitivity.



**Figure 4.2.: Illustration of KATRIN's sensitivity to the neutrino mass.** The graph illustrates KATRIN's sensitivity of 200 meV [KAT05] as per equation (4.11). It shows a hypothetical measurement of the neutrino mass that yields a symmetric confidence interval (90 % C.L.) centrally located around the best fit estimator. If such a classical confidence interval is constructed, KATRIN can reject the null hypothesis of a vanishing neutrino mass at 90 % C.L. if it estimates a neutrino mass of at least 200 meV.

#### 4.6.1. Definition and Construction of KATRIN's Sensitivity

KATRIN's sensitivity can be understood as half the width of a symmetric and central confidence interval (90 % C.L.) for the neutrino mass obtained from a KATRIN neutrino mass measurement. As such, it can be constructed from a  $1\text{-}\sigma$  uncertainty  $\sigma_{\text{tot}}(m_\nu^2)$  on the squared neutrino mass [KAT05]

$$S_{m_\nu}(90\%) = \sqrt{1.645 \cdot \sigma_{\text{tot}}(m_\nu^2)}, \quad (4.11)$$

where the factor 1.645 translates the 68.3 % interval into a 90 % interval in the case of Gaussian uncertainties.

In other words, KATRIN's sensitivity to the neutrino mass can be understood as the minimal neutrino mass that has to be inferred from a KATRIN neutrino mass measurement to exclude the null hypothesis of a vanishing neutrino mass [Kle14] when constructing a classical confidence interval (90 % C.L.). Figure 4.2 illustrates this statement.

For a more comprehensive picture, where not only a classical interval is considered, but also the unified approach according to Feldman and Cousins as well as Bayesian statistics, the reader is referred to [Kle+19]. These techniques also allow to interpret KATRIN's sensitivity as an upper limit that can be established for the neutrino mass if a vanishing neutrino mass is measured.

#### 4.6.2. Sensitivity from Ensemble Tests

In the KATRIN Design Report, the sensitivity to the neutrino mass was evaluated using ensemble tests. An ensemble of many KATRIN measurements was simulated (see section 3.7 on how such a simulation can be conducted) with a true neutrino mass of 0 eV. From each simulated measurement, the squared neutrino mass was inferred in a standard KATRIN four-parameter fit (see section 4.3). The central  $1\text{-}\sigma$  interval of the obtained ensemble of squared neutrino masses was taken as the statistical uncertainty on the squared neutrino mass [KAT05]

$$\sigma_{\text{stat}}^{\text{TDR}}(m_\nu^2) = 0.018 \text{ eV}^2. \quad (4.12)$$

Five main contributions to the overall systematic uncertainty were identified and estimated individually, which yielded an overall systematic uncertainty of  $0.01 \text{ eV}^2$ . Due to the



**Table 4.1.: KATRIN’s sensitivity to the neutrino mass from ensemble tests.** The table lists KATRIN’s sensitivity  $S_{m_\nu}(90\%)$  as defined by equation (4.11). Several works reevaluated the statistical uncertainty according to experimental and theoretical progress. For each reevaluation, a systematic uncertainty of  $\sigma_{\text{sys}}^{\text{TDR}}(m_\nu^2) = 0.017 \text{ eV}^2$  was assumed. A value derived from an ensemble test is a random variable. Corresponding widths of the obtained distributions are reprinted if they were originally stated.

| $\sigma_{\text{stat}}(m_\nu^2)$ (eV <sup>2</sup> ) | $S_{m_\nu}(90\%)$ (meV) | comment                                       | reference |
|--|-------------------------|---|-----------|
| 0.018  | 200                     | design value                                  | [KAT05]   |
| $0.0165 \pm 0.0001$                                | 198                     | updated $\beta$ -spectrum calculation         | [Höt12]   |
| $0.0162 \pm 0.0001$                                | 197                     | further updated $\beta$ -spectrum calculation | [Kle14]   |
| 0.01490  | 193                     | optimized MTD, see section 4.6.3              | [Kle14]   |

early stage of the experiment the systematic uncertainty was conservatively enlarged to a systematic budget at approximately the same scale as the statistical uncertainty [KAT05]

$$\sigma_{\text{sys}}^{\text{TDR}}(m_\nu^2) = 0.017 \text{ eV}^2. \quad (4.13)$$

Adding the statistic and systematic uncertainty quadratically and applying the definition (4.11) yields KATRIN’s design sensitivity [KAT05]

$$S_{m_\nu}^{\text{TDR}}(90\%) = \sqrt{1.645 \cdot \sqrt{\sigma_{\text{stat}}^{\text{TDR}}(m_\nu^2)^2 + \sigma_{\text{sys}}^{\text{TDR}}(m_\nu^2)^2}} \approx 200 \text{ meV}. \quad (4.14)$$

The corresponding investigations were redone in the scope of several works. Table 4.1 lists selected results.

### 4.6.3. Sensitivity using the Profile-Likelihood Method

The approach shown above via ensemble testing yields an ensemble respectively a distribution of estimated squared neutrino masses. The sensitivity is then derived from this ensemble. Instead of using an ensemble of inferred squared neutrino masses, also an ensemble of confidence intervals may be used. This approach is explained in chapter 6. In order to deduce a confidence interval from a single measurement, the profile-likelihood method (see subsequent section 6.3.2) may be used, which finds application in the scope of this thesis.

For comparison with the results from ensemble tests in table 4.1, previous results from [Kle14] based on the profile-likelihood method are shortly reviewed here: Two uncertainties were obtained for two different MTDs in a nominal KATRIN four-parameter fit (see section 4.3). The first MTD had been especially optimized with regard to KATRIN’s sensitivity. The optimization resulted in enlarged measurement times in four isolated ranges of retarding voltage bins. These bins offer enhanced leverage on the four fit parameters of a nominal KATRIN neutrino mass fit and their correlations (for details, the reader is referred to [Kle14]). The obtained statistical uncertainty was  $\sigma_{\text{stat}}(m_\nu^2) = 0.01494 \text{ eV}^2$ . For the second result, the nominal MTD from the KATRIN Design Report (see section 3.6) was used. The corresponding profile likelihood was plotted for an Asimov data set (see section 6.3.3) and  $\sigma_{\text{stat}}(m_\nu^2)$  can be extracted to be between 0.0155 and 0.0165 eV<sup>2</sup>. Both results are in agreement with the results from ensemble tests (see [Kle14] in table 4.1).



# 5. Energy-Dependence of the Cross Section for Inelastic Electron-Scattering within the WGTS

The probability for an electron to scatter when traveling through the WGTS can be characterized by the total scattering cross section  $\sigma_{\text{tot}}$ . Two types of scatterings can be distinguished: elastic and inelastic scattering. The cross section for elastic scattering is smaller than the one for inelastic scattering by one order of magnitude [Kle+19] (also see 3.2.5 for a comparison of the cross section values). This chapter focuses on inelastic scattering and neglects elastic scattering. Within this chapter, the cross section for electrons scattering inelastically off gas molecules is just denoted as “cross section” and with the symbol  $\sigma$ . For ease of notation and reading, the adjective “inelastic” and an index such as “inel” is omitted where the context allows it unambiguously.

The cross section depends on the energy of the incident electrons:  $\sigma \equiv \sigma(E)$ . This dependence has been neglected in the formal modeling of a KATRIN measurement described in chapter 3. This chapter investigates effects related to the incorporation of the energy-dependence. Section 5.1 lists cross section values and formulae from different sources and relates them to each other. Section 5.2 extends the mathematical formalism for a KATRIN measurement in order to incorporate the energy-dependence of the scattering cross section. However, an approximation is made. It is neglected, that an electron that scatters loses energy and becomes more likely to scatter again. This issue is addressed in section 5.4. Section 5.3 discusses the energy-dependence of the scattering cross section within the context of neutrino mass inference using the approximated model. And section 5.5 concludes and offers an outlook.

## 5.1. Cross Section Values for Inelastic Electron-Scattering off Molecular Hydrogen Isotopologues

The goal of this thesis is to investigate to what extent it is important, that the cross section value is not constant but varies with the energy of the incident electrons. In that regard, a formula by [Liu87] is used, that is derived from first principles and denotes the energy-dependent cross section for electrons scattering off hydrogen molecules. This was found to be a feasible approach for the stated goal - at least for first investigations. In section 5.1.1, this formula by [Liu87] is reviewed. For electrons with an energy of  $E_S = 18.6 \text{ keV}$ , its application yields

$$\sigma(18.6 \text{ keV}) = 3.667 \times 10^{-22} \text{ m}^2. \quad (5.1)$$

The formula is in agreement on at least the  $10^{-1}$  level with recent, preliminary estimations obtained through data taken at the KATRIN experiment for electrons scattering off tritium molecules. This hints at its applicability, which is of importance in the light of the following paragraph.

The situation regarding cross section values is not without a certain intricacy. The following list gives an introduction to the matter that shall serve as guidance for the reader:

1. In this work, a theoretical cross section for electrons scattering off hydrogen molecules is used. During a KATRIN neutrino mass measurement, the WGTS is envisaged to be filled with gas of 95 % tritium purity [KAT05]. In that regard, the scattering off tritium molecules is of importance with respect to the KATRIN experiment. How the formulas for scattering off hydrogen molecules can be transformed to tritium molecules is under investigation at the time of writing this thesis. Once, this process is completed, the study presented in this thesis might have to be repeated.
2. The cross section for electrons with an energy of 18.6 keV scattering off tritium molecules was measured at the neutrino mass experiment in Troitsk to be [Ase+00]

$$\sigma(18.6 \text{ eV}) = (3.40 \pm 0.07) \times 10^{-22} \text{ m}^2. \quad (5.2)$$

This value differs by approximately 8 % and 4 standard deviations from the theoretical value given above in equation (5.1).

3. The KATRIN Design Report lists a reference value [KAT05]

$$\sigma_{\text{TDR}} = 3.456 \times 10^{-22} \text{ m}^2. \quad (5.3)$$

This value also differs significantly from the theoretical value given above in equation (5.1). To further complicate the matter, comparability with former results is of importance in the scope of this thesis and several former works are based on  $\sigma_{\text{TDR}}$ . Section 5.1.2 explains why the value of  $\sigma_{\text{TDR}}$  might be erroneous and how this issue is addressed in this work.

### 5.1.1. Theoretical Cross Section Formulae

An expression for the inelastic cross section for electrons scattering off hydrogen molecules can be found in [Liu73]. Two expressions are given: one for relativistic incident electrons and one for non-relativistic incident electrons. In regard to KATRIN, the energies of  $\beta$  electrons from tritium  $\beta$  decay are relevant. The maximum relativistic  $\beta$  factor of electrons from tritium  $\beta$  decay is

$$\beta(E, m) = \sqrt{1 - \frac{1}{\left(\frac{E}{m} + 1\right)^2}} \quad (5.4)$$

$$\Rightarrow \beta_{\text{max,T}} = \beta(E \approx 18.6 \text{ keV}, m_e \approx 511 \text{ keV}) \approx 0.26. \quad (5.5)$$

Traveling at approximately a forth of the speed of light, the  $\beta$  electrons are assumed to behave non-relativisticly. Then, the given expression for the energy-dependent cross section is [Liu73]

$$\sigma(E) = (4\pi a_0^2) \cdot \left(\frac{T(E)}{R}\right)^{-1} \cdot \left[C_1 \cdot \ln\left(\frac{T(E)}{R}\right) + C_2\right] \quad (5.6)$$

with the Bohr radius<sup>1</sup>  $a_0$ , the Rydberg energy<sup>2</sup>  $R$  and two constants  $C_1$  and  $C_2$ . The later two depend on the hydrogen isotopologue. Different values are stated in different works for scattering off hydrogen molecules

$$C_1 = 1.5487 \quad [\text{Liu73}], \quad (5.7a)$$

$$C_2 = 2.2212 \pm 0.0434 \quad [\text{Liu73}], \quad (5.7b)$$

$$C_2 = 1.53 \quad [\text{Ger75}], \quad (5.7c)$$

$$C_2 = 2.4036 \quad [\text{Liu87}]. \quad (5.7d)$$

<sup>1</sup>Bohr radius  $a_0 = 0.52917721067(12) \times 10^{-10} \text{ m}$  [Tan+18]

<sup>2</sup>Rydberg energy  $R = 13.605693009(84) \text{ eV}$  [Tan+18]

The latest of these references, [Liu87], acknowledges that the listed values for  $C_2$  are not compatible.

Furthermore, in equation (5.6),  $T$  denotes the classical kinetic energy<sup>3</sup> using the velocity  $v$  and the electron rest mass  $m_e$

$$\begin{aligned} T &= \frac{1}{2}m_e v^2 = \frac{1}{2}m_e c^2 \frac{v^2}{c^2} \\ &\equiv T(E) = \frac{1}{2}m_e c^2 \beta(E, m_e)^2 \end{aligned} \quad (5.8)$$

with  $\beta(E, m_e)$  as in equation (5.4)<sup>4</sup>. This formula is valid in the center-of-mass frame of the colliding system of the electron and the hydrogen molecule. As the molecule is significantly heavier than an electron, the center of mass-frame was assumed to be the rest frame of the molecule. Also, relative movements due to the gas flow and thermal motion in the WGTS were neglected.

In the scope of this thesis, equation (5.6) with  $C_2$  from [Liu87] (as it is the most up-to-date of the listed ones) is used to include the energy-dependence of the cross section into the mathematical formalism of a KATRIN neutrino mass measurement.

In [Liu73], also a formula for incident electrons with relativistic energies is given. As already mentioned, in this thesis, the  $\beta$  electrons from tritium decay are assumed to behave non-relativistically and the formula is only reprinted here for completeness

$$\sigma(E) = (4\pi a_0^2) \cdot \left(\frac{T(E)}{R}\right)^{-1} \cdot \left[1.5487 \cdot \ln\left(\frac{\beta(E, m_e)}{1 - \beta(E, m_e)^2}\right) + 17.4615\right]. \quad (5.9)$$

The quantities  $T(E)$ ,  $R$  and  $a_0$  match the once in equation (5.6) and  $\beta(E, m_e)$  follows equation (5.4). For energies below 18.6 keV the difference of the formula for relativistic (eq. 5.6) and non-relativistic (eq. 5.9) incident electrons is less than 1%. To what extent relativistic effects in the inelastic scattering process could be of relevance for the KATRIN experiment is not covered in this thesis.

Figure 5.1 shows the theoretical cross-section formulae (5.6) and (5.9) along with the measured value by the Troitsk experiment and the reference value from the KATRIN Design Report.

### 5.1.2. Relation to Cross Section Values Used in Former Works

As already stated and depicted in figure 5.1, the reference cross section from the KATRIN Design Report does not match the theoretical calculations used in this thesis. Nonetheless, several former works (see subsequent section 5.3) are based on the reference value. Comparability is of importance in the scope of this thesis (as will become apparent in section 5.3). How this issue is addressed is explained in the following.

The cross section value stated in the KATRIN Design Report can be recovered from equation (5.6) if an erroneous energy interpretation is applied. If, instead of equation (5.8), one applies the energy interpretation

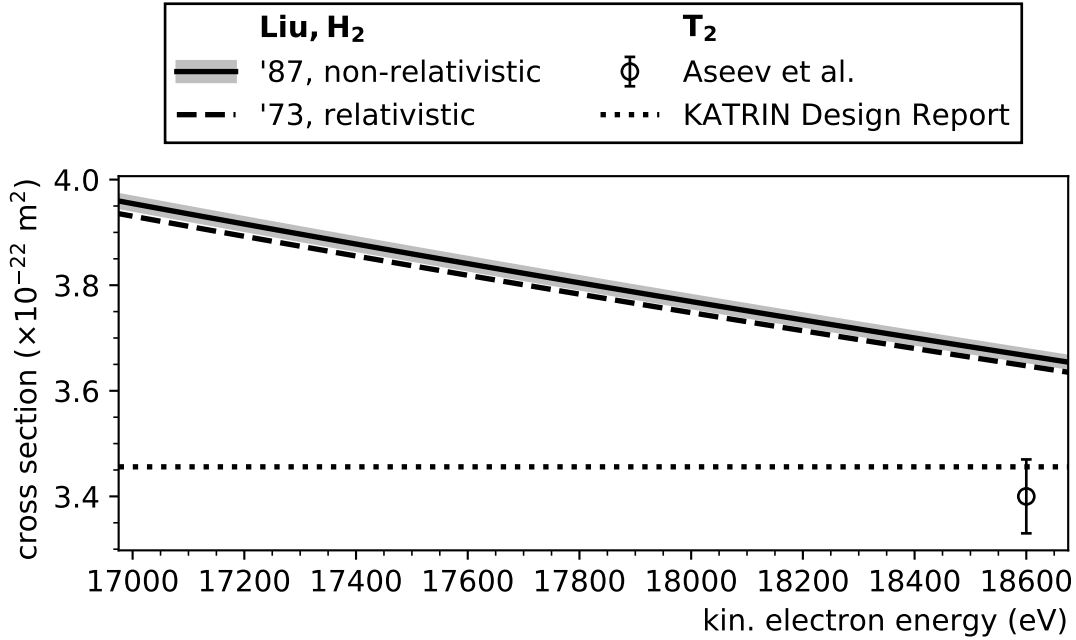
$$T(E) = E, \quad (5.10)$$

the obtained cross section via equation (5.6) is

$$\sigma_{\text{TDR}} \equiv \sigma(E_{\text{TDR}} = 18\,565\text{ eV}) = 3.4559 \times 10^{-22} \text{ m}^2 \approx 3.456 \times 10^{-22} \text{ m}^2 \quad (5.11)$$

<sup>3</sup>I would like to thank Dr. F. Glück for pointing this out. Also see [Ino71] for notations with regard to Bethe theory in scattering processes.

<sup>4</sup>This thesis uses natural units. However, for clarity, it makes sense to explicitly state the speed of light  $c$  in this particular case.



**Figure 5.1.: Inelastic cross section for electrons scattering off molecular hydrogen isotopologues.** The continuous line shows equation (5.6) for incident electrons with non-relativistic energies with constants from equation (5.7a) and (5.7d) where the later is assumed to have an uncertainty according to equation (5.7b). The dashed line shows equation (5.9) for incident electrons with relativistic energies. Also, the measurement by [Ase+00] at the Troitsk neutrino mass experiment and the value stated in the KATRIN Design Report [KAT05] are depicted. The shown energy interval is chosen according to the MTD of the FT measurement campaign.

as stated in the KATRIN Design Report, where the energy  $E_{\text{TDR}}$  is approximately central in the KATRIN design analysis interval (see section 3.6). Whether this had been the approach, that had led to the reference cross section in the KATRIN Design Report is not known.

This work applies the energy interpretation (5.10) when comparability to former works, that used the reference cross section from the KATRIN Design Report, is of importance. Otherwise, equation (5.8) is used. Corresponding indications are given on a case to case basis. The quantitative difference of these two approaches can be assessed by expanding the  $\beta$ -factor (5.4) in the ratio  $E/m_e \approx 18.575/511 \approx 0.036 \ll 1$

$$\beta^2 \approx 2 \frac{E}{m_e} - 3 \left( \frac{E}{m_e} \right)^2. \quad (5.12)$$

The energy interpretation of equation (5.8) then becomes

$$T(E) \approx 0.95 \cdot E, \quad (5.13)$$

which is a shift in energy and hence, in first order, also in the cross section of about 5% compared to the interpretation in equation (5.10). Exact calculations are given in table 5.1.

## 5.2. An Energy-Dependent Scattering Model using the Poisson Distribution

The energy-dependence of the cross section enters into the calculation of the scattering probabilities (3.18). In the derivation that is given in the previous section 3.2.4 the

**Table 5.1.: Probability for severalfold electron-scattering in the WGTS.** The probabilities are averaged over all starting positions and starting pitch angles. Both, the values from a Monte Carlo (MC) particle tracking simulation and the values according to equation (3.19) are given. The cross section was evaluated at an energy of  $E_{\text{TDR}} = 18\,564.374\,63$  eV for the two energy interpretations described by equation (5.10) and (5.8). Further input parameters to the calculations are a constant gas column density  $\rho d = 5 \times 10^{17}$  cm<sup>-2</sup>, a WGTS length of  $d = 10.0820$  m and a maximum acceptance angle of  $\theta_{\text{max}} = 50.7685^\circ$ . The values are given with the precision needed to reproduce the results in the table below in all digits, which is the precision used in SSC. The energy is chosen such, that  $\sigma_{\text{TDR}} = 3.456\,000 \times 10^{-22}$  m<sup>2</sup> is recovered in more than 3 digits as per equation (5.11).

| cross section ( $10^{-22}$ m <sup>2</sup> ) → | 3.456                           | 3.456                        | 3.673      |
|---|---------------------------------|------------------------------|------------|
| source →                                      | MC particle tracking<br>[Gro15] | eq. (3.19)<br>[Gro15; Kle14] | eq. (3.19) |
| scattering count ↓                            |                                 |                              |            |
| 0   | 0.415 ± 0.002                   | 0.41334                      | 0.39564    |
| 1   | 0.292 ± 0.002                   | 0.29266                      | 0.28967    |
| 2   | 0.166 ± 0.001                   | 0.16733                      | 0.17298    |
| 3   | 0.079 ± 0.001                   | 0.07913                      | 0.08590    |
| 4   | 0.031 ± 0.001                   | 0.03178                      | 0.03634    |

dependence on the starting energy  $E_S$  of electrons is neglected. Instead, an average starting energy and hence an average scattering cross section

$$\sigma_{\text{TDR}}(E_{\text{TDR}}) = 3.456 \times 10^{-22} \text{ m}^2 \quad (5.14)$$

(energy interpretation as per equation 5.10) is assumed. Table 5.1 lists the corresponding scattering probabilities averaged over all starting positions and pitch angles of electrons. Additionally, the results of a Monte Carlo particle tracking simulation by [Gro15] and the values using the energy interpretation of equation (5.8) are given. How, instead of assuming energy-independent scattering probabilities, the energy-dependence can be modeled is shown in the following.

### 5.2.1. Formalism of the Poisson Model for Electron-Scattering

An expression for the probability of  $l$ -fold scattering of electrons within the WGTS is derived in the previous section 3.2.4. The given model is independent of the energy of the electrons. Instead of using a constant cross section, the energy-dependence can be respected. The corresponding formulae from section 3.2.4 are repeated below with the energy-dependence made explicit and by using an energy-dependent cross section as per equation (5.6)

$$\mu(E_S, z_S, \theta_S) = \frac{\sigma(E_S)}{\cos \theta_S} \int_{z_S}^{d/2} \rho(z) dz, \quad (5.15a)$$

$$P_l(E_S, z_S, \theta_S) = \text{Poisson}(\mu(E_S, z_S, \theta_S), l), \quad (5.15b)$$

$$\bar{P}_l(E_S) = \frac{1}{d \cdot (1 - \cos(\theta_{\text{max}}))} \int_{-d/2}^{d/2} \int_0^{\theta_{\text{max}}} \sin(\theta_S) \text{Poisson}(\mu(E_S, z_S, \theta_S), l) d\theta_S dz_S. \quad (5.15c)$$

As a reminder,  $\bar{P}_l(E_S)$  in equation (5.15c) denotes the probability for  $l$ -fold scattering of an electron with a starting kinetic energy  $E_S$  averaged over all starting positions and pitch angles in the WGTS. In the following, this model is denoted ‘‘Poisson model’’.

### 5.2.2. Properties of the Poisson Model for Electron-Scattering

Figure 5.2 shows the Poisson model as per equation (5.15). Table 5.1 lists the scattering probabilities for an energy-independent Poisson model (see section 3.2.4) and a reference cross section  $\sigma(E \approx 18\,564\text{ eV}) = 3.673 \times 10^{-22}\text{ m}$ . The energy-dependent Poisson model recovers the energy-independent model exactly at the corresponding energy as expected.

#### Trend of the Energy-Dependence

As depicted in figure 5.2, for an increasing starting energy of electrons within the WGTS, the probability for no and one-fold scattering also increases, while the probability for two-fold scattering decreases. This paragraph aims to give an intuitive argument for this change of sign in the derivative  $d\bar{P}_l(E_S)/dE_S$  in the transition from  $l = 1$  to  $l = 2$ . The Poisson model  $\bar{P}_l(E_S)$  is a probability density in  $l$  for a fixed  $E_S$  (see appendix A for a proof). Hence, the expected scattering count for an energy within the range of starting energies  $E_S \in [17\text{ keV}, 18.6\text{ keV}]$ , that is depicted in figure 5.2, can be calculated (numerically)

$$\begin{aligned}\bar{l}(17.0\text{ keV}) &= \sum_l^{\infty} \bar{P}_l(17.0\text{ keV}) \cdot l \approx 1.23 \\ \bar{l}(18.6\text{ keV}) &= \sum_l^{\infty} \bar{P}_l(18.6\text{ keV}) \cdot l \approx 1.14.\end{aligned}\tag{5.16}$$

In other words, electrons with a starting energy between 17 keV and 18.6 keV are expected to scatter between 1 and 2 times on their way through the WGTS (averaged over all starting positions and pitch angles). Then, the illustrative line of argument is the following:

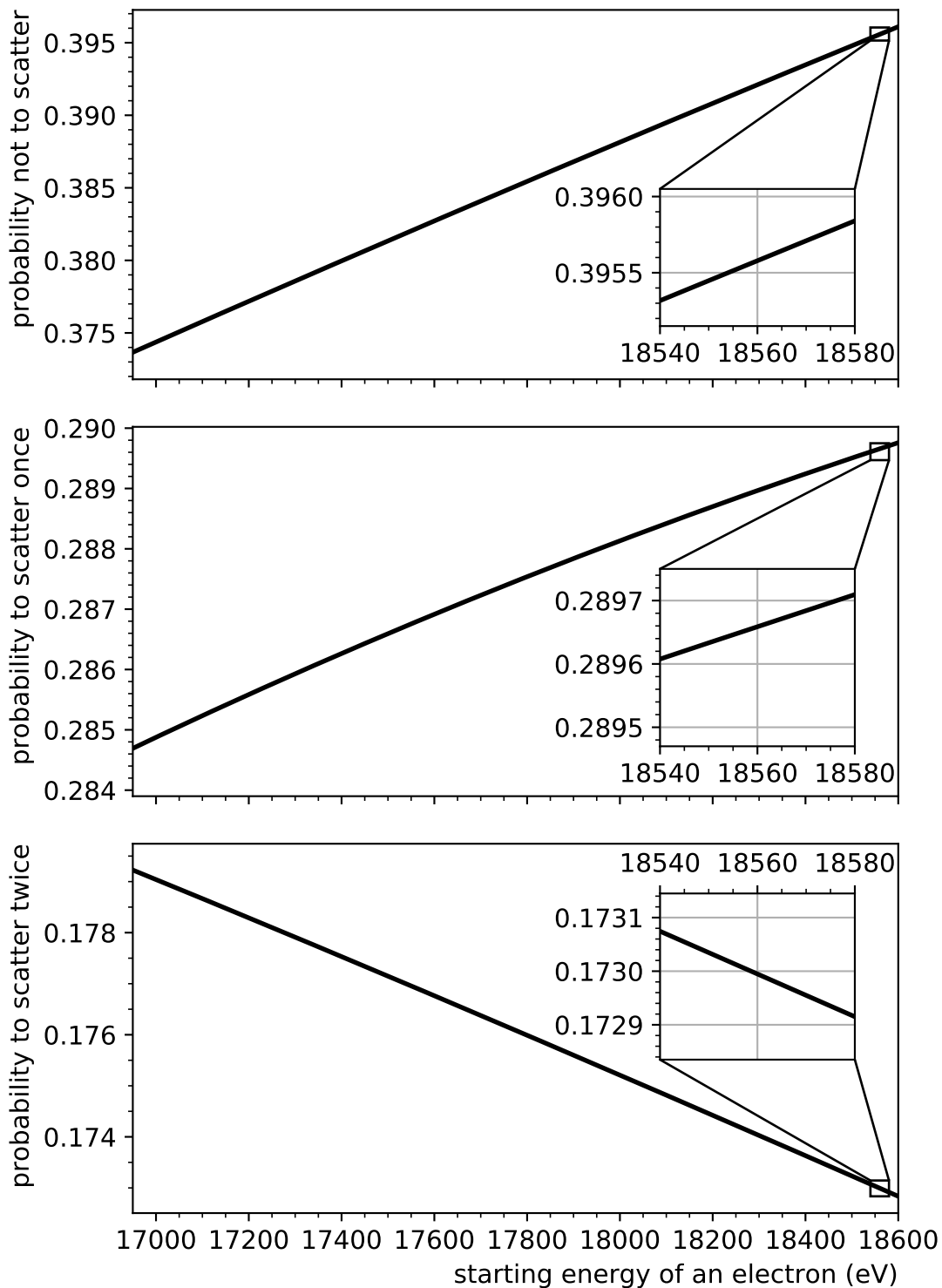
- the starting energy  $E_S$  increases
- $\Rightarrow$  the scattering cross section decreases (see figure 5.1)
- $\Rightarrow$  scattering becomes less likely
- $\Rightarrow \bar{l}$  decreases and moves closer to 1 and away from 2
- $\Rightarrow$  the probability for no and one-fold scattering increases,
- while the probability for more than one scattering decreases.

This is in accordance with the trend of the energy-dependence depicted in figure 5.2. Also, if this reasoning is correct, the sign of  $d\bar{P}_1(E_S)/dE_S$  for one-fold scattering should change for sufficiently high energies and the probability to scatter once should decrease. This is indeed the case as depicted in figure 5.3. The decrease starts at an energy of roughly 24.7 keV. This might for example be of relevance for the measurement of Krypton conversion lines in calibration measurements because the corresponding electron energies are above 30 keV [Vén+18].

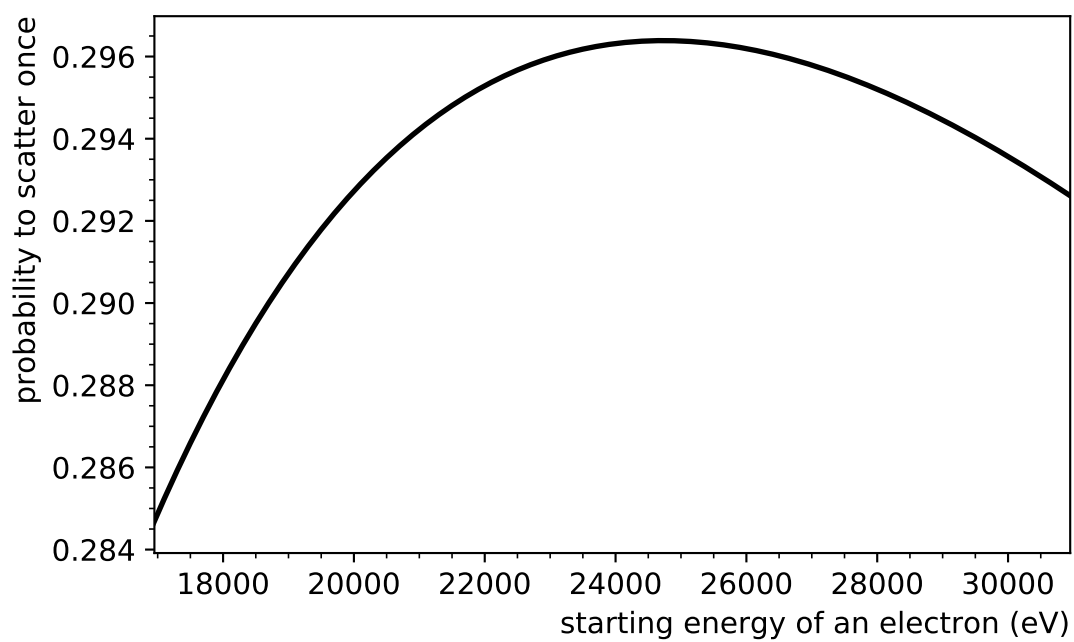
#### Limit of the Poisson Model for Electron-Scattering

The Poisson model is exact for the probability of no scattering  $\bar{P}_0(E_S)$ , but it does not hold for one or more scatterings because an electron loses energy when scattering. The scattering cross section increases with decreasing energy (see figure 5.1) and the electron becomes more likely to scatter again. In other words, the probabilities of individual scattering processes are no longer independent when respecting the dependence on the electron energy. This violates one of the preconditions to model the scattering probabilities via a Poisson distribution. Another model beyond the description via a Poisson distribution is suggested in the subsequent section 5.4. However, at the current stage, this extended model poses computational difficulties as is discussed in section 5.4.2. The aim of this thesis is the assessment of energy-dependent effects. This can already be achieved with the Poisson model as shown in the previous paragraphs. Therefore, the Poisson model was further used to address the energy-dependent effects in the scope of neutrino mass inference.





**Figure 5.2.: Probability for severalfold electron-scattering in the WGTS.** From top to bottom, the probability for no, one-fold and two-fold scattering are shown averaged over all starting positions and starting pitch angles of electrons in the WGTS according to the Poisson Model in equation (5.15). The shown energy range matches the measurement range of the FT measurement campaign and the inset shows an energy span around the endpoint of the tritium  $\beta$  spectrum.



**Figure 5.3.:** Probability for one-fold electron-scattering in the WGTS in an extended energy range. The graph shows the probability of one-fold scattering according to equation (5.15c). It emphasizes the change of sign in the derivative at approximately 24.7 keV.

## Implementation and Performance

Within the scope of this thesis, the energy-dependent Poisson model of equation (5.15) was implemented into the source and spectrum calculation (SSC) software framework. The energy-dependence of the scattering cross section may not be negligible in neutrino mass inference as is explained in the subsequent section 5.3. For that reason, the impact on the fitting run time by using an energy-dependent cross section was probed. Depending on the MTD, a fit might become slower by a factor of 40 to 120. This is due to the fact, that, when integrating over the energy loss in the response function in equation (3.29), the scattering probabilities have to be recomputed in every step of the numerical integration. It might be beneficial to investigate whether the evaluation can be speed up in the future. For example, it might be sufficient to only compute the energy-dependent scattering probabilities for a limited set of energies and interpolate them in-between those energies.

### 5.3. Effects of Systematic Cross-Section Offsets in Neutrino Mass Inference

It was investigated, how much the squared neutrino mass that is inferred from a KATRIN measurement would be shifted if the energy-dependence of the scattering cross section is neglected in the corresponding fitting procedure. The comparability to former results is of importance within this section. For that reason, the energy interpretation of equation (5.10) is used, which yields a cross section of  $\sigma_{\text{TDR}} = 3.456 \times 10^{-22} \text{ m}^2$  at an energy of 18 565 eV within the KATRIN design analysis interval. There are several previous works that studied similar effects [Ant15; Gro15; Sei19; Kuc16; Kuc+18]. In section 5.3.1, the results from [Gro15] are reviewed representatively for the mentioned works because they might intuitively contradict the results obtained in this thesis. In section 5.3.2, the results of this thesis are listed. And in section 5.3.3, an argument is given, why both sets of results are assumed to be in accordance.

#### 5.3.1. Neutrino Mass Shifts for an Energy-Independent Cross-Section Offset from Former Works

In [Gro15], it was investigated how much a constant offset of the cross section would shift the inferred squared neutrino mass if the offset were neglected in the analysis. A rule of thumb for the shift in dependence on the relative offset of the cross section  $\Delta\sigma/\sigma_{\text{TDR}}$  is given [Gro15]

$$\frac{\Delta m_{\nu}^2(\Delta\sigma)}{10^{-3} \text{ eV}^2} = -0.45 - 1204 \cdot \frac{\Delta\sigma}{\sigma_{\text{TDR}}} \quad \text{with} \quad \sigma_{\text{TDR}} = 3.456 \times 10^{-22} \text{ m}^2. \quad (5.17)$$

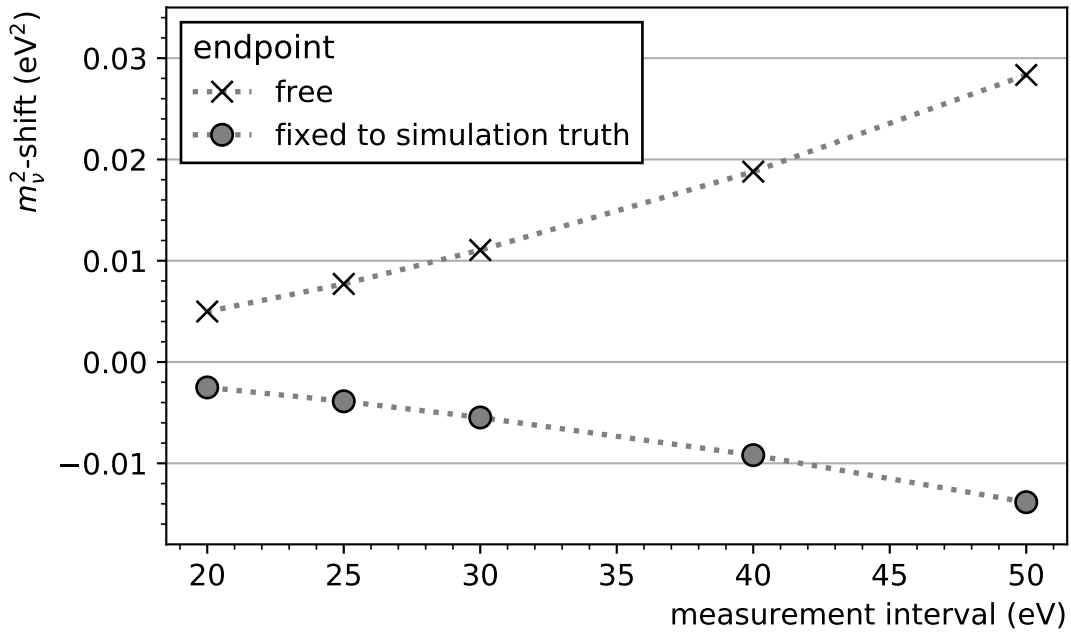
This rule of thumb was deduced via a linear approximation of several neutrino mass shifts in dependence of the relative cross section offset. Although, it can only be used for rough estimations, it serves well for the discussion in the following sections. Using the energy interpretation of the KATRIN Design Report (eq. 5.10) and the numerical inversion of the cross section formula (eq. 5.6), in the KATRIN design analysis interval of 30 eV, the relative offset of the cross section caused by the energy-dependence is

$$\left| \frac{\Delta\sigma}{\sigma_{\text{TDR}}} \right| < 0.1 \%. \quad (5.18)$$

Thus, the rule of thumb (5.17) yields

$$\left| \Delta m_{\nu}^2(\Delta\sigma) \right| < 2 \times 10^{-3} \text{ eV}^2 \quad (5.19)$$

within the KATRIN design analysis interval of 30 eV.



**Figure 5.4.: Shift of an inferred squared neutrino mass induced by a neglected energy-dependence of the inelastic scattering cross section.** The neutrino mass was inferred from data that was simulated using an energy-dependent inelastic scattering cross section as per equation (5.6) (with an energy interpretation according to equation (5.10) for compatibility with the KATRIN Design Report) and using a nominal four-parameter fit model that assumes a constant cross section of  $\sigma_{\text{TDR}} = 3.456 \times 10^{-22} \text{ m}^2$ . The procedure was repeated for five different measurement intervals (respectively MTDs)  $[E_0 - \alpha, E_0 + 5 \text{ eV}]$  based on the KATRIN Design Report where  $\alpha$  is denoted on the abscissa and  $E_0 = 18\,575 \text{ eV}$  is the endpoint of the  $\beta$  spectrum. The MTDs are listed in appendix B. The markers show the results for the obtained shift of the squared neutrino mass and the dotted line is an interpolation in-between. The study was done once with a free fit parameter for the endpoint and once having it fixed to the simulation truth in order to show the influence on the sign of the shift of the squared neutrino mass (for details refer to the main text).

### 5.3.2. Neutrino Mass Shifts for an Energy-Dependent Cross-Section Offset

In the scope of this thesis, an energy-dependent offset of the cross section was investigated (in contrast to a constant one in [Gro15]). A KATRIN neutrino mass measurement for a neutrino mass of  $0 \text{ eV}$  was simulated using an energy-dependent cross section<sup>5</sup>. A model that uses a constant cross section was fitted to the simulated spectrum. This procedure was repeated for five different MTDs of different measurement intervals. Figure 5.4 shows the results. For the KATRIN design measurement interval of  $30 \text{ eV}$ , the difference from  $0 \text{ eV}$ , respectively the shift of the squared neutrino mass, is

$$\Delta m_\nu^2 = 1.09 \times 10^{-2} \text{ eV}^2. \quad (5.20)$$

First, it should be noted that this value is greater by a factor of  $\sim 5$  than the one stated above in equation (5.19), that was estimated using the results from [Gro15]. This issue is

<sup>5</sup>In all studies presented in this chapter, the simulated electron detector counts were substituted by their expectation values as opposed to fluctuated according to Poissonian statistics. In other words, no ensemble testing was done. For more details on such an Asimov data set and why it can reasonably be assumed to be representative for an ensemble, see section 6.3.3.

addressed in the subsequent section 5.3.3. Further aspects are discussed in the following paragraph:

### Sign of the Shift of the Squared Neutrino Mass

Using an energy-dependent cross section in the simulation and neglecting it in neutrino mass inference yields a higher  $\beta$  electron count in the endpoint region than expected (see figure 5.6) because the cross section decreases with higher energies and electrons of higher energies are less probable to scatter. On one hand, a higher  $\beta$  electron count for energies in the endpoint region of the  $\beta$  spectrum can be caused by a lower squared neutrino mass. On the other hand, a higher  $\beta$  electron count can also be caused by a higher endpoint of the  $\beta$  spectrum. These two effects counterbalance each other, which makes it difficult to make an intuitive, qualitative statement about the sign of the shift of the squared neutrino mass. As shown in equation (5.20), the shift of the squared neutrino mass is positive. However, if the above reasoning is correct, then one should obtain a flip in the sign of the shift if the endpoint is not taken as a free fit parameter, but fixed to the simulation truth. Figure 5.4 shows that this is indeed the case.

### Significance for the KATRIN Experiment

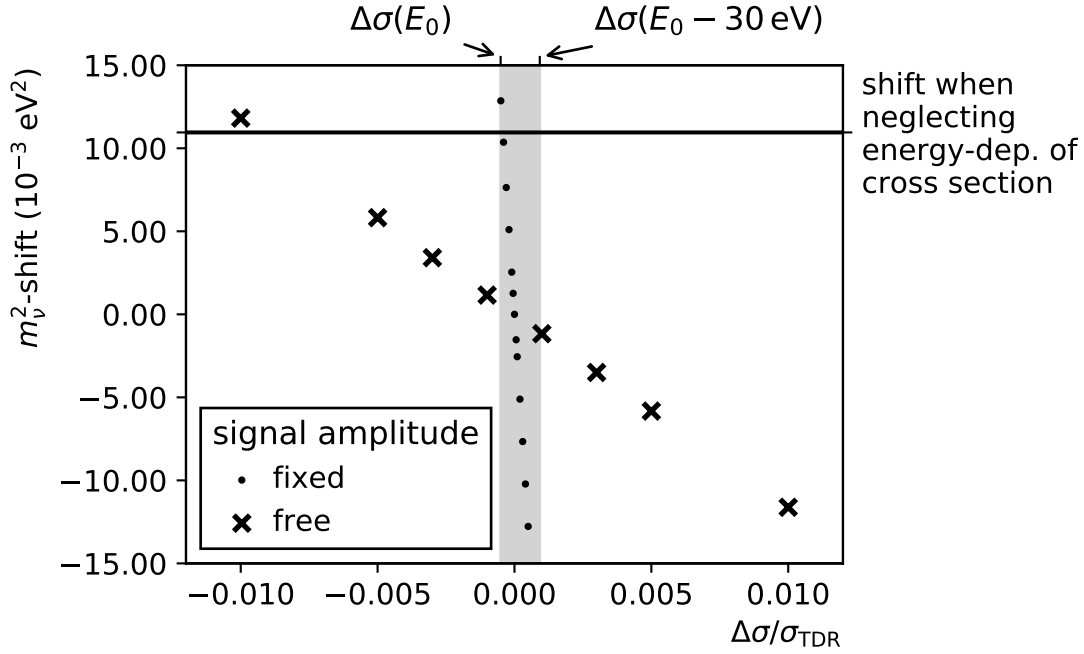
The shift of the squared neutrino mass in equation (5.20) takes a sizable fraction of the systematic budget of  $1.7 \times 10^{-2} \text{ eV}^2$  (see eq. 4.13) of the KATRIN experiment, but is still smaller. Hence, whether the effect of the energy-dependence of the scattering cross section can be neglected depends on the contribution of other systematic effects and can not yet fully be assessed (see for example [Sei19] for a comprehensive list of systematic effects with regard to the KATRIN experiment).

It should be mentioned that KATRIN's design allows to measure the product of the gas column density and the inelastic scattering cross section  $\rho d \cdot \sigma(E)$  for a fixed electron energy  $E$  using the electron gun. In order to respect the energy dependence of the cross section, it is necessary to incorporate a corresponding model, such as the formula for the energy-dependent cross section (eq. 5.6). This might enable the measurement of the two constants  $C_1$  and  $C_2$  in said formula. Therefore, electron spectra have to be recorded for multiple energies  $E$ . Such measurements were already conducted at the KATRIN experiment. However, results that respect the energy-dependence of the scattering cross section in the way described here are not yet available at the time of writing this thesis. But a corresponding analysis should be possible.

### 5.3.3. Relation of the Shifts for an Energy-Dependent and -Independent Cross-Section Offset

In the scope of this thesis, the shift of an inferred neutrino mass that is introduced by neglecting the energy-dependence of the cross section was investigated. Former works investigated the shift when neglecting a constant offset of the cross section. The shift obtained in this thesis in equation (5.20) is  $\sim 5$  times larger than the one obtained by the former study in [Gro15] under equation (5.19), although the energy-dependent offset of the cross section is never greater than the constant offset investigated by [Gro15]. This might be contra intuitive, but a possible explanation is the following: The fit parameter for the signal amplitude  $A_S$  in a nominal KATRIN fit (see section 4.3) can partly compensate for a constant offset of the cross section. However, it cannot compensate for an energy-dependent one equally well. This is demonstrated below:

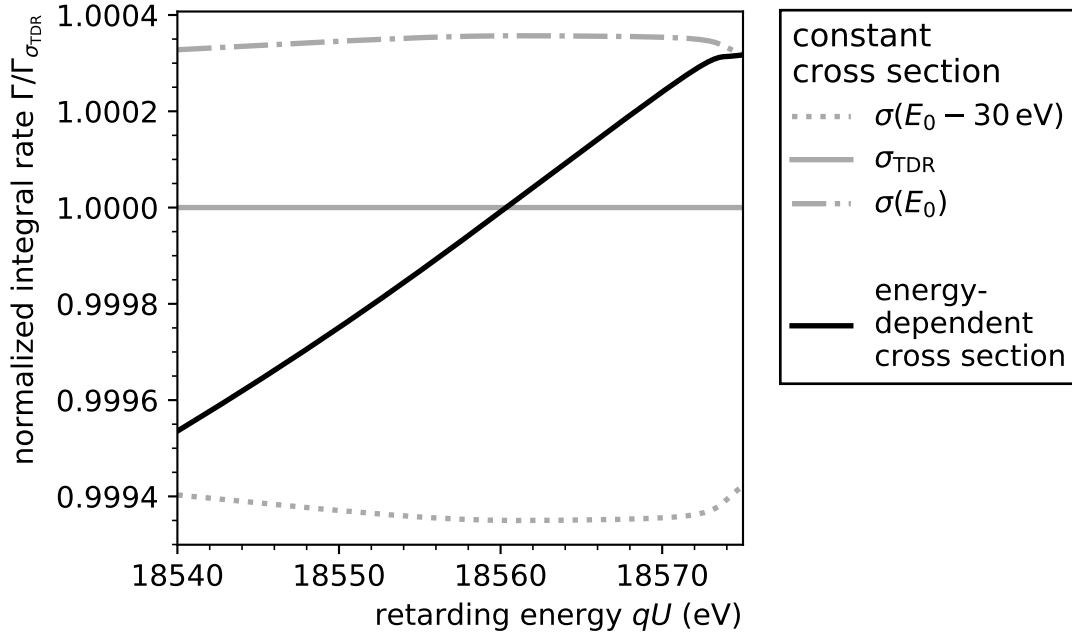
The study from [Gro15] was reproduced with a free fit parameter for the signal amplitude  $A_S$ . The reproduced study was in accordance with the values in [Gro15]. Then,  $A_S = 1$  was fixed.



**Figure 5.5.: Shift of an inferred squared neutrino mass induced by a neglected constant offset of the inelastic scattering cross section.** The neutrino mass was inferred from data that was simulated using different constant inelastic scattering cross sections. The abscissa shows the offset of the used inelastic scattering cross section relative to  $\sigma_{\text{TDR}} = 3.456 \times 10^{-22} \text{ m}^2$ , which was used in the assumed fit model. In an energy-dependent scenario, a cross section corresponds an energy of incident electrons as per equation (5.6). The cross section range, that corresponds the design KATRIN analysis energy interval of 30 eV is depicted as a gray band (energy interpretation as per eq. 5.10). The analysis was done twice: once with a free parameter for the signal amplitude, which reproduced the results given in [Gro15] (figure 6.31 on page 221); the second analysis had this parameter fixed. The shift obtained for  $\Delta\sigma/\sigma_{\text{TDR}} \approx -0.1\%$  and a fixed signal amplitude is greater than the shift obtained when neglecting the energy-dependence of the cross section within the 30 eV measurement interval (depicted by the horizontal line).

The results are shown in figure 5.5. A relative cross section offset of  $|\Delta\sigma/\sigma_{\text{TDR}}| \approx 0.1\%$  (see equation 5.18) yields a shift of the squared neutrino mass, that is larger by almost an order of magnitude with a fixed  $A_S$  as opposed to leaving  $A_S$  as a free parameter. The shift is then greater than the one obtained in the study presented in the previous section 5.3.2 that uses an energy-dependent cross section in the 30 eV measurement interval. This verifies, that the fit parameter for the signal amplitude  $A_S$  can partly compensate for a constant offset of the cross section.

Additionally, the following plausibility argument can be given: Figure 5.6 shows the simulated integral rate for different cross section models normalized to the integral rate obtained by using the model with a constant cross section  $\sigma_{\text{TDR}}$  from the KATRIN Design Report. The integral rates that are based on a model with a constant, energy-independent cross section differ from each other by almost a constant factor (variations  $< 2 \times 10^{-4}$ ). This implies, when one of these models is used in a simulation of a KATRIN measurement and another model is used in a corresponding fit for neutrino mass inference, their difference can be compensated by the fit parameter for the signal amplitude  $A_S$ . However, the model that uses an energy-dependent cross section differs from the one with the constant cross section  $\sigma_{\text{TDR}}$  by a factor that varies over the energy range. Such a difference can not be compensated by the fit parameter for the signal amplitude  $A_S$ .



**Figure 5.6.: Simulated integral  $\beta$ -electron rate using different models of the cross section for inelastic electron-scattering.** The abscissa shows the energy range of the design KATRIN analysis interval of 30 eV. The lines show the integral rate of  $\beta$  electrons for different models for the inelastic scattering cross section normalized to the model that uses the constant cross section  $\sigma_{\text{TDR}} = 3.456 \times 10^{-22} \text{ m}^2$ . The gray lines use a model with a constant cross section. The cross section values can be identified with electron energies as per equation (5.6) and the energy interpretation of equation (5.10) with  $\sigma(E_0 - 30 \text{ eV} = 18\,545 \text{ eV}) = 3.459 \times 10^{-22} \text{ m}^2$  and  $\sigma(E_0 = 18\,575 \text{ eV}) = 3.454 \times 10^{-22} \text{ m}^2$ . The black line shows the simulated integral rate for an energy-dependent cross section. This graph emphasizes that the integral rates for models using a constant cross section differ by an almost constant factor, whereas the rate using an energy-dependent cross section differs from the others by a varying factor.

#### 5.4. An Energy-Dependent Scattering Model beyond Using the Poisson Distribution

In the previous section 5.2 the scattering probabilities are modeled using a Poisson distribution. This introduces an error for the following reason: A scattering electron loses energy. The scattering cross section increases with decreasing energy (see figure 5.1) and the electron becomes more likely to scatter again. In other words, the probabilities of individual scattering processes are not independent when respecting the dependence on the electron energy. This violates one of the preconditions to model the scattering probabilities via a Poisson distribution.

This section aims to quantify the error that is introduced by the approximate modeling via a Poisson distribution. Another model is suggested<sup>6</sup> in section 5.4.1 that partly accounts for the energy loss of electrons in the formalism of the scattering probabilities. The suggested model was evaluated numerically as it includes one limit and two integrals that could not yet be solved analytically. The numerical accuracy had to be good enough to decide whether it differs significantly from the Poisson model or not. At the same time, a balance between the numerical accuracy and the evaluation run time had to be found. This matter

<sup>6</sup>The idea was inspired by a model in [Gro15] that incorporates changes of the electron pitch angle due to scattering.

is described in detail in section 5.4.2. The results of the numerical evaluation are discussed in section 5.4.3.

#### 5.4.1. Formalism of the Extended Model for Electron-Scattering

In the following, a model for the probability of electron-scattering in the WGTS that depends on the electron energy is derived. The difficulty lies in the fact, that an electron that scatters loses energy and hence its probability to scatter a second time is different from the first time. The presented model assumes the same fixed energy loss  $\epsilon$  for each scattering in contrast to a realistic distribution as given by the energy loss function (see figure 3.3). This was sufficient to show a difference in the Poisson model and the model suggested below. However, a more accurate description would incorporate the full energy loss function. This may be the subject of a future study.

The aim of this section is the derivation of an expression  $\bar{P}_l^*(E_S)$  that denotes the probability of  $l$ -fold scattering for an electron with a starting energy  $E_S$  averaged over all starting positions and pitch angles assuming a fixed energy loss  $\epsilon$  per scattering.

The expected amount of scatterings for an electron when traveling through the whole WGTS volume of length  $d$  filled with a gas of constant density  $\rho$  is (see equation 3.17)

$$\mu(E, \theta_S) = \frac{\sigma(E)\rho d}{\cos \theta_S}, \quad (5.21)$$

where  $E$  denotes the electron's kinetic energy;  $\theta_S$  the starting pitch angle; and  $\sigma(E)$  the energy dependent scattering cross section. It should be noted that the energy  $E$  is denoted without and index S here, because, in the following, this formula is reused for different electron energies, not only for the starting energy of an electron.

In the following, the volume of the WGTS is thought of being divided into  $N$  slices of equal width  $w = L/N$  (and later the limit  $N \rightarrow \infty$  is applied).  $N$  is chosen sufficiently large that the probability for an electron to scatter twice within one slice is essentially zero. Then, for large  $N$ , the probability to scatter within one slice is  $\mu(E, \theta_S)/N$ . The probability not to scatter within  $n \leq N$  slices is

$$p_0(E, \theta_S, n) = \left(1 - \frac{\mu(E, \theta_S)}{N}\right)^n. \quad (5.22)$$

Let "Poisson( $\mu, l$ )" denote the Poisson distribution with expectation value  $\mu$  and evaluated at  $l$ . Using the well known limit for the Euler constant, one obtains for  $n = N$  and  $N \rightarrow \infty$  that  $p_0$  is a Poisson distribution with the expectation value  $\mu$  from equation (5.21) and evaluated at  $l = 0$

$$\lim_{N \rightarrow \infty} p_0(E, \theta_S, N) = \lim_{N \rightarrow \infty} \left(1 - \frac{\mu(E, \theta_S)}{N}\right)^N = e^{-\mu(E, \theta_S)} = \text{Poisson}(\mu(E, \theta_S), 0). \quad (5.23)$$

In other words, for no scattering, the Poisson model for the scattering probabilities as described in section 5.2 is recovered.

Assuming a constant energy loss  $\epsilon$  per scattering, the probability to scatter  $l > 0$  times within  $n < N$  slices can be expressed recursively

$$p_l(E, \theta_S, n) = \underbrace{\sum_{k=l}^n}_{(4)} \underbrace{p_{l-1}(E, \theta_S, k-1)}_{(1)} \underbrace{(1 - p_0(E - (l-1)\epsilon, \theta_S, 1))}_{(2)} \underbrace{p_0(E - l\epsilon, \theta_S, n-k)}_{(3)}. \quad (5.24)$$



The idea behind this expression is the following: One imagines an electron with an energy  $E$  that travels through  $n$  WGTS slices, scatters  $l$  times in total and the last time in the  $k$ th slice. Hence, it must have scattered  $l - 1$  times in the  $k - 1$  slices before the  $k$ th slice and must not scatter in the  $n - k$  slices to follow. In that regard, the above expressions have the following meaning:

- (1) Probability to scatter  $l - 1$  times within  $k - 1$  slices with a kinetic energy of  $E$ .
- (2) Probability to scatter once within the  $k$ th slice with a kinetic energy of  $E - (l - 1)\epsilon$ .
- (3) Probability not to scatter within the remaining  $n - k$  slices with a kinetic energy of  $E - l\epsilon$ .
- (4) Sum over all slices  $k$  where the electron could scatter the last time. The sum starts at  $l$  (as opposed to 1) because the probability to scatter  $l - 1$  times within less than  $k = l - 1$  slices (term (1)) is 0 due to the made assumption, that in the limit of a large amount of slices  $N$  an electron does not scatter twice within one slice.

The probability to scatter  $l$  times can be averaged over all starting positions of electrons respectively starting slices. In this regard,  $n_S$  denotes the amount of slices an electron has to pass through depending on its starting slice (including its starting slice). Then the average is

$$\bar{p}_l(E, \theta_S) = \frac{1}{N} \sum_{n_S=1}^N p_l(E, \theta_S, n_S) \approx \frac{1}{d} \int_0^d p_l(E, \theta_S, \left[ N \frac{z_S}{d} \right]) dz_S. \quad (5.25)$$

Here, the averaging sum is approximated by an integral as this helps cutting down on run time in a numerical evaluation. This is due to the fact, that in a numerical evaluation a large number  $N$  ( $\sim 10^5$ ) of slices has to be chosen and the sum would have many terms. However, a numerical integration can achieve an accurate result with a smaller set of supporting points (see subsequent section 5.4.2).

Then, the limit  $N \rightarrow \infty$  can be applied

$$P_l^*(E_S, \theta_S) = \lim_{N \rightarrow \infty} \bar{p}_l(E_S, \theta_S). \quad (5.26)$$

$P_l^*(E_S, \theta_S)$  denotes the probability for an electron to scatter  $l$  times when traveling through the whole WGTS with a starting energy  $E_S$  and pitch angle  $\theta_S$  averaged over all starting positions. Finally, this expression can be averaged over all starting pitch angles in order to obtain the energy dependent scattering probabilities

$$\bar{P}_l^*(E_S) = \frac{1}{1 - \cos \theta_{\max}} \int_0^{\theta_{\max}} \sin \theta_S P_l^*(E_S, \theta_S) d\theta_S. \quad (5.27)$$

$\bar{P}_l^*(E_S)$  denotes the probability for  $l$ -fold scattering of an electron with a starting energy  $E_S$  averaged over all starting positions and pitch angles assuming a fixed energy loss  $\epsilon$  per scattering. In the following, this model is called the “extended model”.

#### 5.4.2. Numerical Accuracy and Cross-Check of the Extended Model for Electron-Scattering

For the evaluation of the energy-dependent scattering probabilities according to the extended model in equation (5.27) the fixed energy loss  $\epsilon = 12.6$  eV per scattering was chosen as it is the most probable energy loss for electrons traveling through tritium gas (see figure 3.3). The mean value would also have been a possible choice. But the energy loss function as depicted in figure 3.3 is not normalized to  $[0, \infty)$ , which poses the problem of finding

a suitable energy interval, with respect to which the mean value is calculated. In other words, at the current stage, a certain arbitrariness concerning the choice of the energy loss per scattering could not be avoided. As already mentioned, a future study might consider a more comprehensive incorporation of the energy loss function.

The extended model in equation (5.27) was evaluated numerically:

- Taking the limit  $N \rightarrow \infty$  in equation (5.26) was replaced by choosing a large  $N$ .
- The averaging integral over the starting positions in equation (5.25) and starting pitch angles in equation (5.27) were computed using Gaussian quadrature.

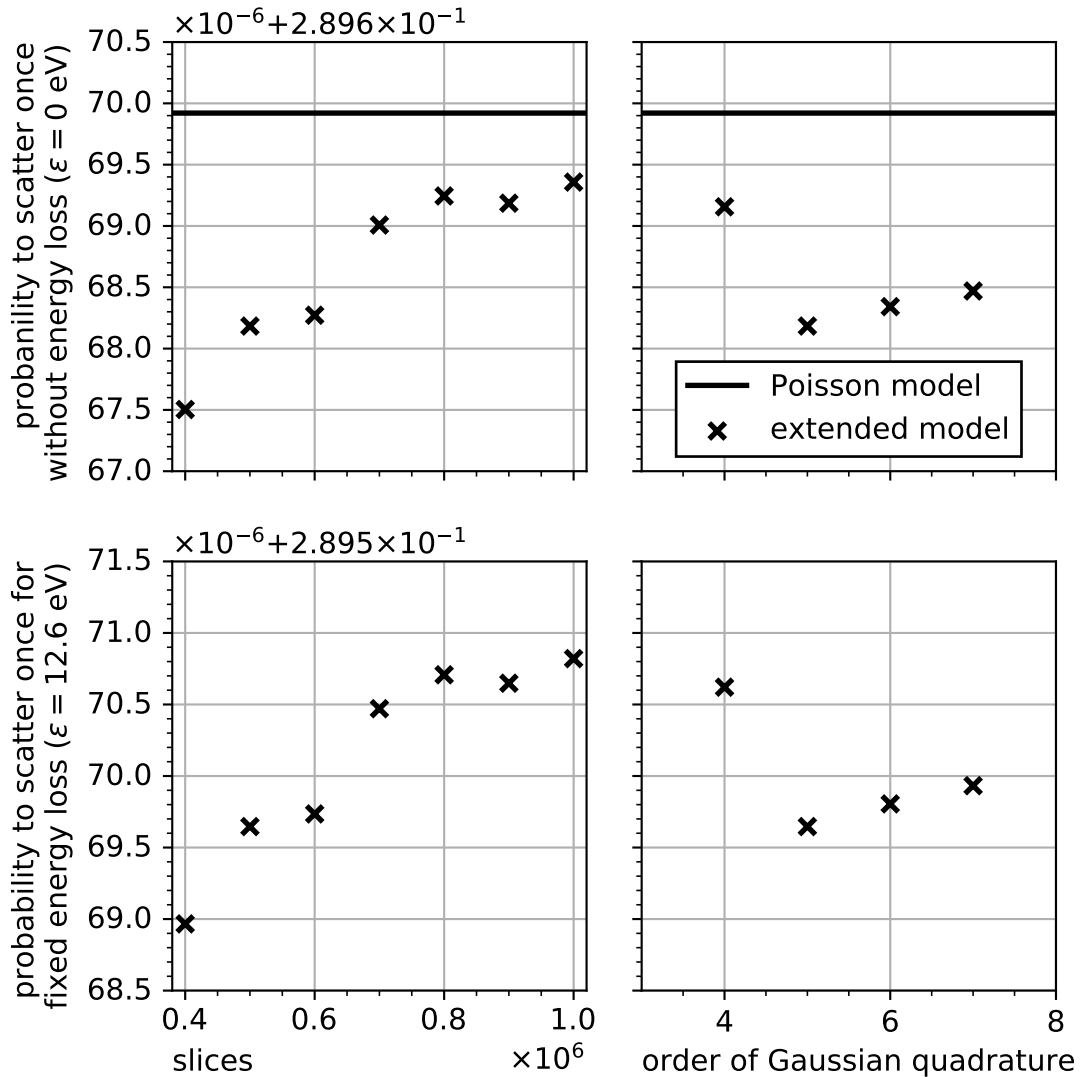
The extended model was introduced because the preconditions to model the scattering probabilities via a Poisson distribution (Poisson model) do not hold. Hence, the numerical evaluation of the extended model must be sufficiently accurate to show the difference to the Poisson model. How accurate this is was not known beforehand and was found out by trial and error. Both,  $N$  and the order of Gaussian quadrature, should be chosen as low as possible to cut down on calculation run time, but sufficiently high for the required accuracy.

A benchmark had to be found in order to determine the accuracy of the numerical evaluation. The following two ideas were used: First, for an energy loss of  $\epsilon = 0$  eV per scattering, the extended model must recover the Poisson model exactly. This can be used to estimate the numerical accuracy in dependence of the number  $N$  of slices and order of Gaussian quadrature. The estimated accuracy when using  $\epsilon = 0$  eV may then be assumed for evaluations when  $\epsilon > 0$  eV. Second, a further cross-check is the convergence of the numerical evaluation with increasing  $N$  and an increasing order of the Gaussian quadrature. Both ideas find application below.

The averaged probability for one-fold scattering  $\bar{P}_1^*$  of the extended model was evaluated for  $\epsilon = 0$  eV. The result is shown in the top row of figure 5.7 in dependence of  $N$  and the order of the Gaussian quadrature. It should recover the Poisson model. For  $N = 5 \times 10^5$  and using Gaussian quadrature of order 5, the Poisson model and the extended model differ less than  $3 \times 10^{-6}$ . The calculations for  $\epsilon = 12.6$  eV are shown in the lower row of figure 5.7. They also show convergence on the  $10^{-5}$  level. Conclusively, the results make it plausible to assume a numerical accuracy on the  $10^{-5}$  level or better for  $N = 10^5$  and using Gaussian quadrature of order 5 for the integrals. Using this configuration, the extended model for one-fold scattering was numerically evaluated for different electron starting energies. The result is depicted in figure 5.8. It was found that the extended model and the Poisson model differ by approximately  $10^{-4}$ . Hence, the numerical accuracy on the  $10^{-5}$  level is sufficient to show the difference of these two models.

The corresponding run time to compute  $\bar{P}_1^*$  is in  $\mathcal{O}(N)$  as it requires a sum over all  $N$  slices in equation (5.24). The extended model is defined recursively and therefore the run time for  $l$ -fold scattering is in  $\mathcal{O}(N^l)$ . Hence, computing the probability for two-fold scattering would take  $5 \times 10^5$  times as long as for one-fold scattering for the same  $10^{-5}$  accuracy. This was not yet found to be feasible.

The results of the numerical evaluation for one-fold scattering are discussed in the subsequent section 5.4.3.



**Figure 5.7.:** Accuracy of the numerical evaluation of the extended model of the probability for electron-scattering in the WGTS. The extended model is given by equation (5.27) and the Poisson model is given by equation (5.15). For both models, the energy of the incident electrons was taken such that the Poisson model yields the values already given in table 5.1. The left column shows the dependence on the number  $N$  of slices of the WGTS. The right column shows the dependence on the order of Gaussian quadrature that was used to evaluate the two integrals in the extended model. The Poisson model is independent of these features and can be evaluated exactly (limited by machine number precision). In the left column, the order of Gaussian quadrature was fixed to 5 and in the right column, the number of slices was fixed to  $5 \times 10^5$ . The upper row shows the numerical evaluation of the extended model for no energy loss per scattering (markers). Its exact solution is given by the Poisson model (line). The lower row shows the extended model for an energy loss of 12.6 eV per scattering. The upper row shows that the numerical evaluation of the extended model is within a distance of less than  $3 \times 10^{-6}$  of its exact solution and the lower row shows convergence on the  $10^{-5}$  level. These results make it plausible to assume a numerical accuracy on the  $10^{-5}$  level or better for the shown configurations of the numerical evaluation of the extended model.

### 5.4.3. Discussion of the Extended Model for Electron-Scattering

For no scattering, the extended model and the Poisson model are equivalent as shown in equation (5.23).

Figure 5.8 depicts the Poisson model as per equation (5.15) and the extended model as per equation (5.27) for the probability of one-fold scattering. The difference between the two models is below  $10^{-4}$ .

For two-fold scattering and higher scattering multiplicities the extended model could not yet be evaluated due to the computational complexity. However an estimation based on theoretical arguments is given below.

#### Difference of the Extended and the Poisson Model for Electron-Scattering

The reasoning concerning the sign of the difference between the Poisson and the extended model is similar to the one in section 5.2 about the change of the scattering probabilities with decreasing or increasing energy: The Poisson model underestimates the total amount of scattering, because it neglects that an electron that scatters loses energy and becomes more likely to scatter again. As already shown, according to the Poisson model, electrons are expected to scatter between one and two times when traveling through the WGTS (see equation 5.16). The extended model is more correct than the Poisson model and hence should yield an increased expected amount of scattering. Therefore, the probability for one-fold scattering should be smaller than the one given by the Poisson model (this is the case as depicted in figure 5.8). For higher scattering multiplicity (which could not be calculated), the sign of the difference between the Poisson and the extended model should be the opposite. In other words, the probability of two-fold scattering as given by the Poisson model should be too low.

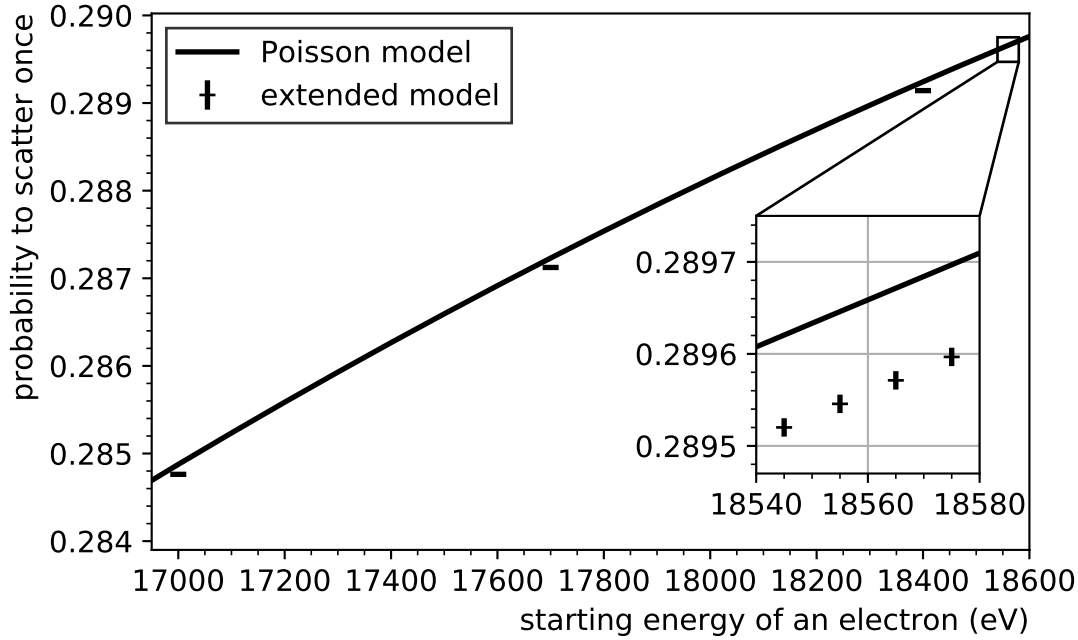
Also, the extended model should be a probability distribution and the sum over all scattering multiplicities should yield unity in the same manner as it does for the Poisson model (see appendix A for a proof that the Poisson model is a probability density). This means that for higher scattering multiplicities, the difference between the extended model and the Poisson model would have to sum up to the difference in the probability for one-fold scattering. Hence, for scattering of higher multiplicity, the absolute difference between the extended model and the Poisson model would not be able to surpass  $10^{-4}$ .

#### The Extended Model for Electron-Scattering in Neutrino Mass Inference

How the error of the Poisson model propagates to neutrino mass inference and whether it would introduce a significant systematic shift of the squared neutrino mass was not probed in the scope of this thesis because the current stage of the extended model is computationally too demanding to be used in a fitting procedure. However, some considerations are given below:

The difference between the extended model and the Poisson model for one-fold scattering is  $\sim 10^{-4}$ , which might be a significant scale because it is similar to the scale on which the probability for one-fold scattering changes when transitioning from an energy-independent to an energy-dependent cross section (see insets of figure 5.2). In the previous section 5.3.2, it is shown that neglecting such an effect might lead to a systematic shift of the squared neutrino mass on the order of  $10^{-2} \text{ eV}^2$ . Nonetheless, the difference of the extended and the Poisson model might be partly compensated by the different fit parameters such as the signal amplitude or the endpoint of the  $\beta$  spectrum. Therefore, it was found difficult to make quantifiable assumptions.

In the following, a procedure is outlined, that might make it possible to estimate the induced systematic shift of the squared neutrino mass in a future study: One could simulate



**Figure 5.8.:** Comparison of the Poisson and the extended model for one-fold scattering of electrons in the WGTS. The line shows the Poisson model and the markers the extended model (see main text for a description of the models). Due to long program run times, the latter was only evaluated at a selection of energies. Furthermore, the numerical evaluation of the extended model is subject to an error of  $\sim 10^{-5}$  (see section 5.4.2), that is depicted as uncertainty bars. The difference of the two models is  $\sim 10^{-4}$ .

data using the Poisson model and additionally introduce an artificial negative shift on the order of  $10^{-4}$  on the probability for one-fold scattering and a positive shift that is smaller than  $10^{-4}$  for the probability for scattering of higher multiplicities because this should approximately resemble the extended model. The fit model can then use the Poisson model without modification. This procedure might make it possible to quantify whether there is a significant shift of the inferred squared neutrino mass caused by neglecting that electrons become more likely to scatter again if they have already scattered.

## 5.5. Conclusion and Outlook

The cross section for inelastic scattering off tritium molecules within the WGTS depends on the energy of the incident electrons. An approximated model that applies the Poisson distribution (Poisson model) was used to estimate the systematic shift on the squared neutrino mass that would be introduced if the energy-dependence were neglected. The shift would be  $\Delta m_\nu^2 = 1.09 \times 10^{-2} \text{ eV}^2$  in a 30 eV analysis interval below the endpoint of the  $\beta$ -tritium spectrum. However, the Poisson model can be used in neutrino mass analysis and hence a corresponding systematic shift can be avoided.

It was also shown that including the energy-dependence in neutrino mass inference via the Poisson model increases the run time of the fitting procedure by a factor of up to 140 depending on the MTD. A suggestion for an improvement of the performance via interpolation of the scattering probabilities is made in section 5.2.2.

The usage of the Poisson model underestimates the total amount of scattering because it neglects that an electron that scatters loses energy and becomes more likely to scatter again. This does not influence the probability for no scattering. However, using the approximation

of a fixed energy loss  $\epsilon = 12.6 \text{ eV}$  per scattering, it could be shown that the probability for one-fold scattering would be shifted by less than  $10^{-4}$ . The effect on scatterings of higher multiplicities or the effect in neutrino mass inference could not be fully quantified, but an approach that might enable the quantification in the future is outlined in section 5.4.3.

The model building in this chapter was based on well argued considerations. Nonetheless, an independent cross-check is recommended. A future study might perform a Monte Carlo particle tracking simulation in order to verify the suggested models presented in this thesis.

## 6. Statistical Methods in the Context of the Empirical Energy Loss Model Derived from KATRIN Data for Electrons Scattering Inelastically off Deuterium

A quantitative accurate description of the scattering processes of  $\beta$  electrons within KATRIN's gaseous tritium source is of crucial importance for KATRIN's sensitivity goal. In modeling the corresponding effects, the energy loss function (see section 3.2.5) plays an important role. The KATRIN Design Report states that the precision of the energy loss functions from literature is not sufficient for KATRIN's envisaged sensitivity. It was planned to deduce a sufficiently accurate model from data taken at KATRIN [KAT05]. In that regard, a preliminary model has successfully been established for electrons scattering off deuterium molecules based on data taken in October 2018 by a designated subgroup of the KATRIN collaboration. In the following, this model is referred to as the "KATRIN model" or the "KATRIN energy loss model". This preliminary energy loss model has partially improved uncertainties with respect to the model by [Ase+00], that is presented in the previous section 3.2.5, and which was used in many previous works [Gro15; Kle14; Kle+19; Sei19]. In the following, the latter model is referred to as the "Aseev model" after the primary author of the corresponding publication [Ase+00]. In addition to the improved uncertainties, the KATRIN model also exhibits features not present in the Aseev model but motivated by data (see section 6.1 and figure 6.1).

As the uncertainties associated with the energy loss function influence KATRIN's sensitivity to the neutrino mass, it is of importance to probe the KATRIN energy loss model in neutrino mass inference and provide corresponding feedback to the team that measures it. Beyond that, the implementation of the required statistical tools into the available software frameworks is of general interest as the approach taken in the scope of this thesis may be applied to other model uncertainties, not only to the ones stemming from the energy loss function.

This chapter is structured as follows: Section 6.1 outlines the KATRIN energy loss model. Section 6.2 discusses the scope of the validity of this study; for example what is expected of the comparison of a model for electrons scattering off deuterium and another model for scattering off tritium molecules. Section 6.3 introduces statistical tools for uncertainty treatment within neutrino mass inference. Special emphasis is put on the use of additional free fit parameters to mitigate systematic effects. These methods find application in section 6.4 in order to evaluate systematic effects as induced by the energy loss model. And section 6.5 concludes and offers an outlook.

## 6.1. The Empirical KATRIN Energy Loss Model

The KATRIN energy loss model was established by a designated subgroup of the KATRIN collaboration. It is partly of empirical nature and partly based on first principles. The model was fitted to data taken at the KATRIN experiment in Fall 2018. A description and a parametrization is given in the following. Section 6.1.1 describes the KATRIN energy model itself and section 6.1.2 describes the fit it was obtained from because there are further fit parameters that can not necessarily be neglected in the uncertainty treatment. It should be noted, that a unified notation for the KATRIN energy loss model has not yet been established and that the notation in this thesis was chosen without following any reference in particular.

### 6.1.1. Description and Parametrization of the KATRIN Energy Loss Model

In the preceding section 3.2.5, the energy loss function  $f_1(\epsilon)$  is introduced. It denotes the probability density for an energy  $\epsilon$  that an electron loses when scattering once (inelastically in the current context). The KATRIN model is such an energy loss function.

The KATRIN model distinguishes two types of energy losses when an electron scatters inelastically off a gas molecule [Pre-Han19-2]. The first type is the excitation of electron states within the target molecule for impact energies below the ionization energy of deuterium  $E_{\text{ion,D}_2} = 15.467 \text{ eV}$  [Shi+93]. The second type is the ionization of the target molecule for impact energies above  $E_{\text{ion,D}_2}$ . The most probable energy losses are among the ones caused by the excitation of electronic states. In other words, the KATRIN model has a peak around these energies between approximately 12 eV and  $E_{\text{ion,D}_2}$ . In the following, this peak is denoted as the “excitation peak”. Also, the part of the KATRIN model for energies above  $E_{\text{ion,D}_2}$  is denoted “ionization tail”.

The excitation peak is empirically modeled by a sum of three scaled Gaussian distributions  $\mathcal{N}$ . This finds inspiration in the three contributions to the vibrational molecular hydrogen states by the Lyman band, the Werner band and higher terms [Pre-Han19-1] (also see [Gei64]). The excitation peak region  $f_{\text{exit}}^{\text{emp}}$  is described by nine parameters  $\boldsymbol{\pi}_{\text{eloss}}$  which comprise the three scales  $A_i$ , means  $m_i$  and standard deviations  $s_i$  ( $i \in \{1, 2, 3\}$ ) of the three Gaussian distributions [Pre-Han19-2]

$$\boldsymbol{\pi}_{\text{eloss}} = (A_1, m_1, s_1, A_2, m_2, s_2, A_3, m_3, s_3)^\top, \quad (6.1)$$

$$f_{\text{exit}}^{\text{emp}}(\epsilon \mid \boldsymbol{\pi}_{\text{eloss}}) = \sum_{i=1}^3 A_i \cdot \mathcal{N}(\epsilon, \mu = m_i, \sigma = s_i). \quad (6.2)$$

The second part, the ionization tail, follows a modified version of the binary-encounter-dipole (BED) model. For the mathematical expression, the reader is referred to [KR94]. Here, this part will be denoted

$$f_{\text{ion}}^{\text{BED}}(\epsilon). \quad (6.3)$$

The BED model is valid for scattering off hydrogen molecules. Therefore, it depends on the ionization energy of hydrogen. The corresponding value was exchanged for the one of deuterium  $E_{\text{ion,D}_2}$ . Whether further modifications with regard to the difference in isotopologues are necessary is currently under investigation. Furthermore, the constant normalization factor of the BED model was removed for the following reason: The transition between the two parts, the excitation peak and the ionization tail, is introduced at the ionization energy of deuterium  $E_{\text{ion,D}_2}$  and in order for the transition to be continuous, a scaling factor for the ionization tail is included

$$S(\boldsymbol{\pi}_{\text{eloss}}) = f_{\text{exit}}^{\text{emp}}(E_{\text{ion,D}_2} \mid \boldsymbol{\pi}_{\text{eloss}}) / f_{\text{ion}}^{\text{BED}}(E_{\text{ion,D}_2}). \quad (6.4)$$



Then, the full parametrization for the KATRIN model reads [Pre-Han19-1]

$$f_1^{\text{KATRIN}}(\epsilon | \boldsymbol{\pi}_{\text{eloss}}) = \begin{cases} 0 & \text{if } \epsilon < 0 \\ f_{\text{exit}}^{\text{emp}}(\epsilon | \boldsymbol{\pi}_{\text{eloss}}) & \text{if } 0 \leq \epsilon \leq E_{\text{ion,D}_2} \\ S(\boldsymbol{\pi}_{\text{eloss}}) \cdot f_{\text{ion}}^{\text{BED}}(\epsilon) & \text{if } \epsilon > E_{\text{ion,D}_2} \end{cases} . \quad (6.5)$$

Figure 6.1 shows the KATRIN model in comparison to the Aseev model. They are not expected to be fully compatible within their uncertainties as they describe the scattering off two different hydrogen isotopologues. For a comparison of the presented KATRIN model with an energy loss model for deuterium from literature, the reader is referred to [Pre-Rod19]. Two of the three Gaussian distributions in the excitation peak region overlap in a way, that they are not distinguishable in the given plot scale. However, one Gaussian forms a second visible, smaller peak in the excitation region which is a feature that is not present in the Aseev model. As the KATRIN model is still in its early stages, this work refrains from a detailed physical interpretation and instead focuses on uncertainty propagation. For a more detailed physical interpretation, the reader is referred to the KATRIN documents [Pre-Rod19; Pre-Han19-1; Pre-Han19-2]. With respect to the uncertainties, the KATRIN model shows an improved uncertainty in the ionization tail and also in large parts of the excitation peak compared to the Aseev model. How this propagates within neutrino mass inference is investigated in this chapter.

### 6.1.2. Nuisance Parameters that Further Influence the KATRIN Energy Loss Model

The uncertainties of the KATRIN model as evaluated at the time of writing this thesis can be divided into two sets: The KATRIN model itself comprises nine parameters (eq. 6.1). Moreover, the model was obtained in a 15-parameter fit. The other six fit parameters are correlated with the parameters of the KATRIN model and hence are not necessarily negligible with regard to uncertainty treatment. In the scope of this thesis, they were incorporated in the statistical treatment of the uncertainties. Whether they may be negligible or can be treated in a simpler fashion in future studies needs further investigation.

For a the full fitting procedure of the KATRIN model to the recorded data, the reader is referred to [Pre-Han19-1], but a description of the further six fit parameters is given in the following. In order to outline their meaning, an overview of the used data sets is required. Four integral spectra were recorded analogously to the described integral  $\beta$  spectrum in chapter 3, but with the electron gun (see section 2.3) as electron source instead of tritium. For each data set a different deuterium column density (see section 2.2) was set in the WGTS: 0 %, 15 %, 50 %, 100 % of the nominal column density of  $\rho d = 5 \times 10^{17}$  molecules/cm<sup>2</sup>. The fit model for the 15 %-measurement was rescaled with respect to the other measurements because the respective data set underwent a different preprocessing than the others. The corresponding scaling factor  $N_{\text{int},15}$  was a free fit parameter. A further scaling factor  $N_{\text{K}}$  for the KATRIN model was introduced in the fit as a free parameter in order for the whole KATRIN model to be normalized to unity. Additionally, a fifth data set recorded at 15 % column density in a time-of-flight mode (see [Bon+99]) was fitted simultaneously. For each data set, a different expected scattering count  $\mu_{\text{tof},15}, \mu_{\text{int},15}, \mu_{\text{int},50}, \mu_{\text{int},100}$  (see eq. 3.17) was fitted, which adds four further parameters (there is no scattering for the 0 % measurement). In summary, the additional parameters in the fit of the KATRIN model are

$$\boldsymbol{\pi}_{\text{eloss}+} = (N_{\text{K}}, \mu_{\text{tof},15}, \mu_{\text{int},15}, \mu_{\text{int},50}, \mu_{\text{int},100}, N_{\text{int},15})^{\text{T}} . \quad (6.6)$$

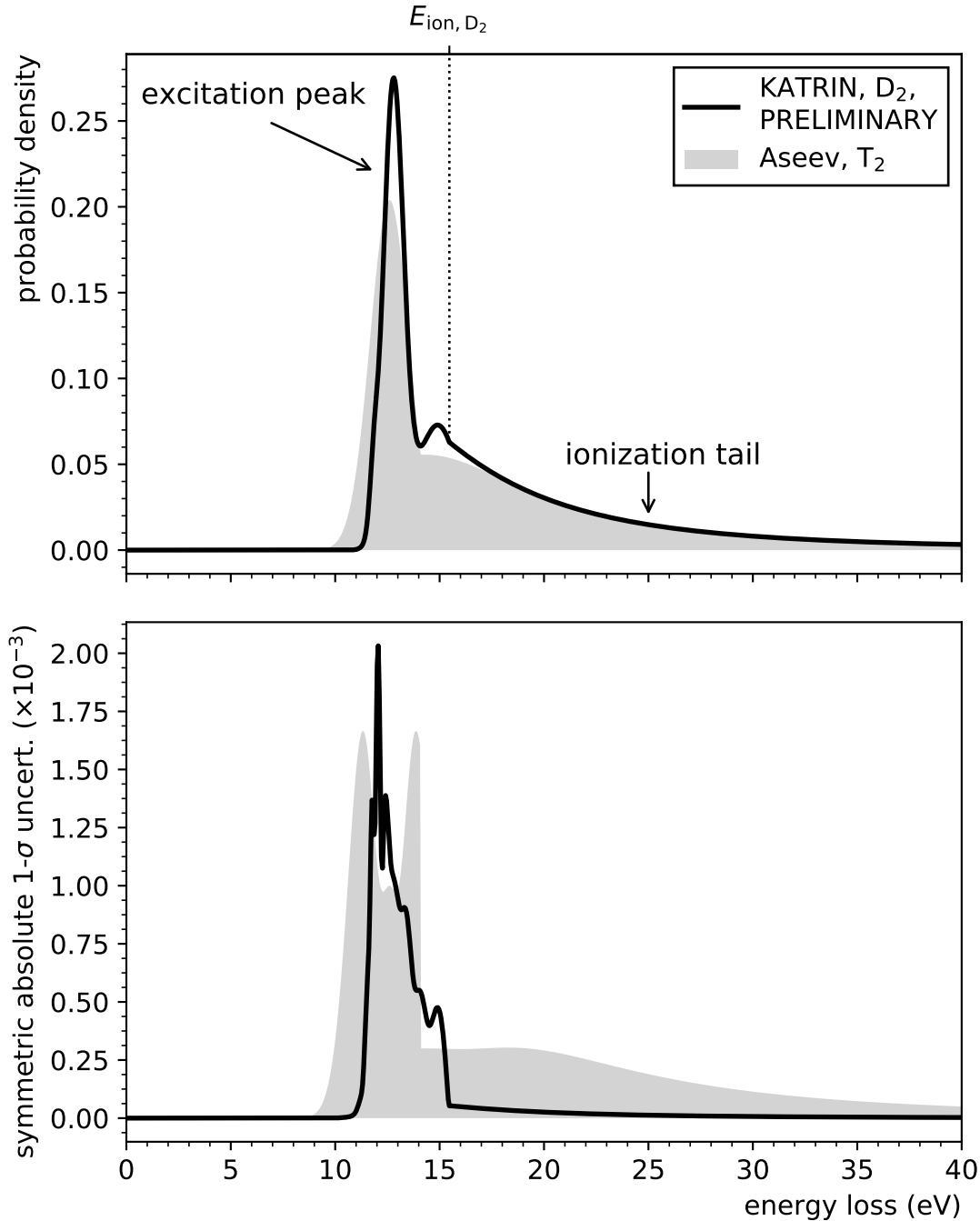
The MLEs, the estimated standard deviations and the estimated correlation matrix of the full 15-parameter set ( $\boldsymbol{\pi}_{\text{eloss}}$  and  $\boldsymbol{\pi}_{\text{eloss}+}$ ) can be found in appendix D. The aim of this chapter is to study the impact of the estimated uncertainties from these 15 parameters on KATRIN's sensitivity to the neutrino mass.

## 6.2. Scope of the Conducted Study

This section lists considerations that should be kept in mind with regard to the study presented in this chapter.

During a neutrino mass measurement, the nominal tritium purity in the WGTS is planned to be above 95% [KAT05]. In this chapter, a sensitivity study is presented that uses an energy loss model for scattering off deuterium molecules. Within the scope of this thesis, the treatment of the uncertainties of the KATRIN energy loss model in neutrino mass inference is of primary interest. Two different, but similar energy loss models are expected for the two different gas species. As long as the KATRIN model for scattering off deuterium molecules is sufficiently similar to a model for scattering off tritium molecules, it is a plausible approach to use it for simulations of KATRIN neutrino mass measurements (for the similarity, see figure 6.1). Also, it is plausible to assume that the parameters for the KATRIN energy loss model for scattering off tritium molecules can be measured with a similar precision to the precision for scattering of deuterium molecules. In that regard, the results for KATRIN's sensitivity derived in this chapter should yield practical indications on what to expect, once the KATRIN energy loss model is available in its final version and for scattering off tritium molecules.

Additionally, it should be noted that the KATRIN model as presented is still preliminary and may be subject to change because the analysis of recent measurements is still ongoing. Furthermore, the KATRIN model does not yet incorporate all known systematic uncertainties. Corresponding efforts for their incorporation are made at the time of writing this thesis. In that regard, a future version of the KATRIN model may exhibit larger uncertainties than the version used in this thesis. But at the same time a continuous modeling effort and an improved understanding of the data may also cause reduced uncertainties. Nonetheless, once the KATRIN model is completed, the study presented in this chapter is expected to be easily repeatable because the necessary statistical tools are now readily implemented.



**Figure 6.1.:** The preliminary empirical KATRIN energy loss model for electrons scattering off deuterium molecules. The black line shows the KATRIN energy loss model as established by a designated subgroup of the KATRIN collaboration for electrons scattering off deuterium molecules (KATRIN model). The corresponding model for scattering off tritium molecules as established by [Ase+00] (Aseev model) is shown for comparison as a shaded area. (That the latter is plotted as an area instead of as a line solely serves readability as the two functions overlap strongly.) The top panel shows the probability densities of the energy loss. The transition from the excitation peak to the ionization tail is marked by the ionization energy of deuterium at  $E_{\text{ion}, D_2} = 15.467$  eV. The bottom panel shows the absolute symmetric 1- $\sigma$  uncertainties of the models. The uncertainties were obtained through uncertainty propagation via derivatives from the uncertainties of the model parameters (see figure 3.3 for the parameter values of the Aseev model and appendix D for the KATRIN model). Correlations are respected for the KATRIN model. However, there are no published correlations for the Aseev model. The KATRIN model shows particularly improved uncertainties in the ionization tail region in comparison to the Aseev model. (KATRIN model adapted from [Pre-Han19-1].)

### 6.3. Statistical Prerequisites

This section develops the statistical tools used in the scope of this chapter in order to evaluate the impact of the KATRIN energy loss model on KATRIN's sensitivity to the neutrino mass. The methods are described in a general manner (and could be applied to study model uncertainties in general) and then related to the KATRIN energy loss model in each section separately.

#### Concept and Notation

The general concept of the study presented in this chapter is the following: The parameters of the KATRIN energy loss model,  $\pi_{\text{eloss}}$  and  $\pi_{\text{eloss+}}$  (see eq. 6.1 and 6.6), are treated as free parameters in a neutrino mass fit. This makes the model more general than fixing these parameters to their best estimates. Such an approach, the inclusion of further free parameters in the model of a measurement, may eliminate or at least reduce the effect of a possible systematic bias in parameter inference, but their presence will also result in an increased statistical uncertainty on the parameter of interest [Tan+18], here, the squared neutrino mass. In that regard, the following notation is used for the uncertainties of a parameter of interest within this chapter:

- $\sigma_{\text{stat}}$  denotes the statistical uncertainty without using additional free fit parameters beyond the set of a nominal four-parameter KATRIN neutrino mass fit (see section 4.3),
- $\sigma_{\text{tot}}$  denotes the statistical uncertainty in the presence of further free fit parameters and
- $\Delta\sigma_{\text{stat}} = \sigma_{\text{tot}} - \sigma_{\text{stat}}$  denotes the difference.

As an aside, the mitigation of a systematic bias by additional fit parameters has already been demonstrated in another context in section 5.3.3, where the parameter for the signal amplitude  $A_S$  could partly compensate for a constant offset of the scattering cross section.

#### Outline of the Statistical Prerequisites

In the study presented in this chapter, the parameters of the KATRIN energy loss model are treated as free fit parameters. A KATRIN neutrino mass measurement is however hardly sensitive to these parameters. Without further measures, one would obtain a greatly enlarged statistical uncertainty because the likelihood is relatively flat in the dimension of these additional parameters. In contrary, the measurement of the KATRIN energy loss was especially designed to be sensitive to its parameters. In other words, there are data that can be used to constrain the parameters in neutrino mass inference. Section 6.3.1 presents a corresponding concept for the combination of a neutrino mass and a calibration measurement.

Treating the parameters of the KATRIN energy loss model as free leads to the presence of additional nuisance parameters. Section 6.3.2 explains the profile-likelihood method for the treatment of such nuisance parameters when determining a one-dimensional confidence interval (as required for the squared neutrino mass).

In order to incorporate the statistical nature of the detector counts measured by the KATRIN detector into neutrino mass inference, ensemble tests can be applied (see section 4.6.2). Ensemble testing may be computational demanding. Section 6.3.3 introduces the idea of a so-called Asimov data set that can represent an ensemble of simulated data sets and hence simplify statistical studies.

In the scope of this chapter, the combination of the statistical tools mentioned above was applied in order to study how the uncertainties of the KATRIN energy loss model inflict KATRIN's sensitivity to the neutrino mass.

### 6.3.1. Combination of a Calibration and a Neutrino Mass Measurement

If two models share a set of parameters  $\boldsymbol{\theta}_s$  and, additionally, have an individual set of parameters  $\boldsymbol{\theta}_1$  and  $\boldsymbol{\theta}_2$  and for each model exists a statistically independent set of observations, then a combined likelihood  $L$  is given by the product of the likelihoods  $L_1$  and  $L_2$  of each measurement [Tan+18]

$$\begin{aligned} -2 \ln L(\boldsymbol{\theta}_s, \boldsymbol{\theta}_1, \boldsymbol{\theta}_2) &= -2 \ln L_1(\boldsymbol{\theta}_s, \boldsymbol{\theta}_1) - 2 \ln L_2(\boldsymbol{\theta}_s, \boldsymbol{\theta}_2) \\ &\equiv -2 \ln L_1(\boldsymbol{\theta}_{s,1}) - 2 \ln L_2(\boldsymbol{\theta}_{s,2}), \end{aligned} \quad (6.7)$$

where, for ease of notation, the combined parameter vectors  $\boldsymbol{\theta}_{s,1} \equiv (\boldsymbol{\theta}_s, \boldsymbol{\theta}_1)$  and  $\boldsymbol{\theta}_{s,2} \equiv (\boldsymbol{\theta}_s, \boldsymbol{\theta}_2)$  are introduced. In the scope of this chapter, it makes sense to identify the first measurement with a KATRIN neutrino mass measurement and the second with the measurement of the KATRIN energy loss model (or the Aseev model). In order to emphasize generality, this identification is postponed until the end of this section and the second measurement is called a calibration measurement throughout this section.

#### Approximated Combination of Likelihoods

For practicality, in this thesis, an approximation is applied: The calibration measurement is treated as evaluated independently (which is the case for the evaluation of the measurement of the KATRIN energy loss model). Hence, there are estimates  $\hat{\boldsymbol{\theta}}_{s,2}$ , and an estimated covariance matrix  $\hat{V}_{s,2}$  for the parameters of the calibration measurement (in this chapter, variables with a hat mark an estimator). These can in turn be used to approximate the likelihood  $L_2$ . A choice that stands to reason for the approximation of  $L_2$  is a multivariate normal distribution  $\mathcal{N}$ . For the purpose of parameter inference,  $-2 \ln L_2$  needs only to be accurately approximated within the contour that is needed to extract confidence intervals. The choice of a multivariate normal distribution corresponds a symmetric approximation in second order of  $-\ln L_2$  around its minimum. The combined likelihood (6.7) then reads

$$\begin{aligned} -2 \ln L(\boldsymbol{\theta}_s, \boldsymbol{\theta}_1, \boldsymbol{\theta}_2) &\approx \chi^2(\boldsymbol{\theta}_{s,1}) - 2 \ln \mathcal{N}(\boldsymbol{\theta}_{s,2}, \hat{\boldsymbol{\theta}}_{s,2}, \hat{V}_{s,2}) + \text{constants} \\ &= \underbrace{\chi^2(\boldsymbol{\theta}_{s,1})}_{(1)} + \underbrace{(\boldsymbol{\theta}_{s,2} - \hat{\boldsymbol{\theta}}_{s,2})^\top \hat{V}_{s,2}^{-1} (\boldsymbol{\theta}_{s,2} - \hat{\boldsymbol{\theta}}_{s,2})}_{(2)} + \text{constants} \end{aligned} \quad (6.8)$$

Here, (1) is the chi-square likelihood for a KATRIN neutrino mass measurement (see eq. 4.4). And (2) resembles the negative log-likelihood of the calibration measurement approximated by a multivariate normal distribution. Terms having a form like (2) may be called “pull terms” because in the minimization of the likelihood they “pull” the parameters  $\boldsymbol{\theta}_{s,2}$  towards the corresponding values in  $\hat{\boldsymbol{\theta}}_{s,2}$ .

#### Chi-Square Characteristics

The chi-square term (1) in equation (6.8) is assumed to be a sum of standard normal distributed random variables (as discussed in section 4.2). Hence, a likelihood only composed of the chi-square term (1) offers a goodness-of-fit criteria via the Pearson chi-square statistic. Whether the same criteria can be applied to the combined likelihood has to be investigated individually from case to case. A corresponding study with regard to the KATRIN energy loss model is presented in the subsequent section 6.4.2.

#### Implementation in the KaFit Software Framework

The KaFit software module (see section 4.5) had allowed to use one-dimensional Gaussian “pull terms” (term (2) in eq. 6.8). In the scope of this thesis, the software was extended to allow for arbitrary dimensions with corresponding correlations by using a multivariate

normal distribution. Albeit not of particular interest in this chapter, for completeness, the following shall be mentioned: By comparing equations (4.6) and (6.8) it becomes apparent that such “pull terms” take the same mathematical form as Bayesian priors. For that reason, further term forms apart from the multivariate normal distribution were implemented in order to be used as priors in a Bayesian analysis. For a documentation of the software features see appendix C.

### Application to the Energy Loss Model

With regard to the study presented in this chapter, the following identification can be made:  $\theta_1$  comprises the parameter of a nominal four-parameter KATRIN neutrino mass fit (see section 4.3). Furthermore, the calibration measurement can be identified with the measurement of the KATRIN energy loss model  $\theta_s = \pi_{\text{eloss}}$ ,  $\theta_2 = \pi_{\text{eloss+}}$ ,  $\hat{\theta}_s = \hat{\pi}_{\text{eloss}}$ ,  $\hat{\theta}_2 = \hat{\pi}_{\text{eloss+}}$  (see equations 6.1 and 6.6 for  $\pi_{\text{eloss}}$  and  $\pi_{\text{eloss+}}$ ) and the corresponding estimator for the covariance matrix  $\hat{V}_{s,2} = \hat{V}_{\text{eloss,eloss+}}$ . The numerical values for the three estimators can be found in the form of the means, the standard deviations and the correlation matrix in appendix D. In the same manner, the calibration measurement can instead be identified with the measurement of the Aseev model  $\theta_s = (A_1, A_2, \epsilon_2, \omega_1, \omega_2)^\top$  without additional nuisance parameters  $\dim(\theta_2) = 0$  and a diagonal, estimated covariance matrix  $\hat{V}_{s,2} \equiv \hat{V}_{\text{Aseev}}$  (diagonal because there are no published correlations of the parameters of the Aseev model). See equation (3.21) for the meaning of the parameters of the Aseev model and see the caption of figure 3.3 for their values.

### 6.3.2. Nuisance Parameters and the Profile-Likelihood Method

Apart from the parameters of interest  $\theta$ , the KATRIN likelihood can depend on further nuisance parameters  $\pi$ . With regard to the study presented in this chapter, the parameter of interest is the squared neutrino mass  $\theta \equiv m_\nu^2$  and the nuisance parameters are the further three parameters of a nominal four-parameter KATRIN neutrino mass fit as well as the parameters of the energy loss model (KATRIN or Aseev). The derivation of a combined confidence region for the full parameter set may cause long run times (or be impossible) due to the dimensionality (here: 15 parameters of the KATRIN model and additional four parameters of a nominal KATRIN neutrino mass fit). And, more importantly, as indicated by the naming conventions, only a confidence interval for the squared neutrino mass is of interest. The aim of KATRIN is the deduction of a stand-alone confidence interval for the neutrino mass as opposed to an interval that depends on many other nuisance parameters. This section outlines the profile-likelihood method, that effectively handles nuisance parameters in the derivation of a one-dimensional confidence interval.

### Confidence Regions for the Parameters Of Interest Only

Section 4.4.2 already introduced a method for deriving confidence regions for the full parameter set that the likelihood depends on (parameters of interest and nuisance parameters) based on the likelihood ratio as a test statistic. It can be used to reject estimators for the combined parameter set with a significance of  $\alpha$  and include all non-rejected estimators in a confidence region of confidence level  $\alpha$ . The same approach may also be used to construct a confidence region only for the parameters of interest. But, in the strict frequentist approach, values for the parameters of interest are excluded from the confidence region only if they are rejected for all possible values of the nuisance parameters [Tan+18]. To show the rejection for all possible values of the nuisance parameters may be difficult [Tan+18] (corresponding studies were not pursued in the scope of this thesis). The difficulty may be circumvented if a test statistic is found that solely depends on the parameters of interest. Exact independence from the nuisance parameters is only achievable in special cases, it can, however, be achieved approximately by the use of the profile-likelihood ratio as test statistic [Tan+18].

### The Profile-Likelihood Ratio

In the following, the profile likelihood is defined. It only depends on the parameters of interest  $\boldsymbol{\theta}$  and is independent of the nuisance parameters  $\boldsymbol{\pi}$ . Its values correspond to the likelihood values evaluated at  $\boldsymbol{\theta}$  in the dimensions of the parameters of interest and maximized in the dimensions of the nuisance parameters [Tan+18]

$$L_p(\boldsymbol{\theta}) = L(\boldsymbol{\theta}, \hat{\boldsymbol{\pi}}(\boldsymbol{\theta})), \quad (6.9)$$

where the double-hat indicates the maximization respectively the “profiling”. Also, the profile-likelihood ratio can be defined [Tan+18]

$$\lambda_p(\boldsymbol{\theta}) = \frac{L_p(\boldsymbol{\theta})}{L_p(\hat{\boldsymbol{\theta}})}. \quad (6.10)$$

According to Wilks’ theorem [Wil38], the distribution of  $-2 \ln \lambda_p(\hat{\boldsymbol{\theta}})$ , where  $\hat{\boldsymbol{\theta}}$  is the MLE (see section 4.1), approaches a chi-square distribution in the limit of a large data sample, independently of the values of the nuisance parameters  $\boldsymbol{\pi}$  [Tan+18]. Hence, the profile-likelihood ratio offers a test statistic, based on which, values for the parameters of interest can be rejected independently of the values of the nuisance parameters. In other words, the profile-likelihood method is a constructive approach on how to derive a confidence region for the parameters of interest.

### 6.3.3. An Asimov Data Set in Relation to Ensemble Tests in Sensitivity Studies

If one were to repeat the KATRIN experiment many times, one would obtain an ensemble of confidence intervals for the neutrino mass (see section 4.4.2 about confidence intervals). KATRIN’s sensitivity can be deduced from a confidence interval (see section 4.6.1 about KATRIN’s sensitivity). In that sense, if many KATRIN measurements are simulated, one also obtains an ensemble respectively a distribution of sensitivities. It should be noted, that such a distribution of sensitivities is narrow compared to its expectation value (see for example the subsequent eq. 6.12 where the spread of the statistical uncertainty (68 % C.L.) on the squared neutrino mass is smaller than the mean value by two orders of magnitude). In that sense, the whole distribution of sensitivities is not necessarily of interest, but the mean or the median sensitivity can provide sufficient information. One way to retrieve the mean or the median sensitivity is to simulate many KATRIN neutrino mass measurements and fluctuate the measured electron counts according to Poissonian statistics (or Gaussian statistics as an approximation) and extract them from the obtained distribution. The simulation of many measurements was for example applied within the scope of the KATRIN Design Report [KAT05].

However, simulating a KATRIN neutrino mass measurement can be time-consuming depending on its level of detail. In that regard, using only one simulation would be more practical. The median of KATRIN’s sensitivity may be retrieved from one simulation by replacing the electron count rates with their expectation value instead of fluctuating them according to Poissonian statistics [Cow+11]. This reasoning is based on Walt’s theorem [Wal44] that concerns the sum of Gaussian distributed random variables in the limit of many summands (which approximately conforms the KATRIN-chi-square likelihood). Such a single simulated data set is called an Asimov<sup>1</sup> data set. However, whether this approach is applicable may be hard to verify without doing an ensemble test, which would nullify the practicality here.

---

<sup>1</sup>The name of the Asimov data set is inspired by the short story “Franchise”, by Isaac Asimov. In the story, elections are held by selecting the single most representative voter to replace the entire electorate [Cow+11].

In the scope of this chapter, both approaches, using an Asimov data set and simulating an ensemble, were applied. The outcome verifies that an Asimov data set would be representative in the study presented in this chapter and indeed yields the median sensitivity.

#### 6.4. Free Fit Parameters of the Energy Loss Model in Neutrino Mass Inference

This section lists the increase of the statistical uncertainty by additional fit parameters using the Aseev or the KATRIN energy loss model, respectively. The sensitivity study was conducted using a mathematical model of a KATRIN neutrino mass measurement as described in chapter 3 and the Aseev model as described in section 3.2.5 using a nominal configuration based on the KATRIN Design Report [KAT05] with a squared neutrino mass of  $0 \text{ eV}^2$  and a measurement time of three years. Correspondingly, the SSC and KaFit modules were used (see section 4.5). The full configuration of the study can be found in appendix E.

The confidence interval for the squared neutrino mass was extracted via the profile-likelihood method as described in section 6.3.2.

The following three cases were investigated using an Asimov data set and the second case also via ensemble testing:

1. The simulation- and fit-model use the KATRIN energy loss model, but it was assumed to be without uncertainties. In other words, only the four parameters of a nominal KATRIN-neutrino-mass fit (see section 4.3) were treated as free parameters and the parameters of the KATRIN energy loss model were fixed to their best estimates. In the following, this case will be denoted the “nominal case”.
2. The simulation- and fit-model use the KATRIN energy loss model, and all its parameters were treated as free in order to incorporate the uncertainties of parameters of the KATRIN model, which results in a 19-parameter fit.
3. The simulation- and fit-model use the Aseev model, and all its parameters were treated as free in order to incorporate the uncertainties of the parameters of the Aseev model, which results in a nine-parameter fit.

For the last two cases, the parameters of the energy loss models were constrained via “pull terms” as described in section 6.3.1 about the combination of a neutrino mass and a calibration measurement.

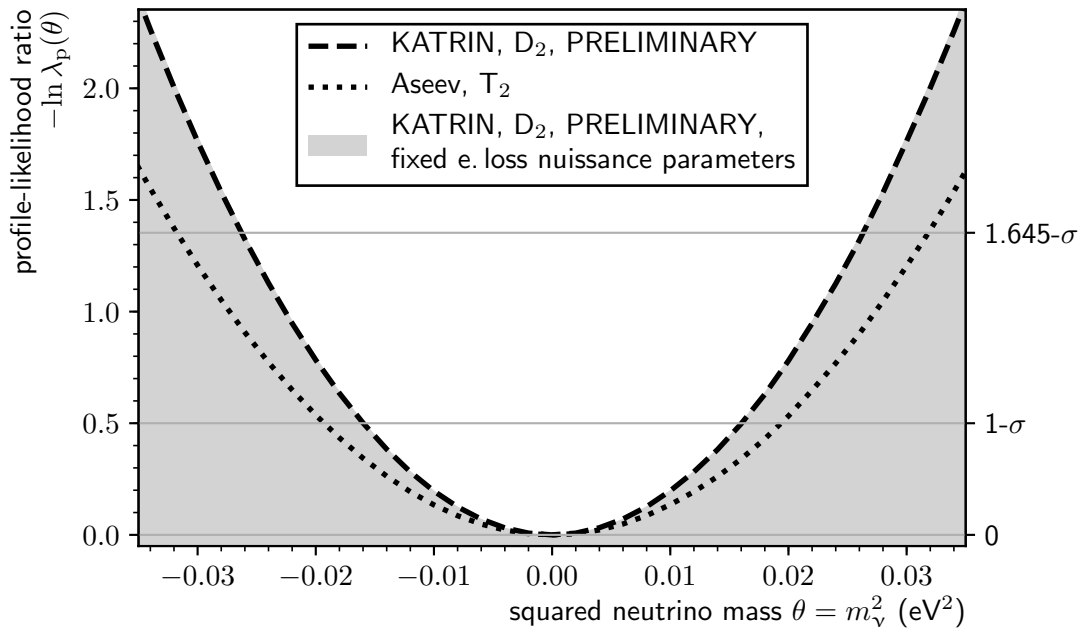
Section 6.4.1 presents the results using an Asimov data set. And section 6.4.2 presents the same results as obtained by an ensemble test for the second of the cases listed above.

##### 6.4.1. Sensitivity from an Asimov Data Set

Figure 6.2 shows the profile-likelihood ratios and table 6.1 lists the extracted confidence intervals and obtained sensitivities for the three cases mentioned in the introduction to this section.

The result of the nominal four-parameter KATRIN neutrino mass fit should be comparable to the results of former works listed in table 4.1. Here, a  $\sim 4 \times 10^{-4} \text{ eV}^2$  smaller statistical uncertainty on the squared neutrino mass is obtained compared to the results by [Kle14; Höt12]. This is mainly due to the fact, that in the study presented in this chapter a more up-to-date [Ams+15] detection efficiency of 95 % was used, whereas the other works used 90 %.





**Figure 6.2.: Profile-likelihood ratio of a KATRIN measurement from an Asimov data set.** The graph shows the profile-likelihood ratio  $\lambda_p$  incorporating the uncertainties of the following three cases: 1) KATRIN model (dashed line), 2) Aseev model (dotted line) or 3) nominal case (shaded area). For a description of the different cases, the reader is referred to the main text. (That the nominal case is plotted as an area instead of as a line solely serves readability because a line would overlap with the line for the case of the KATRIN model.) The MLE recovers the true simulated squared neutrino mass of  $0 \text{ eV}^2$  with a corresponding likelihood ratio (and profile likelihood ratio) of 1 ( $\Rightarrow -\ln \lambda_p(\hat{\theta}) = 0$ ). The horizontal  $s$ - $\sigma$  lines are drawn at  $s^2/2$  as per equation (4.9). Their intersections with  $-\ln \lambda_p$  mark the confidence intervals for the neutrino mass at 68% respectively 90% confidence level. The width of the two intervals indeed relate through the factor 1.645 due to the parabolic shape of  $-\ln \lambda_p$ . The uncertainties of the KATRIN model hardly have an impact compared to the nominal case because the corresponding profile likelihoods are almost equal (also see table 6.1 for quantitative values).

Also, the tabulated values show, that the extrapolation of the likelihood to negative squared neutrino masses as discussed in section 4.4.2 yields a slightly asymmetric (order of  $\sim 10^{-4}$ ) profile likelihood for the nominal case and the case that treats the parameters of the KATRIN model as free. This might mean, that a perfectly symmetric extrapolation might need fine-tuning of the corresponding parameters on a case-by-case basis.

Under the restrictions listed in section 6.2, the following conclusions can be drawn:

- Treating the parameters of the KATRIN model as free parameters and constraining them as described in section 6.3.1 about the combination of a calibration and neutrino mass measurement results in an enlarged statistical uncertainty (68% C.L.) for the squared neutrino mass

$$\Delta\sigma_{\text{stat}}(m_\nu^2) = 1.2 \times 10^{-4} \text{ eV}^2. \quad (6.11)$$

as compared to a nominal four-parameter KATRIN neutrino mass fit. This enlargement is negligible with respect to the KATRIN systematic budget as it would change KATRIN's sensitivity by less than 1 meV.

**Table 6.1.: Neutrino mass confidence intervals and sensitivities obtained from an Asimov data set.** The table lists values that can be extracted from the profile-likelihood ratio depicted in figure 6.2 for the three conducted studies denoted in the first column (also refer to the main text for a description of these three cases). The following quantities are reported: the lower bound of the confidence interval (68 % C.L.) on the squared neutrino mass  $l(m_\nu^2)$ , the upper bound  $u(m_\nu^2)$ , half the width of the interval  $\sigma_{\text{stat}}(m_\nu^2)$  or  $\sigma_{\text{tot}}(m_\nu^2)$  and KATRIN’s sensitivity on the neutrino mass as per equation (4.11). For the calculation of the sensitivity an additional systematic budget of  $0.017 \text{ eV}^2$  was included (see section 4.6.2).

|                 | $l(m_\nu^2)$<br>( $10^{-2} \text{ eV}^2$ ) | $u(m_\nu^2)$<br>( $10^{-2} \text{ eV}^2$ ) |   | $S_{m_\nu}$ (90 %)<br>(meV) |
|-----------------|--|--|---|-----------------------------|
| 1. nominal case | -1.586                                     | 1.592                                      | $\sigma_{\text{stat}}(m_\nu^2) = 1.589 \times 10^{-2} \text{ eV}^2$ | 196                         |
| 2. KATRIN model | -1.598                                     | 1.604                                      | $\sigma_{\text{tot}}(m_\nu^2) = 1.601 \times 10^{-2} \text{ eV}^2$  | 196                         |
| 3. Aseev model  | -1.931                                     | 1.939                                      | $\sigma_{\text{tot}}(m_\nu^2) = 1.935 \times 10^{-2} \text{ eV}^2$  | 206                         |

- Using the KATRIN model at its current stage yields an improvement of 10 meV in sensitivity to the neutrino mass compared to using the Aseev model.
- The logarithm of the profile-likelihood ratio has a parabolic shape, which translates to a Gaussian shape for the likelihood. This justifies the usage of the factor 1.645 to convert a confidence interval of 68 % confidence level into one of 90 %.

#### 6.4.2. Cross-Check and Extension of the Asimov Data Set via Ensemble Testing

An ensemble of 4046 KATRIN neutrino mass measurements with a true neutrino mass of  $0 \text{ eV}$  was simulated using the KATRIN energy loss model and incorporating its uncertainties in the same manner as for the Asimov data set described in the last section 6.3.3. However, as opposed to the Asimov data set, in the simulation of the ensemble, the detector counts were fluctuated according to Poissonian statistics. Several aspects were investigated as listed below:

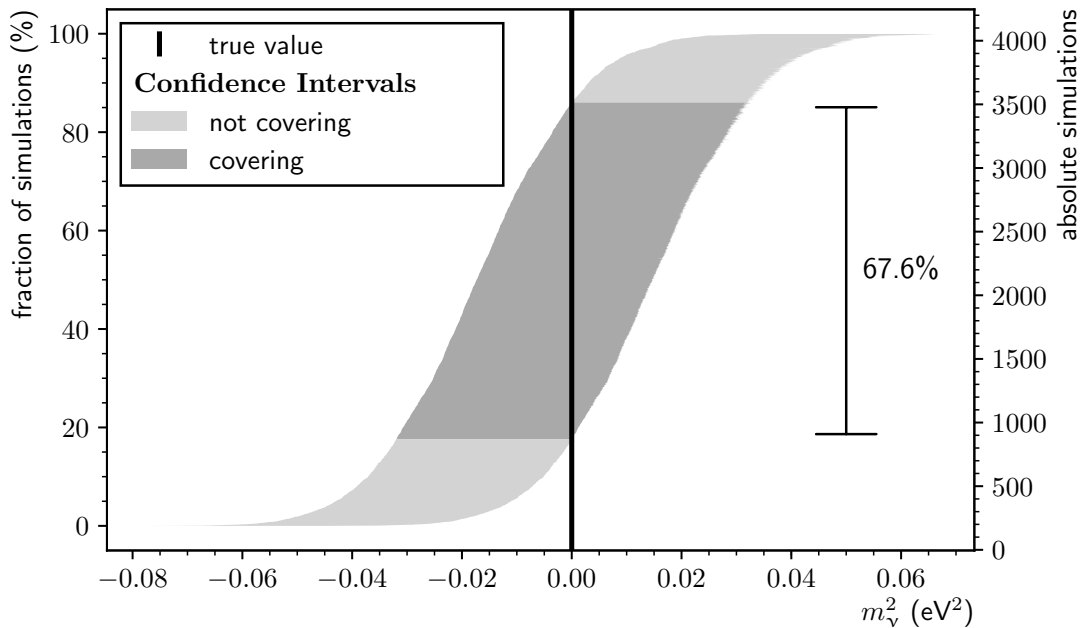
##### Test of Coverage

The extraction of a confidence interval via the profile-likelihood method as done in the previous section 6.3.3 for the Asimov data set should per construction yield a coverage probability of 68.2 %. If all conditions are met for this to hold was tested. In other words, 68.2 % of the obtained confidence intervals in the conducted ensemble test should cover the true simulated squared neutrino mass of  $0 \text{ eV}^2$ . The obtained coverage is 67.6 % as illustrated in figure 6.3. The slight undercoverage on the  $10^{-3}$  scale may stem from a limited ensemble test size. Furthermore, it must be emphasized that here the likelihood of the measurement of the KATRIN energy loss model is approximated by a multivariate normal distribution and does not fluctuate in the presented ensemble test. For the result to have full validity, the measurement of the KATRIN energy loss model has to be simulated along with the KATRIN neutrino mass measurement. This may be the aim of a future analysis.

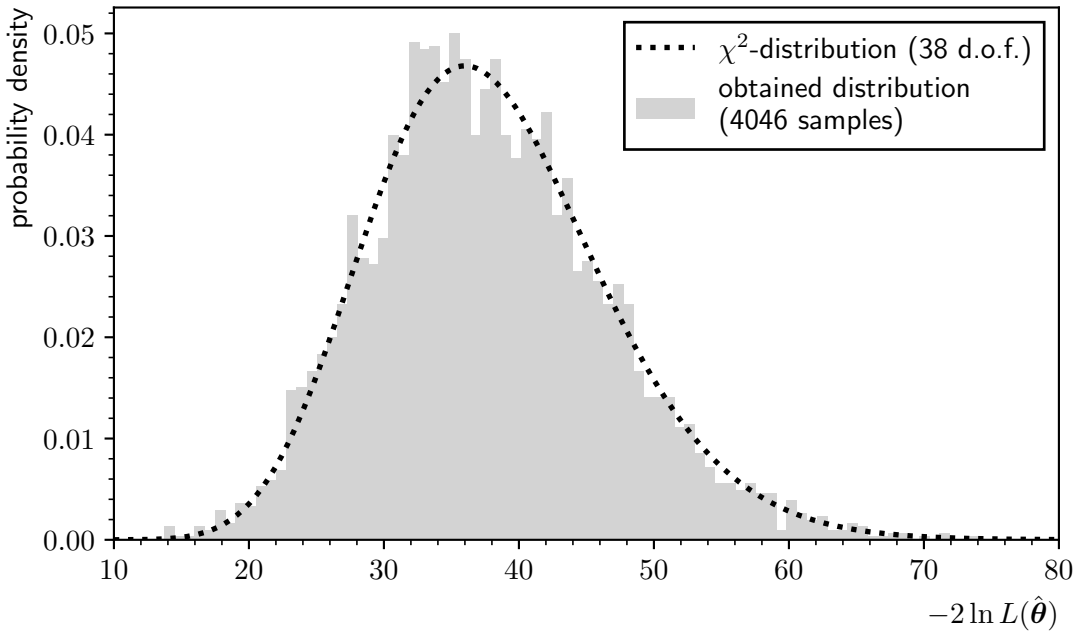
##### Chi-Square Characteristics

The combined likelihood of a neutrino mass and a calibration measurement evaluated at the MLE  $-2 \ln L(\hat{\theta})$  might not follow the chi-square statistic as mentioned in section 6.3.1. Figure 6.4 shows the obtained distribution of  $-2 \ln L(\hat{\theta})$  in the conducted ensemble test. An MTD with 41 retarding potentials was used. Hence, there are 41 summands in the

likelihood for the KATRIN neutrino mass measurement. In the combined likelihood, there are additional 15 summands to approximate the likelihood of the measurement of the KATRIN energy loss model. In total, there are 19 fit parameters, four for the nominal KATRIN neutrino mass fit (see section 4.3) and 15 in order to incorporate the uncertainties of the KATRIN model. Thus, the hypothesis stands to reason that the obtained distribution follows a chi-square distribution with  $41 + 15 - 19 = 37$  degrees of freedom. A corresponding Kolmogorov–Smirnov test yields a  $p$ -value of  $p = 6 \times 10^{-6}$ . In other words, this hypothesis has to be rejected with a significance of  $4\sigma$ . Repeating the test for 38 respectively 39 degrees of freedom yields  $p = 0.14$  respectively  $p = 1 \times 10^{-14}$ . Hence, a chi-square distribution of 38 degrees of freedom may not be rejected. This is important because it means that the chi-square statistic can not necessarily be used as a measure for goodness-of-fit when incorporating the uncertainties of the KATRIN energy loss model into the neutrino mass inference in the way described in this thesis. Or, at least, one has to be careful about the choice of degrees of freedom. The origin of this effect is that, here, the likelihood of the measurement of the KATRIN energy loss model is approximated by a multivariate normal distribution and hence does not fluctuate in the given ensemble test.



**Figure 6.3.:** Test of coverage for the ensemble of confidence intervals obtained in the sensitivity study using the KATRIN energy loss model. The graph illustrates the coverage probability. An ensemble of 4046 KATRIN neutrino mass measurements with a true squared neutrino mass of  $0 \text{ eV}^2$  was simulated. For each simulated measurement a confidence interval was constructed using the profile-likelihood method. The obtained confidence intervals were sorted by their lower limit and plotted stacked which yields the gray band. The confidence intervals that cover the true value are depicted in dark gray, while the ones that do not cover the true value are depicted in light gray. In total 67.6% coverage is obtained. In the limit of an infinite ensemble size, a coverage of 68.3% would be expected per construction (also refer to the main text for a limitation of this statement). It should also be noted, that there is little fluctuation in the width of the confidence intervals, which hints at the representative qualities of an Asimov data set.



**Figure 6.4.:** Chi-square distribution for the simulated ensemble of neutrino mass measurements obtained in the sensitivity study using the KATRIN energy loss model. An ensemble of 4046 KATRIN neutrino mass measurements was simulated. The histogram shows the obtained distribution of the likelihood  $L$  evaluated at the MLE  $\hat{\theta}$ . The obtained distribution may follow a chi-square distribution with 38 degrees of freedom. For details the reader is referred to the main text.

### Representative Qualities of the Asimov Data Set

The estimated mean and standard deviation of the distribution of the confidence intervals (68% C.L.) as obtained by the ensemble test is

$$\hat{\sigma}_{\text{tot}}(m_\nu^2) = (1.599 \pm 0.013) \times 10^{-2} \text{ eV}^2 \quad (6.12)$$

in agreement with the one obtained through the Asimov data set in table 6.1. Furthermore, the median confidence interval

$$\tilde{\sigma}_{\text{tot}}(m_\nu^2) = 1.598 \times 10^{-2} \text{ eV}^2 \quad (6.13)$$

recovers the one of the Asimov data set on the  $10^{-5} \text{ eV}^2$  level. This verifies the Asimov data set as representative for the study on the KATRIN energy loss model presented in this chapter.

#### 6.4.3. Reduction of Systematic Effects by Additional Fit Parameters

The motivation for the inclusion of further free fit parameters is their capability to compensate for systematic shifts. In order to check whether this has a significant effect with regard to the KATRIN energy loss model, the following study was conducted: A KATRIN neutrino mass measurement was simulated using the KATRIN energy loss model as in the previous sections (see appendix E for the full configuration). An Asimov data set was used with the electron counts substituted by their expectation values instead of fluctuated according to Poissonian statistics. In the simulation, all but one parameter of the KATRIN energy loss model were taken as the given mean values in appendix D. One parameter was displaced from this value by the amount of its  $1\text{-}\sigma$ -uncertainty (one time negatively and

one time positively). The neutrino mass was inferred from the simulated data in order to quantify the systematic shift that such a displacement can induce. Two types of fits were performed:

1. The fit parameters of the KATRIN energy loss model were fixed to their means as given in appendix D. This causes a systematic shift of the squared neutrino mass because one parameter of the KATRIN energy loss model is displaced in the simulation.
2. The parameters of the KATRIN energy loss model ( $\pi_{\text{eloss}}$  and  $\pi_{\text{eloss+}}$  from eq. 6.1 and eq. 6.6) are left as free fit parameters constrained by “pull terms” as described in section 6.3.1 about the combination of a calibration and a neutrino mass measurement. This may partly compensate for a systematic shift of the squared neutrino mass.

This procedure was repeated for each of the nine parameters of the KATRIN energy loss model. In total, 36 shifts of the squared neutrino mass were determined. Two for the positive and negative displacement of a parameter times two for the two fit types listed above times nine repetitions with one for each parameter. Furthermore, the same method was then also applied for the Aseev model (the corresponding parameter means and  $1-\sigma$  uncertainties are given in the caption of figure 3.3). The induced shifts on the inferred squared neutrino mass are listed in table 6.2.

This study aimed to assess whether further free fit parameters can compensate for systematic shifts in neutrino mass inference with regard to the energy loss model. Indeed a reduction of the systematic effects could be achieved in all cases listed in table 6.2. For the KATRIN model, the reduction of systematic shifts is of the order of  $10^{-4} \text{ eV}^2$ . This effect is too small to be significant because the systematic budget for the energy loss model is  $6 \times 10^{-3} \text{ eV}^2$  according to the KATRIN Design Report [KAT05]. For the Aseev model, a significant reduction on the order  $10^{-3} \text{ eV}^2$  could be achieved for the parameters  $\omega_1$  and  $A_1$ . That the additional fit parameters have a greater effect for the Aseev model than for the KATRIN model might be caused by the different dimensionality of the models (5 additional free fit parameters for the Aseev model and 15 additional free fit parameters for the KATRIN energy loss model). A higher dimensionality leaves more degrees of freedom when maximizing the likelihood. The effects of the difference of the fit model and the (simulated) data might therefore be distributed over more parameters. This might allow for finding a parameter configuration that fits the data without reducing the shift of the squared neutrino mass. Also, these effects have to be related to the increase of the statistical uncertainty as per table 6.1. When using the parameters of the Aseev model as free fit parameters, the increase in the statistical uncertainty of  $\Delta\sigma_{\text{stat}}(m_\nu^2) = 3.46 \times 10^{-3} \text{ eV}^2$  (see table 6.1) counterbalances the reduction of the systematic effects.

For the KATRIN model, the greatest shifts are caused by the parameters  $s_3$  and  $s_2$  on a scale that is slightly greater than the allocated systematic budget of  $6 \times 10^{-3} \text{ eV}^2$ . The other shifts are below this threshold. Each shift can be understood as the worst-case scenario on its own because the correlations of the parameters make it unlikely that two or more parameters are displaced strongly at the same time. For example  $s_3$  and  $s_2$  are anti-correlated with a correlation coefficient of  $-0.61$  (see appendix D) and hence the corresponding systematic effects are likely to compensate each other. For the Aseev model, the greatest shift is caused by the parameter  $\omega_1$ , which is also above the systematic budget of  $6 \times 10^{-3} \text{ eV}^2$ . No correlations are published for the parameters of the Aseev model, which makes it difficult to decide whether it is likely that effects cancel out or intensify each other. That cancellation of effects are of importance can also be seen in figure 6.1 that compares the uncertainty band of both energy loss models and here the KATRIN model has significantly smaller uncertainties in the ionization tail. A future study should consider to include the correlations for the KATRIN model when studying the induced

**Table 6.2.: Systematic shifts of the squared neutrino mass induced by offsets of the parameters of the energy loss model.** The left column lists all parameters of the energy loss model. Each one was displaced from its mean value by its  $1\text{-}\sigma$ -uncertainty in a separate simulation. The difference of the inferred squared neutrino mass from the simulation truth of  $0\text{eV}^2$  is given in the two central columns. The sub columns show the shift for a negatively (left) or a positively (right) displaced parameter, respectively. The study was done once for fixed parameters of the energy loss model and once having them as free but constrained fit parameters as described in the main text. The last column shows by which quantity the free fit parameters can compensate for the systematic shift (the difference was calculated before rounding). All shifts became smaller as expected. The upper half of the table shows the results for the KATRIN energy loss model and the lower half for the Aseev model.

| parameter                | fixed fit parameters                      |       | constrained fit parameters                |      | difference  |     |
|--------------------------|---|-------|---|------|---|-----|
|                          | $\Delta m_\nu^2$ ( $10^{-3}\text{eV}^2$ ) |       | $\Delta m_\nu^2$ ( $10^{-3}\text{eV}^2$ ) |      | $\Delta \Delta m_\nu^2 $ ( $10^{-3}\text{eV}^2$ ) |     |
| KATRIN energy loss model |   |       |   |      |   |     |
| $A_1$                    | 3.3                                       | -3.4  | 3.2                                       | -3.3 | 0.1   | 0.1 |
| $m_1$                    | -0.2                                      | 0.1   | 0.0                                       | 0.1  | 0.2   | 0.0 |
| $s_1$                    | 4.1                                       | -4.2  | 4.1                                       | -4.1 | 0.1   | 0.1 |
| $A_2$                    | 5.4                                       | -5.4  | 5.3                                       | -5.3 | 0.1   | 0.1 |
| $m_2$                    | -0.8                                      | 0.8   | -0.8                                      | 0.8  | 0.0   | 0.0 |
| $s_2$                    | 9.3                                       | -9.3  | 9.1                                       | -9.1 | 0.2   | 0.2 |
| $A_3$                    | 3.8                                       | -3.9  | 3.5                                       | -3.7 | 0.3   | 0.1 |
| $m_3$                    | -0.5                                      | 0.5   | -0.4                                      | 0.5  | 0.1   | 0.0 |
| $s_3$                    | 10.2                                      | -10.1 | 9.9                                       | -9.9 | 0.4   | 0.2 |
| Aseev energy loss model  |   |       |   |      |   |     |
| $A_1$                    | 4.1                                       | -4.1  | 1.4                                       | -1.4 | 2.7   | 2.7 |
| $A_2$                    | 1.1                                       | -1.1  | 0.2                                       | -0.2 | 0.9   | 0.9 |
| $\omega_1$               | 8.0                                       | -8.0  | 3.0                                       | -3.0 | 5.0   | 5.0 |
| $\omega_2$               | 0.4                                       | -0.4  | 0.0                                       | 0.0  | 0.4   | 0.4 |
| $\epsilon_2$             | 0.3                                       | -0.3  | 0.0                                       | 0.0  | 0.3   | 0.3 |

systematic shifts. In a rough approximation, one can look at the average of the 18 absolute shifts in table 6.2 for the KATRIN energy loss model, which yields

$$\langle \Delta m_{\nu,\text{fix}}^2 \rangle = 4.2 \times 10^{-3} \text{eV}^2$$

for the case of fixed energy loss parameters and

$$\langle \Delta m_{\nu,\text{free}}^2 \rangle = 4.1 \times 10^{-3} \text{eV}^2$$

for the case of free and constrained parameters. Both values are within the systematic budget of  $6 \times 10^{-3} \text{eV}^2$ . Given the increased statistical uncertainty by  $\Delta\sigma_{\text{stat}}(m_\nu^2) = 0.1 \times 10^{-3} \text{eV}^2$  (eq. 6.11) in the case of free parameters, the improvement of the systematic shift is completely compensated.

It should be noted, that the presented study only investigates the uncertainties of the energy loss models, but does not make further statements about their applicability. In other words, there may be reasons to transit to the KATRIN model beyond the improvement of the sensitivity on simulated data sets. For example, the inclusion of the second peak as depicted in figure 6.1 in the energy loss function may prove to be more correct when used with real data.

## 6.5. Conclusion

A general statistical framework was developed that enables the treatment of model uncertainties by additional free fit parameters. Leaving the parameters of the KATRIN energy loss model as free in order to reduce systematic effects would increase the statistical uncertainty on the squared neutrino mass by  $\Delta\sigma_{\text{tot}}(m_\nu^2) = 1.2 \times 10^{-4} \text{ eV}^2$ , which is negligible given the KATRIN systematic budget. Additionally, the reduction of systematic effects is on the order of  $10^{-4} \text{ eV}^2$ , which is not significant given the systematic budget on the order of  $10^{-3} \text{ eV}^2$  for the energy loss function.

Aside from systematic effects, further conclusions could be drawn from the statistical studies: It was shown, that Asimov data sets of the presented kind may reasonably be assumed to be representative for an ensemble of neutrino mass measurements. Also, it was demonstrated, that precautions have to be taken when using the chi-square statistic as a measure for goodness-of-fit at the same time as incorporating “pull terms”. And it was confirmed that the profile-likelihood method yields a confidence interval of the expected coverage of 68% in the presented study.





## 7. Summary

In the scope of this thesis, two aspects concerning the inelastic electron scattering off gas molecules in the WGTS were investigated:

### **Energy-dependence of the Inelastic Scattering Cross Section**

A model for electron-scattering in the WGTS was established in dependence of the incident electron energy. The main finding was that, if the energy-dependence of the scattering cross section in neutrino mass inference were neglected, it could lead to a systematic bias of the estimated squared neutrino mass of  $\Delta m_\nu^2 = 1.09 \times 10^{-2} \text{ eV}^2$  when using the design KATRIN analysis interval of 30 eV below the endpoint of the  $\beta$  spectrum. However, the effect does not have to be neglected, but can be incorporated in the analysis using the model that was established. It should be mentioned, that certain approximations were made in the modeling process. For details and for further findings, the reader is referred to the conclusion and outlook section 5.5.

### **Statistical Methods and the KATRIN Energy Loss Model**

The main achievement is the transfer of statistical methods into the context of the KATRIN experiment to enable the treatment of a high-dimensional model of a KATRIN measurement in neutrino mass inference. The statistical tools were applied to assess whether systematic effects stemming from the energy loss function can be reduced by additional free fit parameters. The reduction was on the order of  $10^{-4} \text{ eV}^2$  or less for the inferred squared neutrino mass, which is not significant given the systematic budget of  $6 \times 10^{-3} \text{ eV}^2$  allocated for the energy loss function by the KATRIN Design Report. For further findings, the reader is referred to the conclusion section 6.5.



# Appendix

## A. Proof that the Poisson Model for Electron-Scattering is a Probability Density

In the following, it is shown that the energy-dependent Poisson model  $\bar{P}_l(E_S)$  (see equation 5.15) resembles a probability density in  $l$  independent of a constant starting energy  $E_S$  of electrons. Accordingly, the two properties that have to be verified are

$$\forall E_S > 0 : \bar{P}_l(E_S) \geq 0 \quad (\text{A.1})$$

$$\sum_l^{\infty} \bar{P}_l(E_S) = 1. \quad (\text{A.2})$$

The first condition holds because all quantities in the calculation of  $\bar{P}_l(E_S)$  are positive. For the second condition, one can use that the Poisson distribution is a probability density that sums to one

$$\begin{aligned} \sum_l^{\infty} \bar{P}_l(E_S) &= \frac{1}{d \cdot (1 - \cos(\theta_{\max}))} \int_{-d/2}^{d/2} \int_0^{\theta_{\max}} \sin(\theta_S) \sum_l^{\infty} \text{Poisson}(\mu(E_S, z_S, \theta_S), l) \, d\theta_S \, dz_S \\ &= \frac{1}{d \cdot (1 - \cos(\theta_{\max}))} \int_{-d/2}^{d/2} \int_0^{\theta_{\max}} \sin(\theta_S) \cdot 1 \, d\theta_S \, dz_S \\ &= 1. \end{aligned}$$

This verifies that the Poisson model is a probability density in the amount of scatterings  $l$ .

## B. Measurement Time Distributions

Below, the five different MTDs, one per column, as used in the study in section 5.3.2 are given. The total measurement time amounts to three years each. The retarding energies  $qU$  are denoted relative to  $E_0 = 18\,575\text{ eV}$ .

| $qU - E_0$ (eV) | measurement time (s) |          |          |          |          |
|-----------------|----------------------|----------|----------|----------|----------|
| -50.0           | 0                    | 0        | 0        | 0        | 550000   |
| -49.0           | 0                    | 0        | 0        | 0        | 550000   |
| -48.0           | 0                    | 0        | 0        | 0        | 550000   |
| -47.0           | 0                    | 0        | 0        | 0        | 550000   |
| -46.0           | 0                    | 0        | 0        | 0        | 550000   |
| -45.0           | 0                    | 0        | 0        | 0        | 550000   |
| -44.0           | 0                    | 0        | 0        | 0        | 550000   |
| -43.0           | 0                    | 0        | 0        | 0        | 550000   |
| -42.0           | 0                    | 0        | 0        | 0        | 550000   |
| -41.0           | 0                    | 0        | 0        | 0        | 550000   |
| -40.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -39.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -38.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -37.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -36.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -35.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -34.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -33.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -32.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -31.0           | 0                    | 0        | 0        | 720000   | 550000   |
| -30.0           | 0                    | 0        | 1050000  | 720000   | 550000   |
| -29.0           | 0                    | 0        | 1050000  | 720000   | 550000   |
| -28.0           | 0                    | 0        | 1050000  | 720000   | 550000   |
| -27.0           | 0                    | 0        | 1050000  | 720000   | 550000   |
| -26.0           | 0                    | 0        | 1050000  | 720000   | 550000   |
| -25.0           | 0                    | 1350000  | 1050000  | 720000   | 550000   |
| -24.0           | 0                    | 1350000  | 1050000  | 720000   | 550000   |
| -23.0           | 0                    | 1350000  | 1050000  | 720000   | 550000   |
| -22.0           | 0                    | 1350000  | 1050000  | 720000   | 550000   |
| -21.0           | 0                    | 1350000  | 1050000  | 720000   | 550000   |
| -20.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -19.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -18.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -17.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -16.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -15.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -14.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -13.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -12.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -11.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -10.0           | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -9.0            | 1910000              | 1350000  | 1050000  | 720000   | 550000   |
| -8.0            | 908164               | 908164   | 908164   | 908164   | 908164   |
| -7.5            | 908164               | 908164   | 908164   | 908164   | 908164   |
| -7.0            | 908164               | 908164   | 908164   | 908164   | 908164   |
| -6.5            | 908164               | 908164   | 908164   | 908164   | 908164   |
| -6.0            | 908164               | 908164   | 908164   | 908164   | 908164   |
| -5.5            | 908164               | 908164   | 908164   | 908164   | 908164   |
| -5.0            | 1975257              | 1975257  | 1975257  | 1975257  | 1975257  |
| -4.5            | 12214808             | 12214808 | 12214808 | 12214808 | 12214808 |
| -4.0            | 20728847             | 20728847 | 20728847 | 20728847 | 20728847 |
| -3.5            | 2020665              | 2020665  | 2020665  | 2020665  | 2020665  |
| -3.0            | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |
| -2.0            | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |
| -1.0            | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |
| 0.0             | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |
| 1.0             | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |
| 2.0             | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |
| 3.0             | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |
| 4.0             | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |
| 5.0             | 3300000              | 3300000  | 3300000  | 3300000  | 3300000  |

### C. Documentation of KaFit-Likelihood Extensions

In the scope of this thesis, the implementation of the likelihood in the KaFit framework was extended as described in section 6.3.1. The possibility was introduced to multiply the likelihood by different function types that resemble term (2) in equation (6.8).

KaFit is configured using an XML-like syntax. Example excerpts from KaFit-XML-configurations are given below. The documentation is given to a level that enables the usage by a user already familiar with KaFit. `[Double]` is used as a placeholder for a number and `[Index*]` for a parameter index. For example, the squared neutrino mass has the parameter index 0. A new `Penalty`-tag was introduced as sub-tag of the already established `LoglikelihoodKatrin`-tag:

```
<LoglikelihoodKatrin
  Name="myKatrinLogL" PDF="Gauss" RunSource="myRunGen"
  SpectrumSimulator="mySpecSim">
  <Penalty>
    <!-- Penalty Type -->
  </Penalty>
</LoglikelihoodKatrin>
```

`<!-- Penalty Type -->` can be substituted by one or more of the tags listed below.

#### Multivariate Normal Distribution

The `Parameter`-tag represents a parameter the likelihood depends on. Its attribute `Mean` specifies the mean of a parameter, `Std` the standard deviation, and one or more `Correlation`-sub-tags the correlations between the parameters.

```
<MultivarNorm>
  <Parameter Index="[Index1]" Mean="[Double]" Std="[Double]" />
  <Parameter Index="[Index2]" Mean="[Double]" Std="[Double]" />
  <!-- ... -->
  <Correlation Index1="[Index1]" Index2="[Index2]" Value="[Double]" />
  <!-- ... -->
</MultivarNorm>
```

#### One-Dimensional Gaussian Distribution

Analogously to the multivariate normal distribution, a one-dimensional Gaussian distribution can be used:

```
<Gaussian ParamIndex="[Index]" Mean="[Double]" Std="[Double]" />
```

#### Uniform Neutrino Mass Prior

A uniform prior on the neutrino mass in a Bayesian analysis can be set via the following tag:

```
<ConstInSqrt ParamIndex="0" />
```

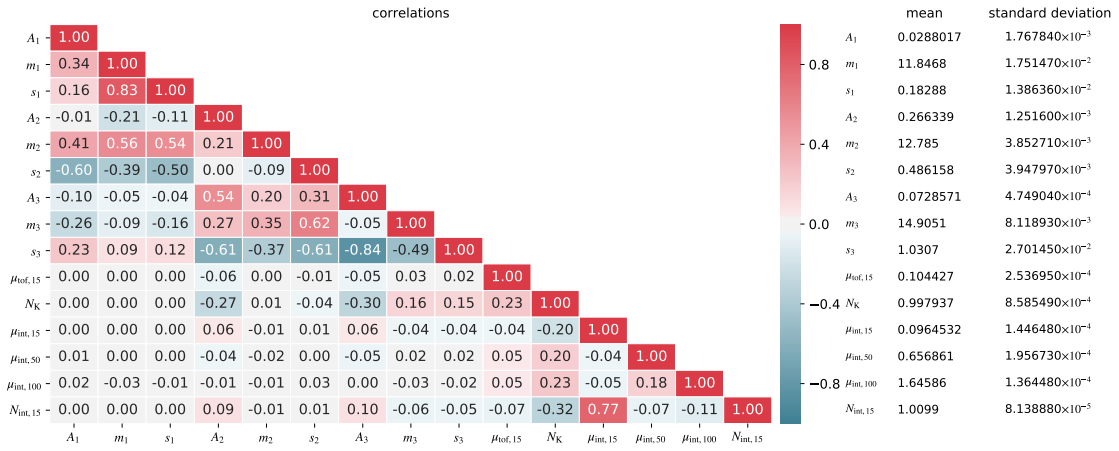
The former implementation in KaFit only allowed for a constant prior on the squared neutrino mass. The following lines derive the form of a prior on  $m_\nu^2$  that resembles a uniform prior on  $m_\nu$ . Let  $f(m_\nu) = C = \text{constant}$  be the prior on  $m_\nu$  and  $g(m_\nu^2)$  be the

prior on  $m_\nu^2$ . Starting from conservation of probability one derives

$$\begin{aligned} f(m_\nu) dm_\nu &= g(m_\nu^2) dm_\nu^2 \\ \Rightarrow g(m_\nu^2) &= f(m_\nu) \left( \frac{dm_\nu^2}{dm_\nu} \right)^{-1} \\ \Rightarrow g(m_\nu^2) &= C \frac{1}{2\sqrt{m_\nu^2}}. \end{aligned}$$

#### D. Preliminary Parameter Values of the KATRIN Energy Loss Model

The best-fit values, standard deviations and correlations of the fit parameters for the KATRIN energy loss model for electrons scattering off deuterium molecules as used in this thesis are listed below [Pre-Han19-1]. The parameter names follow section 6.1. The parameters that actually enter into the KATRIN energy loss model are the first nine from  $A_1$  to  $s_3$ . Their uncertainties are derived using the HESSE algorithm [S-MIN] of the MINUIT2 software package.



## E. Configuration of the Sensitivity Study using the Empirical KATRIN Energy Loss Model

The configuration of the SSC and KaFit modules as used in the sensitivity study in chapter 6 is listed below.

| <b>WGTS</b>                    |  |
|--------------------------------|--|
| gas column density             | $5 \times 10^{21}$ molecules/m <sup>2</sup> (constant, no density profile) |
| slices                         | 1  |
| length                         | 10.082 m   |
| tritium purity                 | 95 %   |
| <b>Differential Spectrum</b>   |  |
| final molecular states         | by Saenz, emulating Doppler effect   |
| theo. corrections              | screening, radiation (reference energy: molecular final states)            |
| Fermi function                 | approximately relativistic   |
| endpoint energy                | 18 575 eV  |
| squared neutrino mass          | 0 eV <sup>2</sup>  |
| <b>Energy Loss</b>             |  |
| energy loss function           | KATRIN or Aseev model  |
| inel. scattering cross section | $3.456 \times 10^{-22}$ m <sup>2</sup>                                     |
| elas. scattering               | neglected  |
| <b>Transmission Function</b>   |  |
| general configuration          | relativistic, not detailed   |
| mag. field in analyzing plane  | $3 \times 10^{-4}$ T   |
| mag. field of pinch magnet     | 6 T  |
| mag. field in WGTS             | 3.6 T  |
| <b>Detector</b>                |  |
| efficiency                     | 95 %   |
| <b>MTD</b>                     |  |
| voltages                       | reference MTD from the KATRIN Design Report                                |
| range                          | $[E_0 - 30 \text{ eV}, E_0 + 5 \text{ eV}]$                                |
| duration                       | 3 years  |
| <b>Background</b>              |  |
| rate                           | 10 mcps  |





# Acronyms

**BIXS** beta-induced X-ray spectroscopy.

**CKrS** condensed  $^{83m}\text{Kr}$  source.

**CPS** cryogenic pumping section.

**DPS** differential pumping section.

**EMCS** Earth magnetic field compensation system.

**FBM** forward beam monitor.

**FPD** focal plane detector.

**FT** KATRIN First Tritium measurement campaign.

**KaFit** KATRIN Fitter.

**KASPER** KATRIN Analysis and Simulations Package.

**KNM1** KATRIN neutrino mass measurement campaign 1.

**LARA** laser Raman system.

**LEP** Large Electron Positron Collider.

**LFCS** low-field correction system.

**MAC-E** magnetic adiabatic collimation with electrostatic filtering.

**MLE** maximum likelihood estimator.

**MTD** measurement time distribution.

**PULCINELLA** precision ultra-low current integrating normalization electrometer for low-level analysis.

**RMMS** radial magnetic measuring system.

**RS** rear section.

**SDS** spectrometer and detector section.

**SM** Standard Model of Particle Physics.

**SSC** source and spectrum calculation.

**SSM** standard solar model.

**STS** source and transport section.

**TDR** KATRIN Technical Design Report.

**TLK** Tritium Laboratory Karlsruhe.

**VMMS** vertical magnetic measuring system.

**WGTS** windowless gaseous tritium source.

# Bibliography

## References

- [Abd+09] J. N. Abdurashitov et al. “Measurement of the solar neutrino capture rate with gallium metal. III. Results for the 2002–2007 data-taking period”. *Phys. Rev. C* **80**.1 (2009), 015807. DOI: 10.1103/PhysRevC.80.015807.
- [Abd+17] D. N. Abdurashitov et al. “Electron scattering on hydrogen and deuterium molecules at 14–25 keV by the ‘Troitsk nu-mass’ experiment”. *Phys. Part. Nuclei Lett.* **14**.6 (2017), 892–899. DOI: 10.1134/S1547477117060024.
- [Abe+16] Y. Abe et al. “Measurement of  $\theta_{13}$  in Double Chooz using neutron captures on hydrogen with novel background rejection techniques”. *J. High Energ. Phys.* **2016** (2016), 163. DOI: 10.1007/JHEP01(2016)163.
- [Acc+98] M. Acciarri et al. “Determination of the number of light neutrino species from single photon production at LEP”. *Phys. Lett.* **431**.1 (1998), 199–208. DOI: 10.1016/0370-2693(89)90704-1.
- [Agh+18] N. Aghanim et al. “Planck 2018 results. VI. Cosmological parameters”. (2018). arXiv: 1807.06209 [astro-ph.CO].
- [Ago+18] M. Agostini et al. “Improved limit on neutrinoless double- $\beta$  decay of  $^{76}\text{Ge}$  from GERDA phase II”. *Phys. Rev. Lett.* **120**.13 (2018), 132503. DOI: 10.1103/PhysRevLett.120.132503.
- [Agu+09] A. A. Aguilar-Arevalo et al. “The MiniBooNE detector”. *Nucl. Instrum. Meth. A* **599**.1 (2009), 28–46. DOI: doi.org/10.1016/j.nima.2008.10.028.
- [Aha+13] B. Aharmim et al. “Combined analysis of all three phases of solar neutrino data from the Sudbury Neutrino Observatory”. *Phys. Rev. C* **88**.2 (2013), 025501. DOI: 10.1103/PhysRevC.88.025501.
- [Alt+05] M. Altmann et al. “Complete results for five years of GNO solar neutrino observations”. *Phys. Lett. B* **616**.3 (2005), 174–190. DOI: doi.org/10.1016/j.physletb.2005.04.068.
- [Ams+15] J. Amsbaugh et al. “Focal-plane detector system for the KATRIN experiment”. *Nucl. Instrum. Methods Phys. Res.* **778** (2015), 40–60. ISSN: 0168-9002. DOI: 10.1016/j.nima.2014.12.116.
- [Ant15] J. Antoni. “Monte-Carlo simulations of systematic effects due to inelastic scattering in the KATRIN tritium source”. Diploma thesis. KIT (2015).
- [Are+16] M. Arenz et al. “Commissioning of the vacuum system of the KATRIN main spectrometer”. *J. Instrum.* **11**.04 (2016), P04011. DOI: 10.1088/1748-0221/11/04/p04011.
- [Are+18a] M. Arenz et al. “First transmission of electrons and ions through the KATRIN beamline”. *J. Instrum.* **13**.04 (2018), P04020–P04020. DOI: 10.1088/1748-0221/13/04/p04020.

- [Are+18b] M. Arenz et al. “The KATRIN superconducting magnets: overview and first performance results”. *J. Instrum.* **13**.08 (2018), T08005. DOI: 10.1088/1748-0221/13/08/t08005.
- [Ase+00] V. Aseev et al. “Energy loss of 18 keV electrons in gaseous T and quench condensed D films”. *Eur. Phys. J. D* **10**.1 (2000), 39–52. DOI: 10.1007/s100530050525.
- [Ase+11] V. N. Aseev et al. “An upper limit on electron antineutrino mass from Troitsk experiment”. *Phys. Rev. D* **84** (2011), 112003. DOI: 10.1103/PhysRevD.84.112003.
- [Ass+96] K. Assamagan et al. “Upper limit of the muon-neutrino mass and charged-pion mass from momentum analysis of a surface muon beam”. *Phys. Rev. D* **53**.11 (1996), 6065–6077. DOI: 10.1103/PhysRevD.53.6065.
- [Bab+12] M. Babutzka et al. “Monitoring of the operating parameters of the KATRIN windowless gaseous tritium source”. *New J. Phys.* **14**.10 (2012), 103046. DOI: 10.1088/1367-2630/14/10/103046.
- [Bab14] M. Babutzka. “Design and development for the rearsection of the KATRIN experiment”. PhD thesis. KIT (2014). DOI: 10.5445/IR/1000045598.
- [Bah96] J. N. Bahcall. “Solar neutrinos: Where we are, where we are going”. *Astrophys. J* **467** (1996), 475. DOI: 10.1086/177624.
- [Bar+98] R. Barate et al. “An upper limit on the tau-neutrino mass from three-prong and five-prong tau decays”. *Eur. Phys. J. C* **2** (1998), 395–406. DOI: 10.1007/s100520050149.
- [Bau14] S. Bauer. “Energy calibration and stability monitoring of the KATRIN experiment”. PhD thesis. WWU (2014). URL: <https://miami.uni-muenster.de/Record/da3d6759-9202-4b68-bb80-711f52b208d8>.
- [BC38] H. A. Bethe and C. L. Critchfield. “The formation of deuterons by proton combination”. *Phys. Rev.* **54**.4 (1938), 248–254. DOI: 10.1103/PhysRev.54.248.
- [Bec00] H. Becquerel. “Contribution à l’étude du rayonnement du radium”. *J. Phys. Theor. Appl.* **9**.1 (1900), 190–199.
- [Bec95] H. Becquerel. “Sur les Radiations Invisibles Émises par les Corps Phosphorescents”. *C. R. Acad. Sci.* **122** (1895), 501–503.
- [Bet39] H. A. Bethe. “Energy production in stars”. *Phys. Rev.* **55**.5 (1939), 434–456. DOI: 10.1103/PhysRev.55.434.
- [Boh32] N. Bohr. “Faraday lecture. Chemistry and the quantum theory of atomic constitution”. *J. Chem. Soc.* **0** (1932), 349–384. DOI: 10.1039/JR9320000349.
- [Bon+99] J. Bonn et al. “A high resolution electrostatic time-of-flight spectrometer with adiabatic magnetic collimation”. *Nucl. Instrum. Methods Phys. Res.* **421**.1 (1999), 256–265. DOI: 10.1016/S0168-9002(98)01263-7.
- [BPB01] J. N. Bahcall, M. H. Pinsonneault, and S. Basu. “Solar models: current epoch and time dependences, neutrinos, and helioseismological properties”. *Astrophys. J.* **555**.2 (2001), 990–1012. DOI: 10.1086/321493.
- [BPR15] L. I. Bodine, D. S. Parno, and R. G. H. Robertson. “Assessment of molecular effects on neutrino mass measurements from tritium  $\beta$  decay”. *Phys. Rev. C* **91**.3 (2015), 035505. DOI: 10.1103/PhysRevC.91.035505.

- [BPT80] G. Beamson, H. Q. Porter, and D. W. Turner. “The collimating and magnifying properties of a superconducting field photoelectron spectrometer”. *J. Phys. E* **13**.1 (1980), 64–66. DOI: 10.1088/0022-3735/13/1/018.
- [Cha14] J. Chadwick. “Intensitätsverteilung im magnetischen Spectrum der  $\beta$ -Strahlen von Radium B + C”. *Verhandl. Dtsch. Phys. Ges.* **16** (1914), 383.
- [Cle+98] B. T. Cleveland et al. “Measurement of the solar electron neutrino flux with the Homestake chlorine detector”. *Astrophys. J.* **496** (1998), 505–526. DOI: 10.1086/305343.
- [Cow+11] G. Cowan et al. “Asymptotic formulae for likelihood-based tests of new physics”. *Eur. Phys. J. C* **71** (2011), 1554. DOI: 10.1140/epjc/s10052-011-1554-0.
- [Cow+56] C. L. Cowan et al. “Detection of the free neutrino: a confirmation”. *Science* **124**.3212 (1956), 103–104. DOI: 10.1126/science.124.3212.103.
- [Dan+62] G. Danby et al. “Observation of high-energy neutrino reactions and the existence of two kinds of neutrinos”. *Phys. Rev. Lett.* **9** (1962), 36–44. DOI: 10.1103/PhysRevLett.9.36.
- [Dor+04] A. Doroshkevich et al. “Large scale structure in the SDSS galaxy survey”. *Astron. Astrophys.* **418** (2004), 7–23.
- [Dos+06] N. Doss et al. “Molecular effects in investigations of tritium molecule beta decay endpoint experiments”. *Phys. Rev. C* **73** (2006), 025502. DOI: 10.1103/PhysRevC.73.025502.
- [Dyb19] S. Dyba. “Background reduction by the inner wire electrode and set-up of the condensed krypton source at the neutrino mass experiment KATRIN”. PhD thesis. WWU (2019).
- [Ell+17] E. Ellinger et al. “Monitoring the KATRIN source properties within the beamline”. *J. Phys. Conf. Ser.* **888** (2017), 012229. DOI: 10.1088/1742-6596/888/1/012229.
- [Ell19] E. Ellinger. “Development and investigation of the Forward Beam Monitor for the KATRIN experiment”. To be published. PhD thesis. Bergische Universität Wuppertal (2019).
- [Erh+14] M. Erhard et al. “High-voltage monitoring with a solenoid retarding spectrometer at the KATRIN experiment”. *J. Instrum.* **9**.06 (2014), P06022–P06022. DOI: 10.1088/1748-0221/9/06/p06022.
- [Erh+18] M. Erhard et al. “Technical design and commissioning of the KATRIN large-volume air coil system”. *J. Instrum.* **13**.02 (2018), P02003. DOI: 10.1088/1748-0221/13/02/p02003.
- [Est+19] I. Esteban et al. “Global analysis of three-flavour neutrino oscillations: synergies and tensions in the determination of  $\theta_{23}$ ,  $\delta_{cp}$ , and the mass ordering”. *J. High Energy Phys.* **2019**.1 (2019), 106. DOI: 10.1007/JHEP01(2019)106.
- [EW27] C. D. Ellis and W. A. Wooster. “The average energy of disintegration of radium E”. *Proc. Royal Soc. Lond. A* **117** (1927), 109. DOI: 10.1098/rspa.1927.0168.
- [FC98] G. J. Feldman and R. D. Cousins. “Unified approach to the classical statistical analysis of small signals”. *Phys. Rev. D* **57**.7 (1998), 3873–3889. DOI: 10.1103/PhysRevD.57.3873.

- [Fer34] E. Fermi. “Versuch einer Theorie der  $\beta$ -Strahlen”. *Z. Phys.* **88.3** (1934), 161–177. DOI: 10.1007/BF01351864.
- [Frä10] F. M. Fränkle. “Background investigations of the KATRIN pre-spectrometer”. PhD thesis. KIT (2010). DOI: 10.5445/IR/1000019392.
- [Fuk+98] Y. Fukuda et al. “Evidence for oscillation of atmospheric neutrinos”. *Phys. Rev. Lett.* **81** (1998), 1562–1567. DOI: 10.1103/PhysRevLett.81.1562.
- [Gan+16] A. Gando et al. “Search for Majorana neutrinos near the inverted mass hierarchy region with KamLAND-Zen”. *Phys. Rev. Lett.* **117.8** (2016), 082503. DOI: 10.1103/PhysRevLett.117.082503.
- [Gei64] J. Geiger. “Streuung von 25 keV-Elektronen an Gasen”. *Z. Phys.* **181.4** (1964), 413–425. DOI: 10.1007/BF01380873.
- [Ger75] D. E. Gerhart. “Comprehensive optical and collision data for radiation action. I. H<sub>2</sub>”. *J. Chem. Phys.* **62.3** (1975), 821–832. DOI: 10.1063/1.430533.
- [Gla61] S. L. Glashow. “Partial symmetries of weak interactions”. *Nucl. Phys.* **22** (1961), 579–588. DOI: 10.1016/0029-5582(61)90469-2.
- [Gol+58] M. Goldhaber et al. “Helicity of neutrinos”. *Phys. Rev.* **109.3** (1958), 1015–1017. DOI: 10.1103/PhysRev.109.1015.
- [Gro+08] S. Grohmann et al. “Cryogenic design of the KATRIN source cryostat”. *AIP Conf. Proc.* **985.1** (2008), 1277–1284. DOI: 10.1063/1.2908483.
- [Gro15] S. Groh. “Modeling of the response function and measurement of transmission properties of the KATRIN experiment”. PhD thesis. KIT (2015). DOI: 10.5445/IR/1000046546.
- [Har15] F. T. Harms. “Characterization and minimization of background processes in the KATRIN main spectrometer”. PhD thesis. KIT (2015). DOI: 10.5445/IR/1000050027.
- [Hei18] F. Heizmann. “Analysis tools and methods for tritium data taking with the KATRIN experiment”. To be published. PhD thesis. KIT (2018).
- [Hig64] P. W. Higgs. “Broken symmetries and the masses of gauge bosons”. *Phys. Rev. Lett.* **13** (1964), 508–509. DOI: 10.1103/PhysRevLett.13.508.
- [Höt12] M. Hötzel. “Simulation and analysis of source-related effects for KATRIN”. PhD thesis. KIT (2012). DOI: 10.5445/IR/1000031259.
- [Ino71] M. Inokuti. “Inelastic collisions of fast charged particles with atoms and molecules—the Bethe theory revisited”. *Rev. Mod. Phys.* **43.3** (1971), 297–347. DOI: 10.1103/RevModPhys.43.297.
- [Jac75] J. Jackson. “Classical electrodynamics”. Wiley, Hoboken, 1975. ISBN: 9780471431329.
- [Jan15] A. Jansen. “The cryogenic pumping section of the KATRIN experiment - design studies and experiments for the commissioning”. PhD thesis. KIT (2015). DOI: 10.5445/IR/1000047146.
- [Käf12] W. Käfer. “Sensitivity studies for the KATRIN experiment”. PhD thesis. KIT (2012). DOI: 10.5445/IR/1000026021.
- [KAT05] J. Angrik et al. “KATRIN design report 2004”. *Wissenschaftliche Berichte / Forschungszentrum Karlsruhe* **7090** (2005).
- [Kir98] T. Kirsten. “GALLEX solar neutrino results”. *Prog. Part. Nucl. Phys.* **40** (1998), 85–99. DOI: doi.org/10.1016/S0146-6410(98)00013-1.

- [Kle+19] M. Kleesiek et al. “ $\beta$ -decay spectrum, response function and statistical model for neutrino mass measurements with the KATRIN experiment”. *Eur. Phys. J. C* **79.3** (2019), 204. DOI: 10.1140/epjc/s10052-019-6686-7.
- [Kle14] M. Kleesiek. “A data-analysis and sensitivity-optimization framework for the KATRIN experiment”. PhD thesis. KIT (2014). DOI: 10.5445/IR/1000043301.
- [Klein19] M. Klein. “Tritium ions in KATRIN: blocking, removal and detection”. PhD thesis. KIT (2019). DOI: 10.5445/IR/1000093526.
- [Kod+01] K. Kodama et al. “Observation of tau neutrino interactions”. *Phys. Lett. B* **504** (2001), 218–224. DOI: 10.1016/S0370-2693(01)00307-0.
- [Kos12] A. Kosmider. “Tritium retention techniques in the KATRIN transport section and commissioning of its DPS2-F cryostat”. PhD thesis. KIT (2012). DOI: 10.5445/IR/1000028959.
- [KR94] Y.-K. Kim and M. E. Rudd. “Binary-encounter-dipole model for electron-impact ionization”. *Phys. Rev. A* **50.5** (1994), 3954–3967. DOI: 10.1103/PhysRevA.50.3954.
- [Kra+05] C. Kraus et al. “Final results from phase II of the Mainz neutrino mass search in tritium decay”. *Eur. Phys. J. C* **40** (2005), 447–468.
- [Kuc+18] L. Kuckert et al. “Modelling of gas dynamical properties of the KATRIN tritium source and implications for the neutrino mass measurement”. *Vacuum* **158** (2018), 195–205. DOI: 10.1016/j.vacuum.2018.09.036.
- [Kuc16] L. Kuckert. “The windowless gaseous tritium source of the KATRIN experiment - characterisation of gas dynamical and plasma properties”. PhD thesis. KIT (2016). DOI: 10.5445/IR/1000065077.
- [Let+18] J. Letnev et al. “Technical design and commissioning of a sensor net for fine-meshed measuring of the magnetic field at the KATRIN spectrometer”. *J. Instrum.* **13.08** (2018), T08010. DOI: 10.1088/1748-0221/13/08/t08010.
- [Liu73] J. W. Liu. “Total inelastic cross section for collisions of  $H_2$  with fast charged particles”. *Phys. Rev. A* **7.1** (1973), 103–109. DOI: 10.1103/PhysRevA.7.103.
- [Liu87] J. W. Liu. “Total cross sections for high-energy electron scattering by  $H_2$  ( $^1\Sigma_g^+$ ),  $N_2$  ( $^1\Sigma_g^+$ ), and  $O_2$  ( $^3\Sigma_g^-$ )”. *Phys. Rev. A* **35.2** (1987), 591–597. DOI: 10.1103/PhysRevA.35.591.
- [LL02] T. J. Loredo and D. Q. Lamb. “Bayesian analysis of neutrinos observed from supernova SN 1987A”. *Phys. Rev. D* **65.6** (2002), 063002. DOI: 10.1103/PhysRevD.65.063002.
- [LY56] T. D. Lee and C. N. Yang. “Question of parity conservation in weak interactions”. *Phys. Rev.* **104.1** (1956), 254–258. DOI: 10.1103/PhysRev.104.254.
- [Mer+19] S. Mertens et al. *J. Phys. G* **46.6** (2019), 065203. DOI: 10.1088/1361-6471/ab12fe.
- [MNS62] Z. Maki, M. Nakagawa, and S. Shoichi. “Remarks on the unified model of elementary particles”. *Prog. Theor. Phys.* **28.5** (1962), 870–880. DOI: 10.1143/PTP.28.870.

- [MS86] S. P. Mikheev and A. Y. Smirnov. “Resonant amplification of neutrino oscillations in matter and solar neutrino spectroscopy”. *Nuovo Cim. C* **9** (1986), 17–26. DOI: 10.1007/BF02508049.
- [Mye+15] E. G. Myers et al. “Atomic masses of tritium and helium-3”. *Phys. Rev. Lett.* **114.1** (2015), 013003. DOI: 10.1103/PhysRevLett.114.013003.
- [Ney37] J. Neyman. “Outline of a theory of statistical estimation based on the classical theory of probability”. *Philos. Trans. Royal Soc. A* **236.767** (1937), 333–380.
- [Noe18] E. Noether. “Invarianten beliebiger Differentialausdrücke” (1918). In: “Nachrichten von der Gesellschaft der Wissenschaften zu Göttingen, Mathematisch-Physikalische Klasse”, 37–44.
- [OW08] E. W. Otten and C. Weinheimer. “Neutrino mass limit from tritium beta decay”. *Rep. Prog. Phys.* **71** (2008), 086201. DOI: 10.1088/0034-4885/71/8/086201.
- [Pau30] W. Pauli. Published in: “Wissenschaftlicher Briefwechsel mit Bohr, Einstein, Heisenberg u.a. Band II: 1930–1939” by V.F. Weisskopf et al. (1985). Springer Berlin Heidelberg. ISBN: 9783540788010.
- [Pon58] B. Pontecorvo. “Inverse beta processes and nonconservation of lepton charge”. *Sov. Phys. JETP* **7** (1958), 172–173.
- [PSB15] F. Priester, M. Sturm, and B. Bornschein. “Commissioning and detailed results of KATRIN inner loop tritium processing system at Tritium Laboratory Karlsruhe”. *Vacuum* **116** (2015), 42–47. DOI: 10.1016/j.vacuum.2015.02.030.
- [Röll15] M. Röllig. “Tritium analytics by beta induced X-ray spectrometry”. PhD thesis. KIT (2015). DOI: 10.5445/IR/1000054050.
- [Röt19] C. Röttele. “Tritium suppression factor of the KATRIN transport section”. To be published. PhD thesis. KIT (2019).
- [Rut99] E. Rutherford. “VIII. Uranium radiation and the electrical conduction produced by it”. *The London, Edinburgh, and Dublin Philosophical Magazine and Journal of Science* **47.284** (1899), 109–163. DOI: 10.1080/14786449908621245.
- [Sal68] A. Salam. “Weak and electromagnetic interactions”. *Proceedings of the Eighth Nobel Symposium* (1968), 367–377.
- [Sch13] M. Schlösser. “Accurate calibration of the Raman system for the Karlsruhe Tritium Neutrino Experiment”. PhD thesis. KIT (2013). DOI: 10.5445/IR/1000034967.
- [Sch19] W. Schmitz. “Particles, Fields and Forces : A Conceptual Guide to Quantum Field Theory and the Standard Model”. 1st ed. Springer International Publishing, 2019. ISBN: 9783030128784.
- [Sei19] H. Seitz-Moskaliuk. “Characterisation of the KATRIN tritium source and evaluation of systematic effects”. PhD thesis. KIT (2019). DOI: 10.5445/IR/1000090748.
- [Shi+93] D. Shiner et al. “H<sub>2</sub>, D<sub>2</sub>, and HD ionization potentials by accurate calibration of several iodine lines”. *Phys. Rev. A* **47.5** (1993), 4042–4045. DOI: 10.1103/PhysRevA.47.4042.
- [Sim81] J. J. Simpson. “Measurement of the  $\beta$ -energy spectrum of <sup>3</sup>H to determine the antineutrino mass”. *Phys. Rev. D* **23.3** (1981), 649–662. DOI: 10.1103/PhysRevD.23.649.



- [Tan+18] M. Tanabashi et al. “Review of particle physics”. *Phys. Rev. D* **98.3** (2018), 030001. DOI: 10.1103/PhysRevD.98.030001.
- [Tho97] J. J. Thomson. “XL. Cathode rays”. *The London, Edinburgh, and Dublin Philosophical Magazine and Journal of Science* **44.269** (1897), 293–316. DOI: 10.1080/14786449708621070.
- [Thü+09] T. Thümmler et al. “Precision high voltage divider for the KATRIN experiment”. *New J. Phys.* **11.10** (2009), 103007. DOI: 10.1088/1367-2630/11/10/103007.
- [Ubi+09] M. Ubieto-Díaz et al. “A broad-band FT-ICR penning trap system for KATRIN”. *Int. J. Mass Spectrom.* **288.1** (2009), 1–5. DOI: 10.1016/j.ijms.2009.07.003.
- [Val04] K. Valerius. “Elektromagnetisches Design für das Hauptspektrometer des KATRIN Experiments”. Diploma thesis. WWU (2004). URL: [https://www.uni-muenster.de/Physik.KP/AGWeinheimer/Files/theses/Diplom\\_Kathrin\\_Valerius.pdf](https://www.uni-muenster.de/Physik.KP/AGWeinheimer/Files/theses/Diplom_Kathrin_Valerius.pdf).
- [Val09] K. Valerius. “Spectrometer-related background processes and their suppression in the KATRIN experiment”. PhD thesis. WWU (2009). URL: <http://repositorium.uni-muenster.de/document/miami/93137705-73f4-404a-a438-09d487cbff63/>.
- [Vén+18] D. Vénos et al. “Properties of 83mKr conversion electrons and their use in the KATRIN experiment”. *J. Instrum.* **13.02** (2018), T02012. DOI: 10.1088/1748-0221/13/02/t02012.
- [Wal44] A. Wald. “On cumulative sums of random variables”. *Ann. Math. Stat.* **15** (1944), 283–296.
- [Wei+93] C. Weinheimer et al. “Improved limit on the electron-antineutrino rest mass from tritium  $\beta$ -decay”. *Phys. Lett. B* **300.3** (1993), 210–216. DOI: 10.1016/0370-2693(93)90355-L.
- [Wei38] C. F. Weizsäcker. “Über Elementumwandlungen im Innern der Sterne”. *Phys. Zeit.* **39.633** (1938), 46.
- [Wei67] S. Weinberg. “A model of leptons”. *Phys. Rev. Lett.* **19.21** (1967), 1264–1266. DOI: 10.1103/PhysRevLett.19.1264.
- [Wil38] S. S. Wilks. “The large-sample distribution of the likelihood ratio for testing composite hypotheses”. *Ann. Math. Statist.* **9.1** (1938), 60–62. DOI: 10.1214/aoms/1177732360.
- [Wol78] L. Wolfenstein. “Neutrino oscillations in matter”. *Phys. Rev. D* **17** (1978), 2369–2374. DOI: 10.1103/PhysRevD.17.2369.
- [Wu+57] C. S. Wu et al. “Experimental test of parity conservation in beta decay”. *Phys. Rev.* **105.4** (1957), 1413–1415. DOI: 10.1103/PhysRev.105.1413.
- [Yèc+17] C. Yèche et al. “Constraints on neutrino masses from Lyman-alpha forest power spectrum with BOSS and XQ-100”. *J. Cosmol. Astropart. Phys.* **1706.06** (2017), 047. DOI: 10.1088/1475-7516/2017/06/047.
- [Zbo11] M. Zbořil. “Feste Elektronenquellen für die Energieskalaüberwachung im KATRIN Experiment”. PhD thesis. WWU (2011). URL: [http://repositorium.uni-muenster.de/document/miami/79b74253-2dcd-493b-a08d-4eeb5a10b8c1/diss\\_zboril.pdf](http://repositorium.uni-muenster.de/document/miami/79b74253-2dcd-493b-a08d-4eeb5a10b8c1/diss_zboril.pdf).
- [Zee03] A. Zee. “Quantum Field Theory in a Nutshell”. Princeton University Press, Princeton, 2003. ISBN: 9780691010199.

- [ZJH12] S. Zhang, G. Ji, and J. Han. “The positional and angular distribution of molecules flowing through cylindrical tube in free molecular flow”. *Phys. Procedia* **32** (2012), 513–524. DOI: 10.1016/j.phpro.2012.03.595.
- [Zub11] K. Zuber. “Neutrino Physics”. 2nd ed. CRC Press, Boca Raton, 2011. ISBN: 9781420064711.

## Preliminary

- [Pre-Han19-1] V. Hannen et al. “Global fit of parametrized energy loss function to measured integral and t.o.f. data”. GIT-Repository, 2019. URL: [https://nuserv.uni-muenster.de:8443/afulst/KATRIN-eloss/blob/master/code/STS3a/global\\_eloss\\_fit-v10.ipynb](https://nuserv.uni-muenster.de:8443/afulst/KATRIN-eloss/blob/master/code/STS3a/global_eloss_fit-v10.ipynb).
- [Pre-Han19-2] V. Hannen. “New energy loss parametrization”. 36th KATRIN Collaboration Meeting, 2019. URL: <https://fuzzy.fzk.de/bscw/bscw.cgi/1254307>.
- [Pre-Rod19] C. Rodenbeck and R. Sack. “Measurement of KATRIN’s energy loss function using a time of flight method”. DPG-Frühjahrstagung, 2019. URL: <https://fuzzy.fzk.de/bscw/bscw.cgi/1256746>.

## Software

- [Ant+09] I. Antcheva et al. “ROOT — A C++ framework for petabyte data storage, statistical analysis and visualization”. *Comput. Phys. Commun.* **180.12** (2009), 2499–2512. DOI: doi.org/10.1016/j.cpc.2009.08.005.
- [S-KAS] KATRIN collaboration. “KATRIN Analysis and Simulations Package (KASPER)”. 2019. URL: <https://userv.uni-muenster.de:8443/katrin-git/kasper>.
- [S-KF] M. Kleesiek et al. “KaFit”. Based on [Kle19+; Kle14] and others, URL: <https://userv.uni-muenster.de:8443/katrin-git/kasper>.
- [S-MIN] F. James. “MINUIT: Function Minimization and Error Analysis Reference Manual” (1998). CERN Program Library Long Writeups. URL: <https://cds.cern.ch/record/2296388>.
- [S-SSC] KATRIN collaboration. “Source and spectrum calculation (SSC)”. Based on [Kle19+; Kle14; Höt12; Gro15; Käf12, Hei18, Kuc+18] and others, URL: <https://userv.uni-muenster.de:8443/katrin-git/kasper>.

# Acknowledgments

I am convinced that there is always something new to learn. Nonetheless, handing in this work is the final act within my official education. It has been a long journey.

I would like to express my gratitude to the people who helped me in completing this thesis:

First, I would like to thank Prof. Dr. Guido Drexlin for being a reviewer of this thesis and giving me the opportunity to work at the KATRIN experiment.

Special thanks goes to Dr. Kathrin Valerius as the leader of the Young Investigator Group that I enjoyed being part of over the course of my thesis. Thank you for the warm welcome, that you enabled my stay the MIT, and that you always made time, when I needed to discuss something. And, of course, thank you, for being a reviewer of this thesis.

I would also like to thank my adviser Moritz Machatschek. Thank you for your enthusiasm when tackling my physics problems as if they were your own. Also, thank you for frequently cross-checking my ideas, hinting me towards errors and offering solutions. Especially, thank you for your easygoing and calm attitude and the thorough proofreading. It was a great experience having you as an adviser.

I would also like to thank Dr. Valérian Sibille. Thank you for advising me before, during and after my stay at the MIT. And thank you for enabling my stay in the first place. I also greatly appreciate the enormous amount of time you took for reviewing my program code and teaching me software engineering as well as physics - and (maybe unintentionally) food culture. I also found it very considerate of you to check back on me on a regular basis.

I would also like to thank Dr. Wonqook Choi. Thank you for being such a cheerful and kind person. And thank you for frequently giving me the opportunity to express my opinion in software matters and appreciating my ideas. Also, thank you for helping me with analysis problems, with the analysis software and for the many discussions.

I would also like to thank Dr. Hendrik Seitz-Moskaliuk. Thank you for having been a pleasant office mate and for always promptly addressing my questions. Also, thank you for being the first to proofread pages of my thesis. Your feedback doubtlessly influenced my further writing style.

I would also like to thank Dr. Jan Behrens. Thank you for proofreading a chapter of my thesis and for helping me many times with the computer infrastructure.

Also, I would like to thank the rest of the Young Investigator Group and the people at KATRIN that helped me on a daily basis - be it practically or motivational: Dr. Carsten Röttele, Emma Malcherek, Fabian Block, Dr. Ferenc Glück, Dr. Florian Fränkle, Dr. Florian Heizmann, Dr. Joachim Wolf, Klaus Mehret, Leonard Köllenberger, Lutz Schimpf, Marion Behechti, Dr. Markus Steidl, Rudolf Sack, Dr. Stephanie Hickford, Thomas Csabo, Dr. Thomas Thümmeler, Dr. Volker Hannen and the ones I forgot.

In a broader scope, I would like to thank my friends and family for always supporting me during my education. Especially, I want to mention my parents and my sisters. Thank you.